

The COHERENT Collaboration and the First Observation of Coherent Elastic Neutrino-Nucleus Scattering

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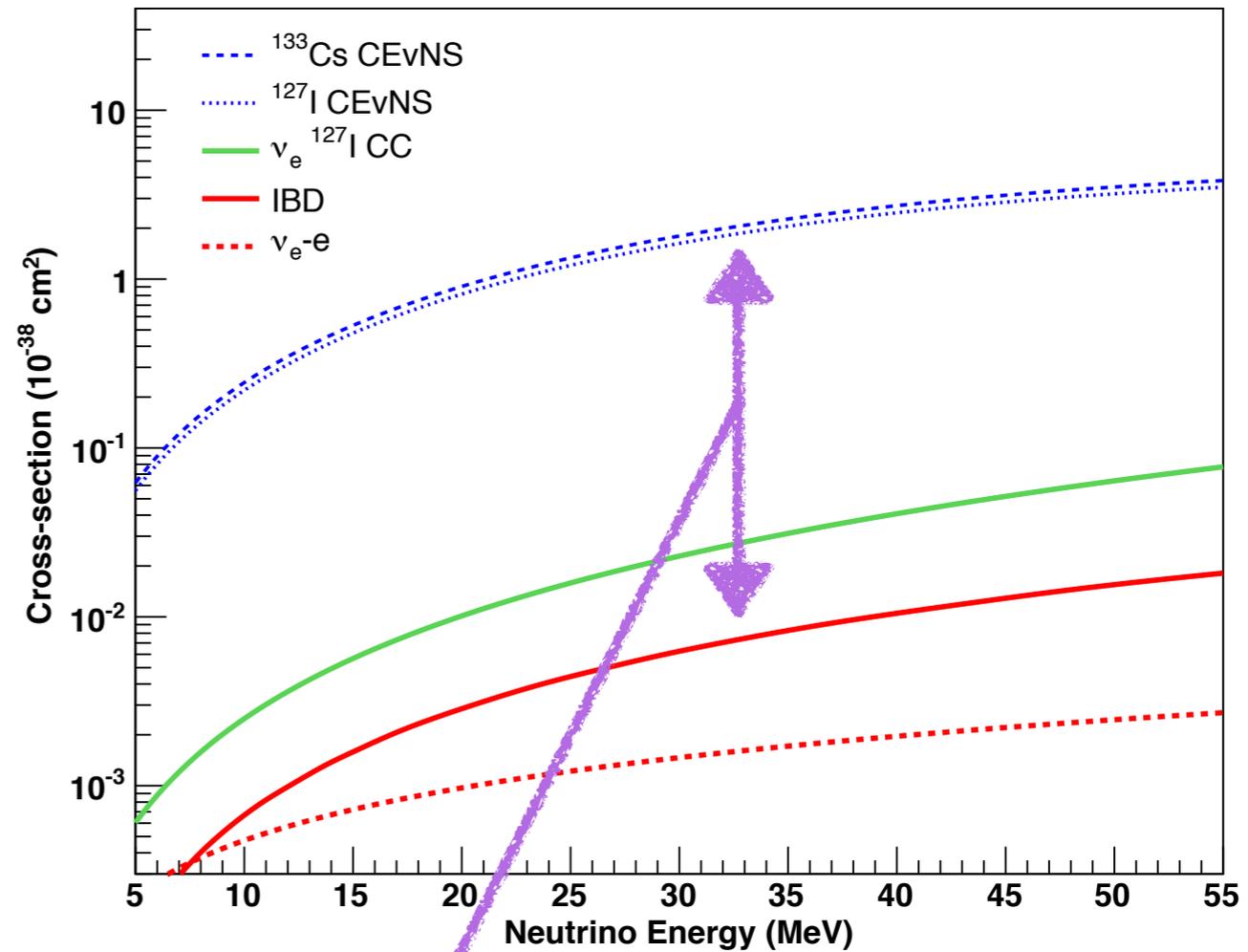


Kavli Institute
for Cosmological Physics
AT THE UNIVERSITY OF CHICAGO

Coherent elastic neutrino-nucleus scattering (CEνNS)

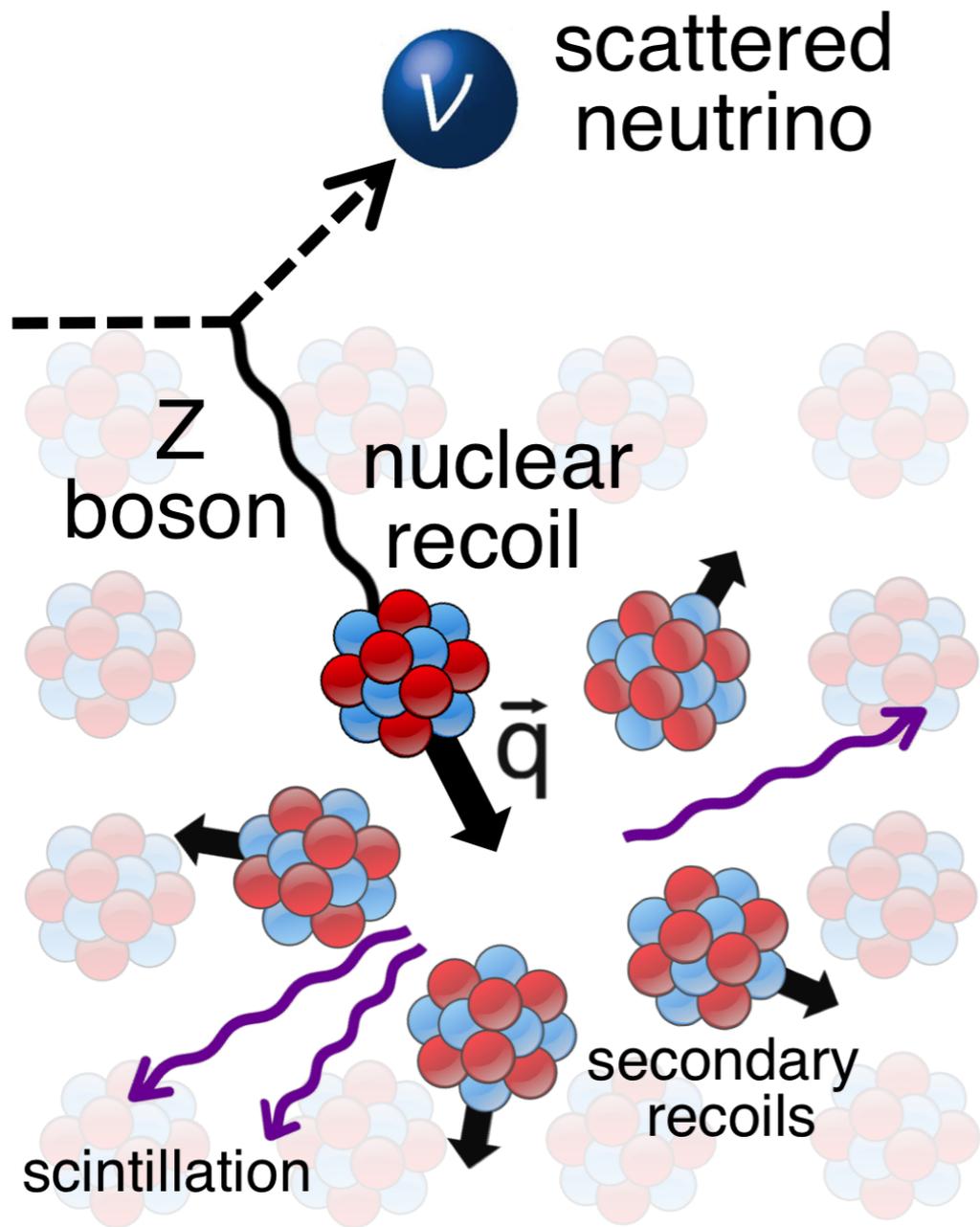
- NC (flavor-independent) process postulated by D.Z. Freedman [1] / Kopeliovich & Frankfurt [2] in 1974
- In a CEνNS interaction, a neutrino scatters off of a nucleus whose nucleons recoil *in phase*, resulting in an enhanced cross section; total cross section scales approximately like N^2

$$\sigma \approx \frac{G_F^2 N^2}{4\pi} E_\nu^2$$



Cross section can be orders of magnitude larger than IBD process used to first observe neutrinos!

“An act of hubris”



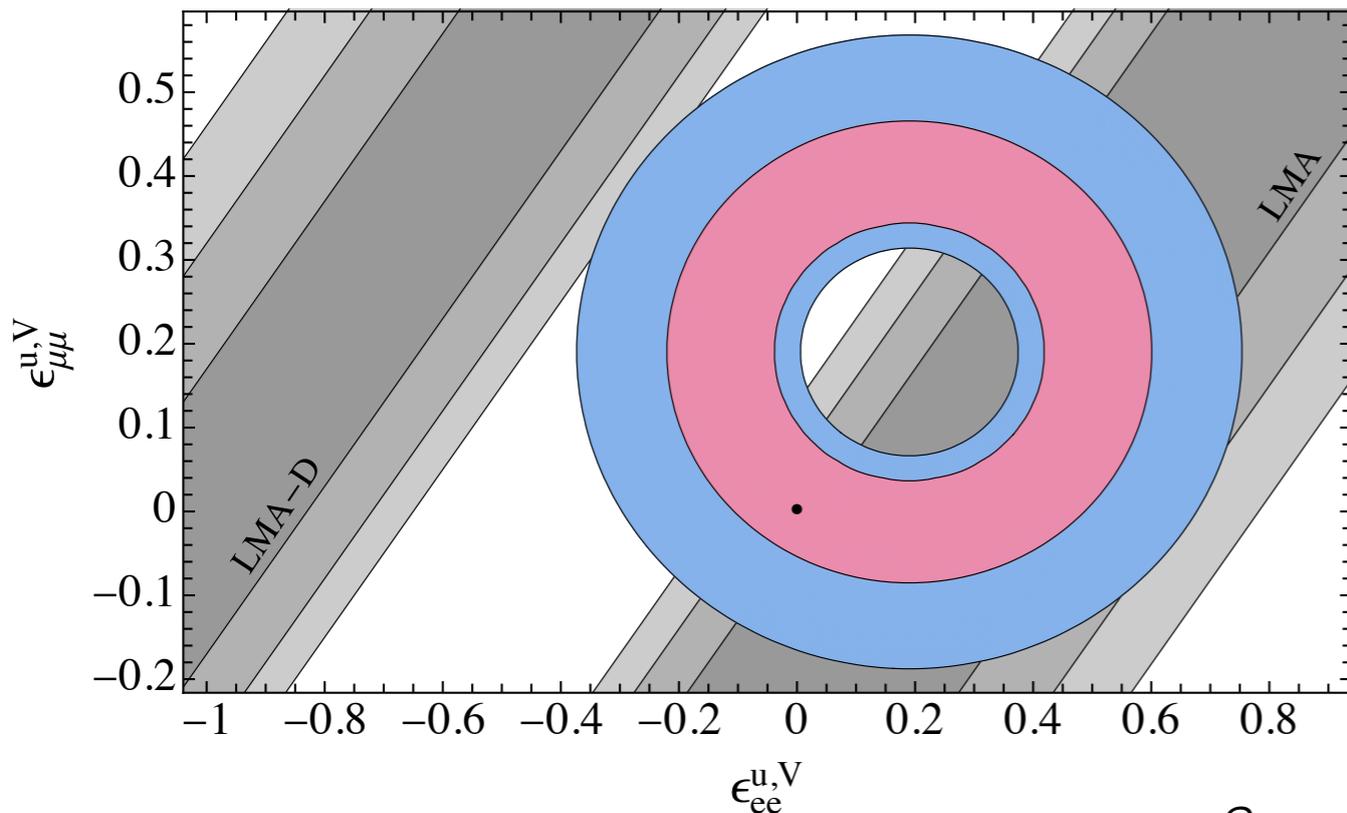
Freedman [1] noted that several factors combine to make CEνNS an exceptionally challenging process to observe

- Only evidence of the interaction is a low-energy recoiling nucleus
 - Heavier nuclei: higher cross section but lower recoil energies
 - Nuclear recoil signal yields are quenched, i.e. reduced compared to signal from electrons of same energy by a factor called the quenching factor (QF)
 - Detector performance hard to calibrate
- Very-low-threshold detectors are very sensitive to backgrounds
 - Neutron backgrounds are particularly dangerous: produce low-energy nuclear recoils just like CEνNS
- Need an appropriate source of neutrinos

Physics from CE ν NS

Supernova physics - Could play a role in dynamics of core-collapse SNe [1] and offers potential way to *observe* SNe neutrinos [2]

Weak mixing angle - Unique probe of Q_W^2 at a unique Q in a region sensitive to dark Z boson models [3]



Non-standard neutrino interactions - explicit dependence on non-universal and flavor-changing neutral currents [4]

Nuclear form factor - Provides a way to measure neutron distributions using neutrino scattering [5], possibly refining nuclear structure models and informing understanding of neutron star EoS [6]

Fundamental properties of neutrinos - sensitivity to effective neutrino charge radius and magnetic moment [7] *and* lift degeneracy of “dark side” solution to θ_{12} that would complicate mass-order determination from oscillation experiments [8]

Neutral-current sterile neutrino search - all-flavor disappearance experiment [9]

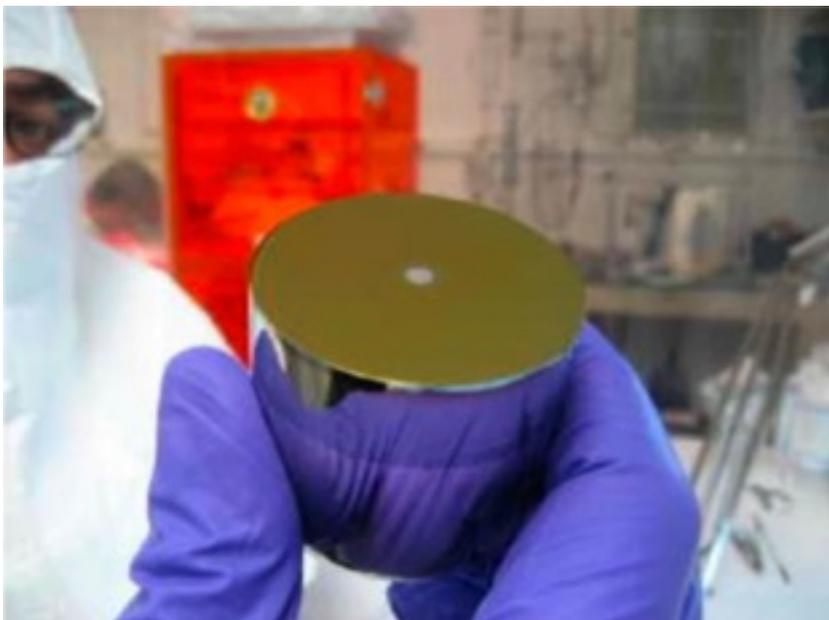
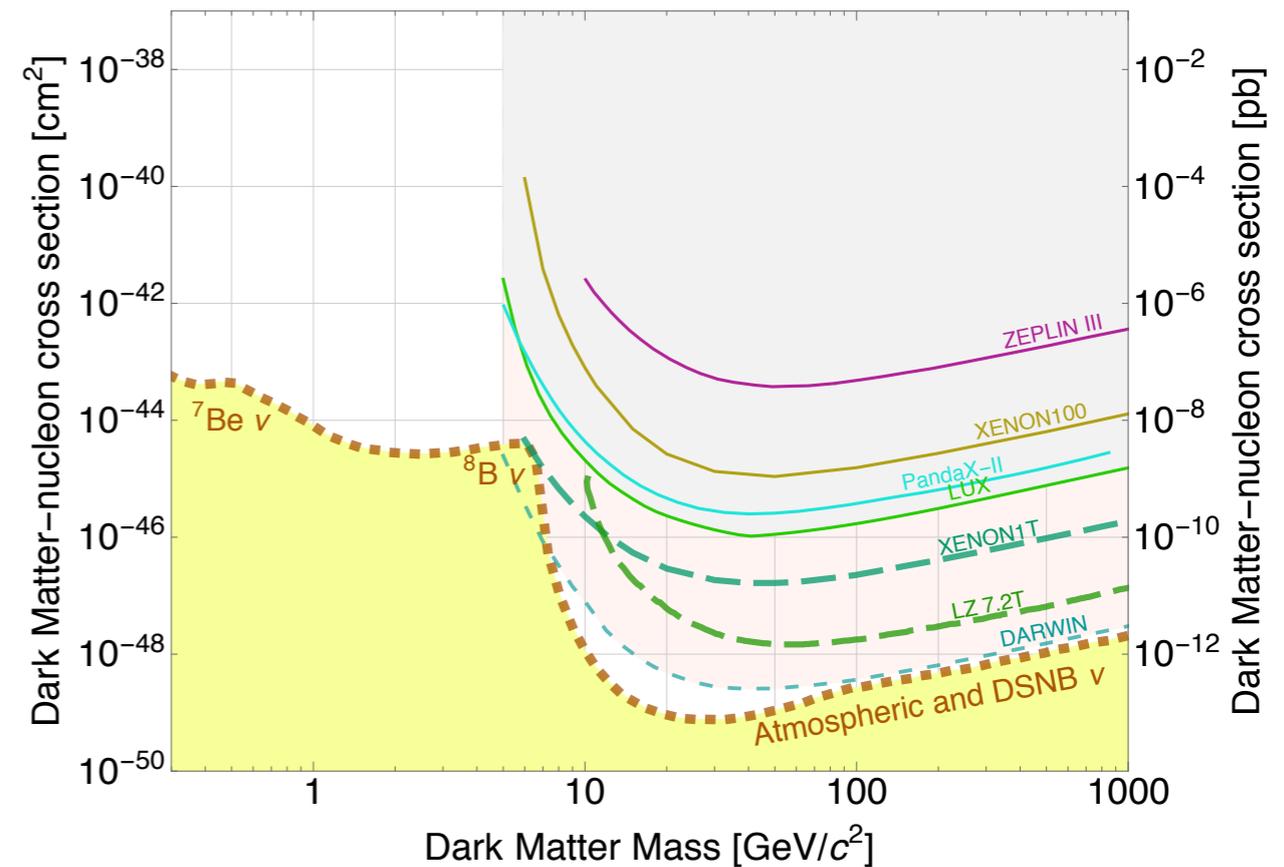
See poster by Sinev & Scholberg on NSI with COHERENT

- [1] D.Z. Freedman, Phys. Rev. D 9 (1974)
- [2] C. Horowitz *et al.*, Phys. Rev. D 68 (2003)
- [3] H. Davoudiasl *et al.*, Phys. Rev. D 89 (2014)
- [4] J. Barranco *et al.*, Phys. Rev. D 76 (2007)
- [5] K. Patton *et al.*, Phys. Rev. C 86 (2012)
- [6] C. Horowitz & J. Piekarewicz, Phys. Rev. Lett. 86 (2000)
- [7] K. Scholberg, Phys. Rev. D 73 (2006)
- [8] P. Coloma *et al.*, Phys. Rev. D 96 (2017)
- [9] A.J. Anderson *et al.*, Phys. Rev. D 86 (2012)

Figure from [8]

CE ν NS becomes a background

- Goodman & Witten recognize utility of CE ν NS-sensitive detectors as potential dark matter detectors [1]
 - DM and CE ν NS interactions are both coherent scattering processes with the same detectable signature (gently recoiling nuclei)
- Numerous instances of proposed CE ν NS detectors turning instead into competitive DM searches



P.S. Barbeau, Ph.D. thesis (UChicago 2009)

- Tremendous advances in detector technology to build more sensitive DM searches
- Next generation of WIMP detectors will begin to be sensitive to CE ν NS from ^8B solar neutrino flux
 - This “neutrino floor” brings the CE ν NS and DM relationship full circle

Physics from CE ν NS

See talks this session by Henry Wong and Omar Miranda for in-depth discussion of CE ν NS physics

Saturday morning session on supernova neutrinos

Enter: The COHERENT Collaboration

- Goal: **unambiguous observation** of CE ν NS using multiple nuclear targets / detector technologies
 - Leverage detector advances from dark-matter community
 - Utilize intense, pulsed neutrino source provided by Spallation Neutron Source (SNS)
 - Use of different nuclear targets allows for measurement of characteristic N^2 cross-section dependence and some added analysis advantages
- Pioneering CE ν NS detector: CsI[Na]

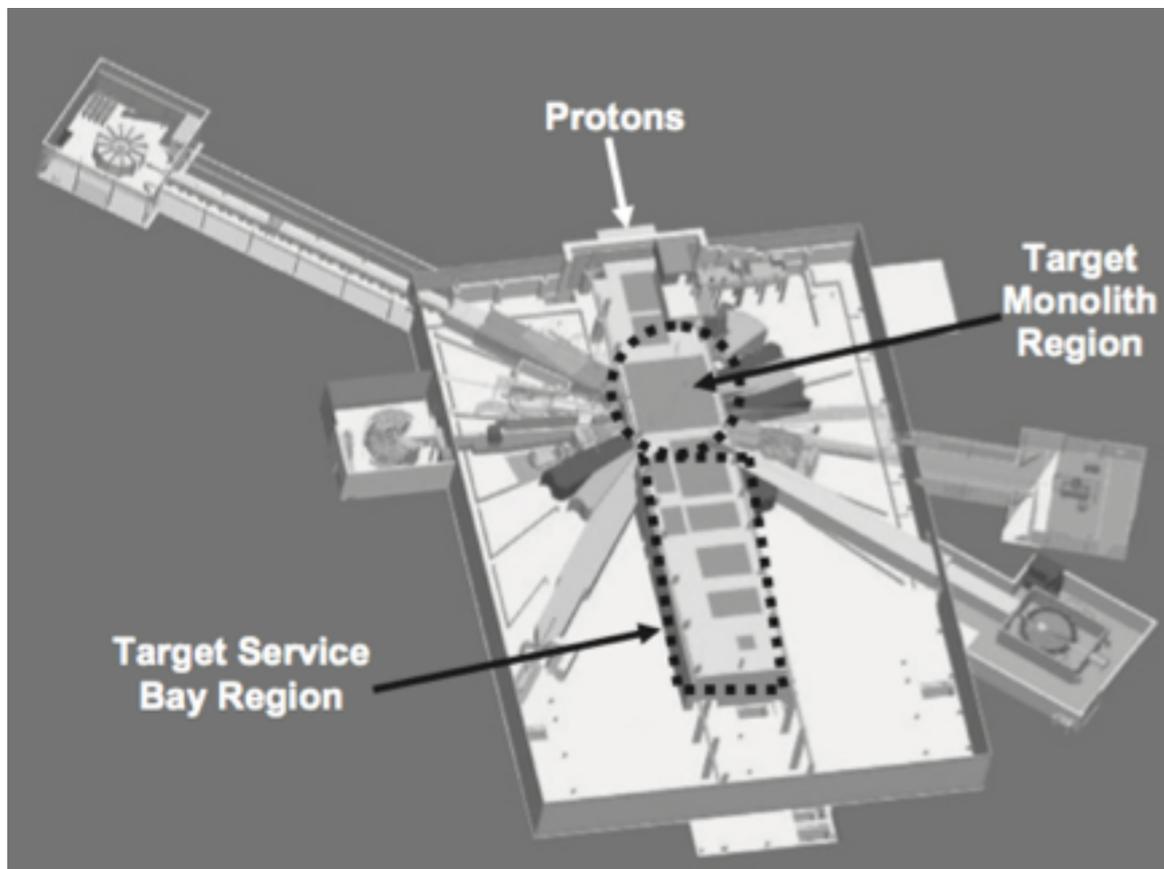
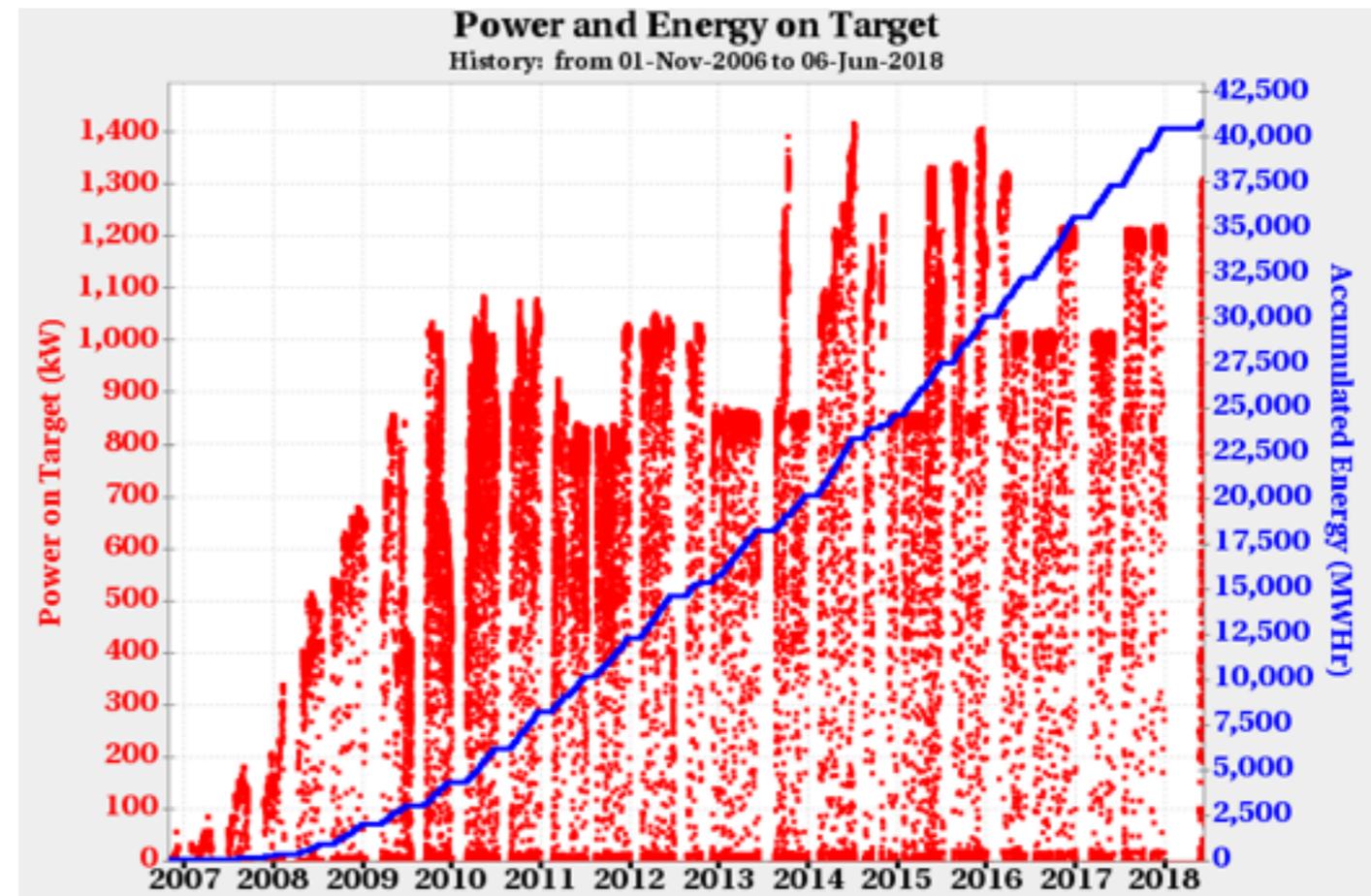


The Spallation Neutron Source



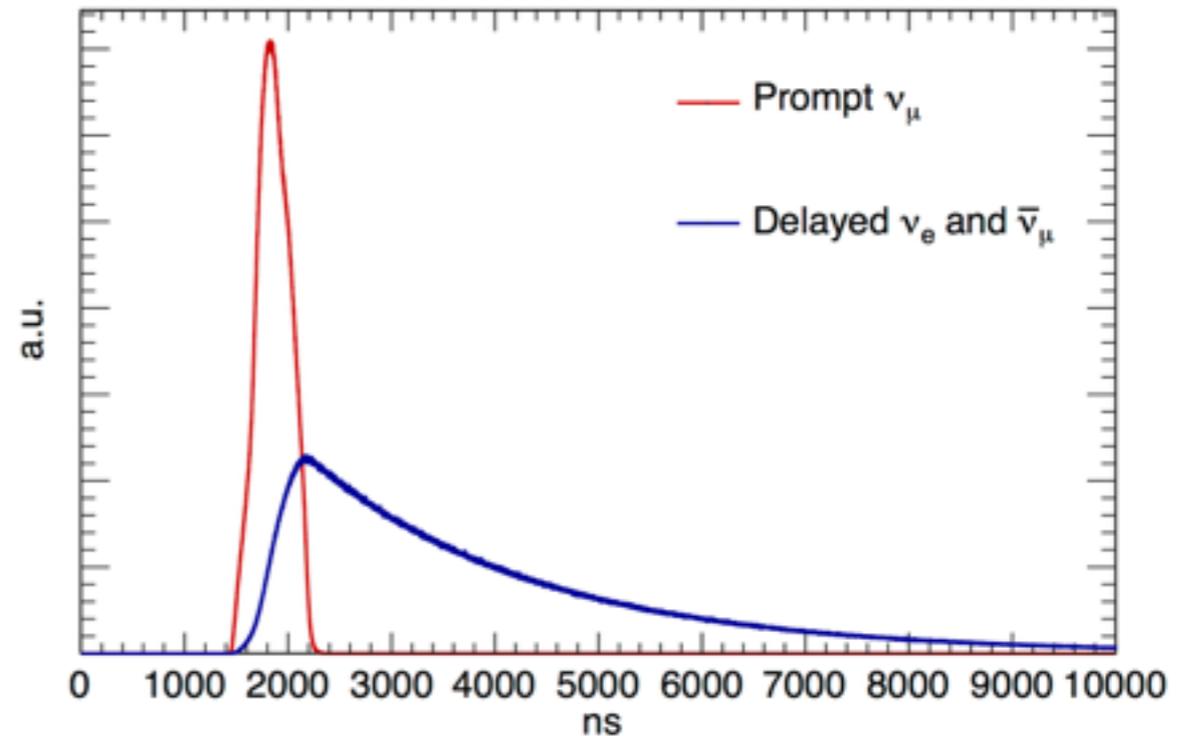
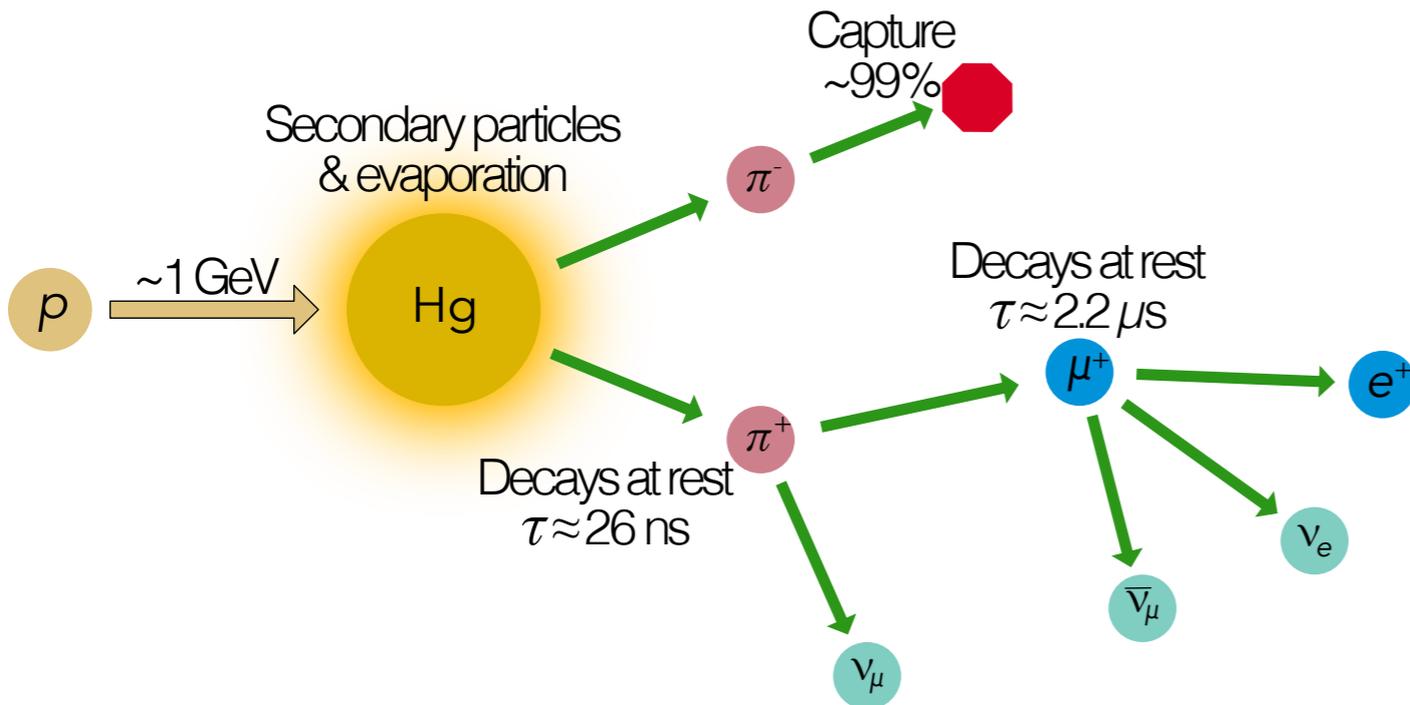
- Located at Oak Ridge National Lab, near Knoxville, TN, USA
- The SNS bombards a liquid mercury target with a ~ 1 -GeV proton beam pulsed at 60 Hz; each beam pulse is ~ 700 -ns wide
- Neutrinos are produced by decay of **stopped** *pions and muons*, resulting in flux with well-defined spectral and timing characteristics

The Spallation Neutron Source



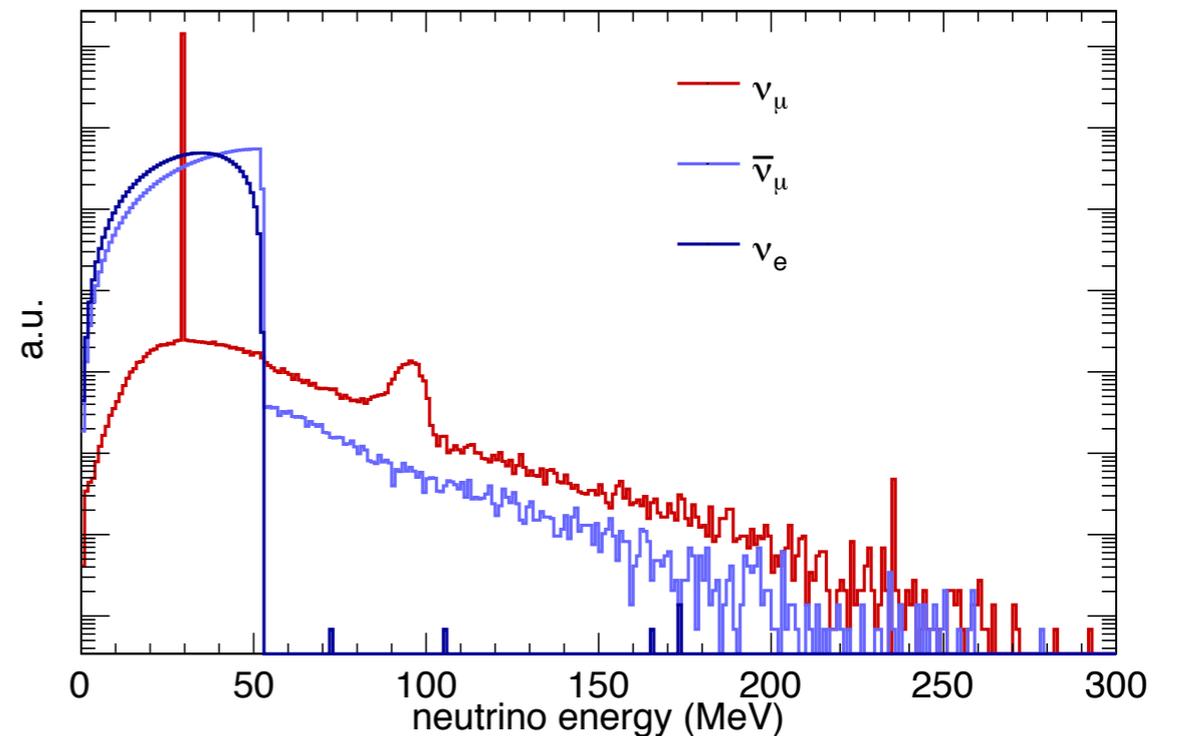
Most intense pulsed neutron source in the world

The Spallation Neutrino Source



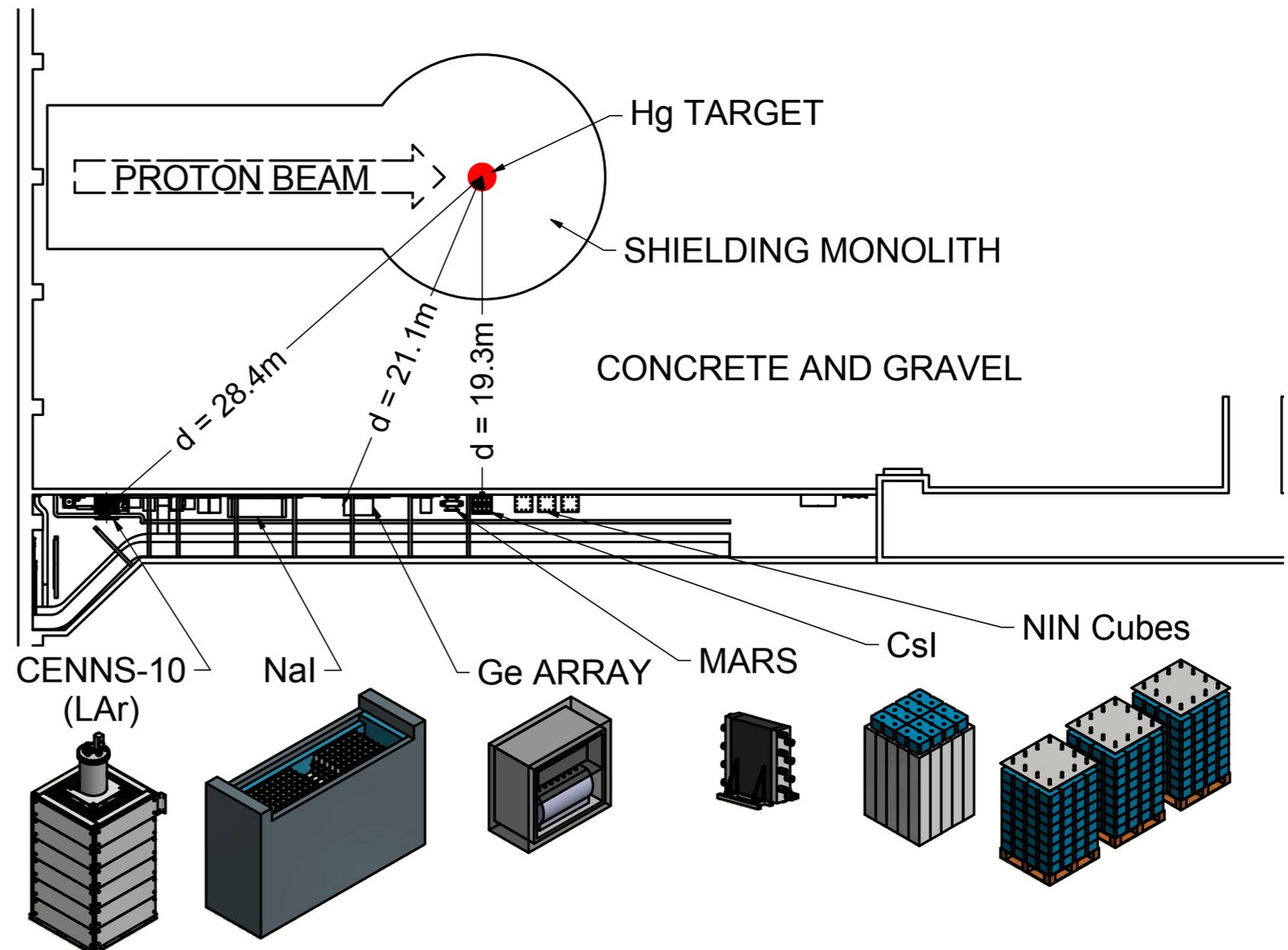
- High-fidelity GEANT4 simulation starts with proton beam; energy spectra very near analytical approximations
- Massive reduction in steady-state backgrounds through timing ($\mathcal{O}(1000)$); facility-wide timing signal can be used to trigger DAQ, both during beam-on and -off periods

Poster on source simulation and flux measurement possibilities from R. Rapp



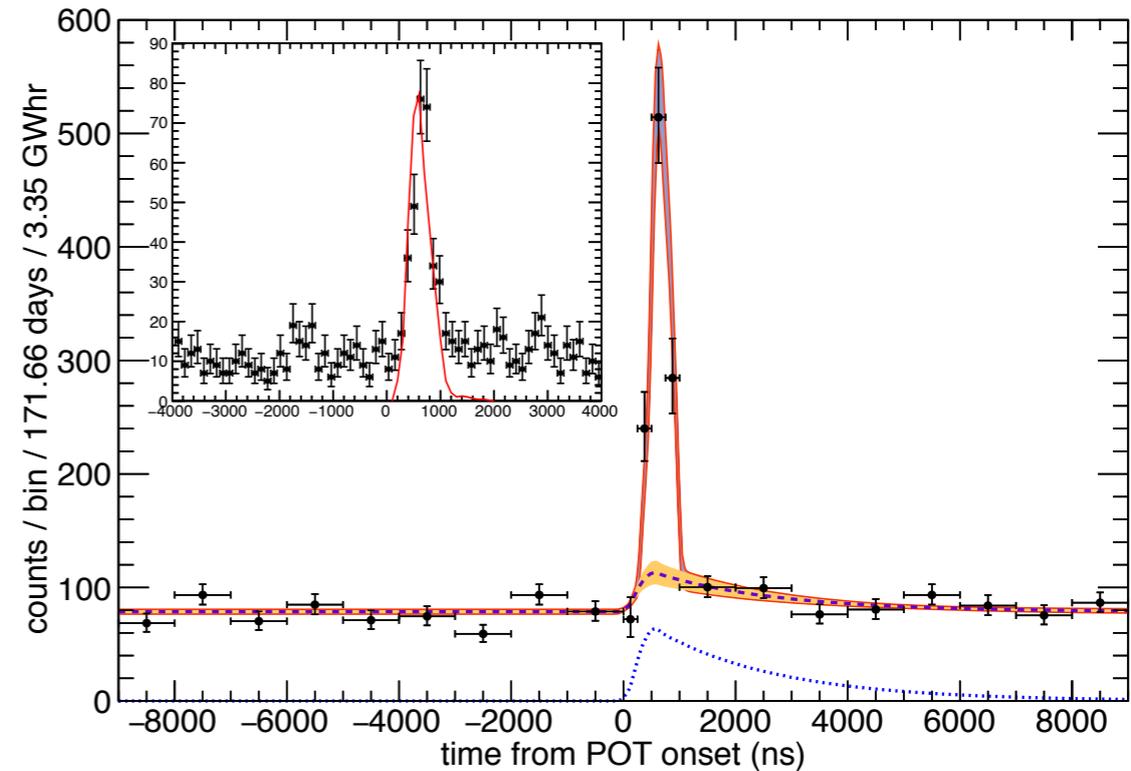
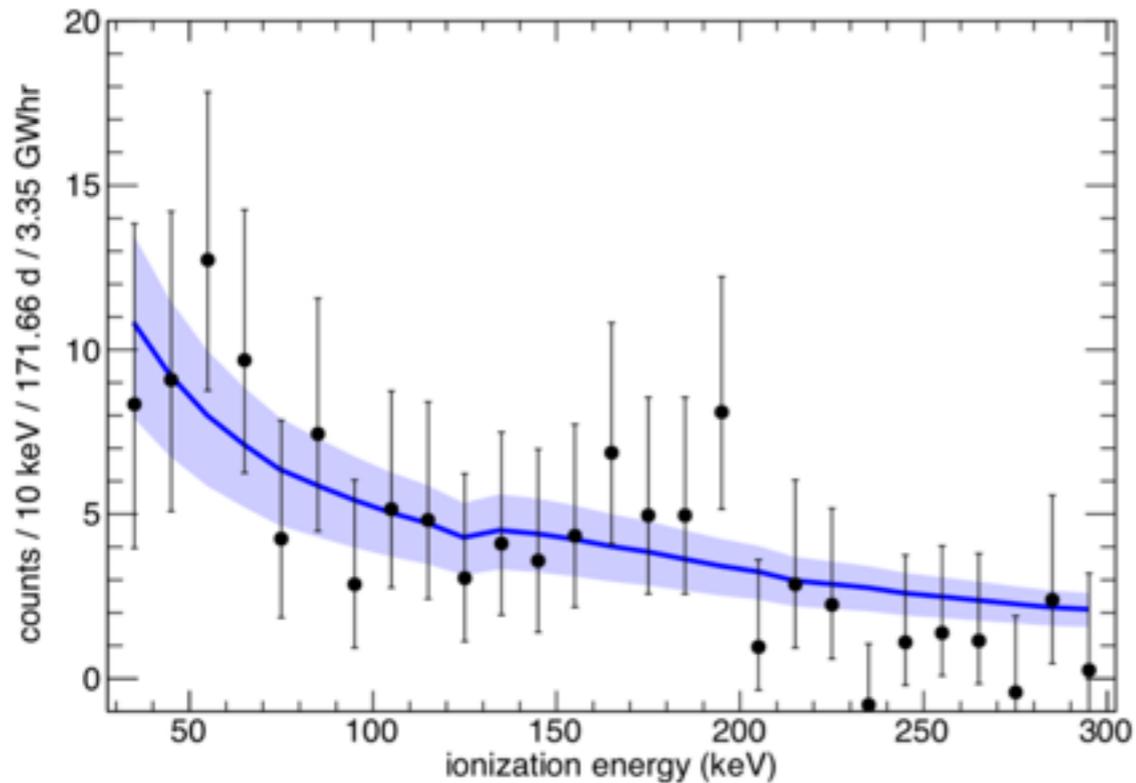
Siting and backgrounds

- Backgrounds depend significantly on siting at SNS
 - Extensive background measurement campaign
- COHERENT experiments located in a ~~basement hallway~~ *neutrino alley*
 - ~8 m.w.e. overburden
 - 20- to 30-m from target
- Primary backgrounds in neutrino alley:
 - Prompt SNS neutrons
 - Neutrino-induced neutrons (NINs)

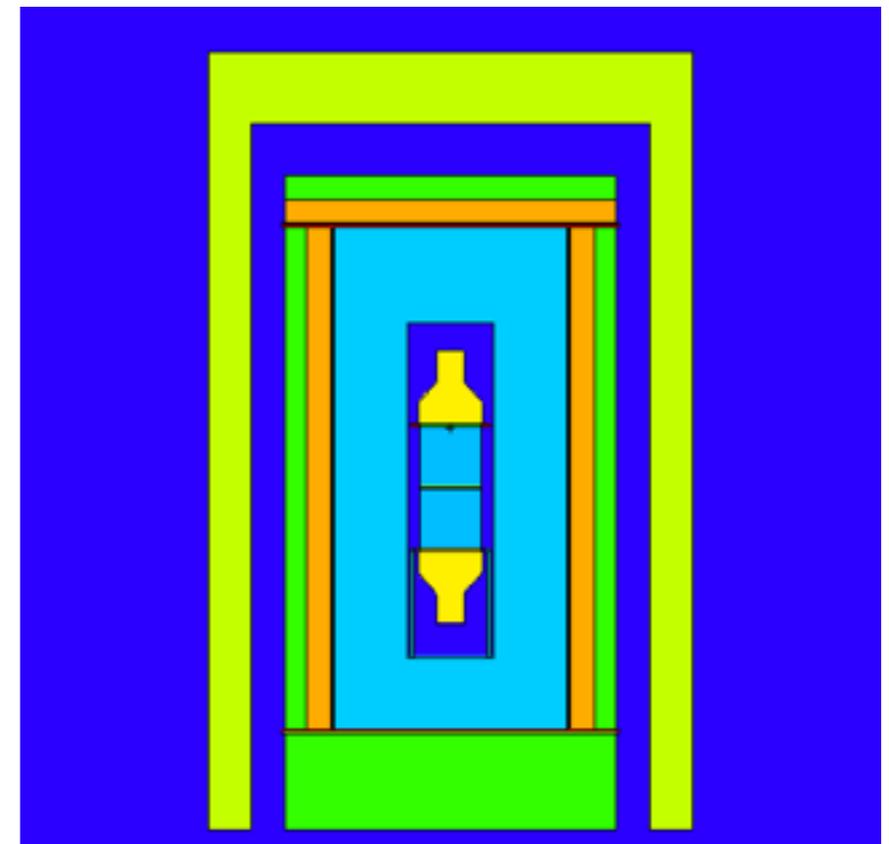


Poster from J. Raybern on neutron background measurement efforts with MARS

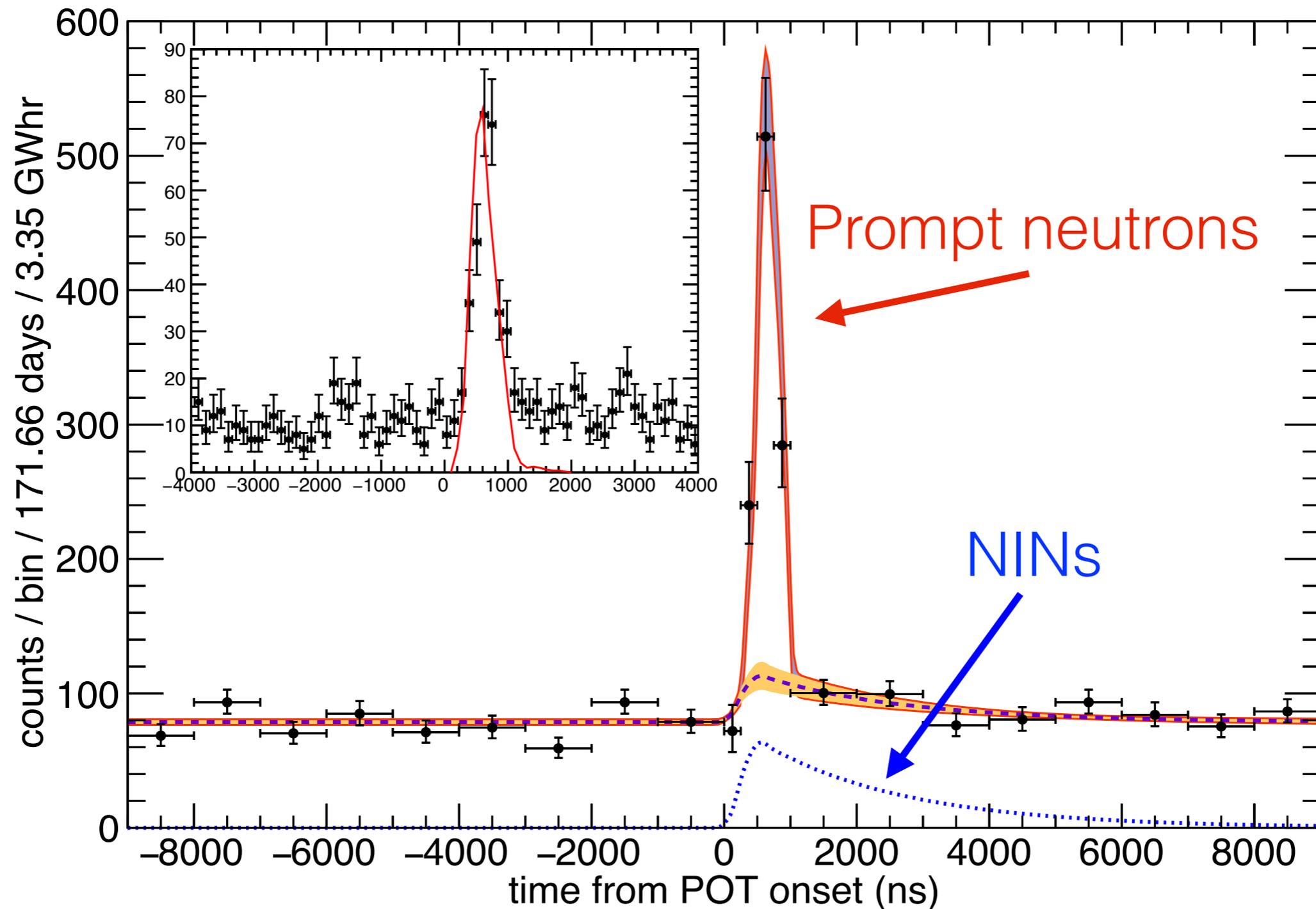
In situ measurement of neutron backgrounds



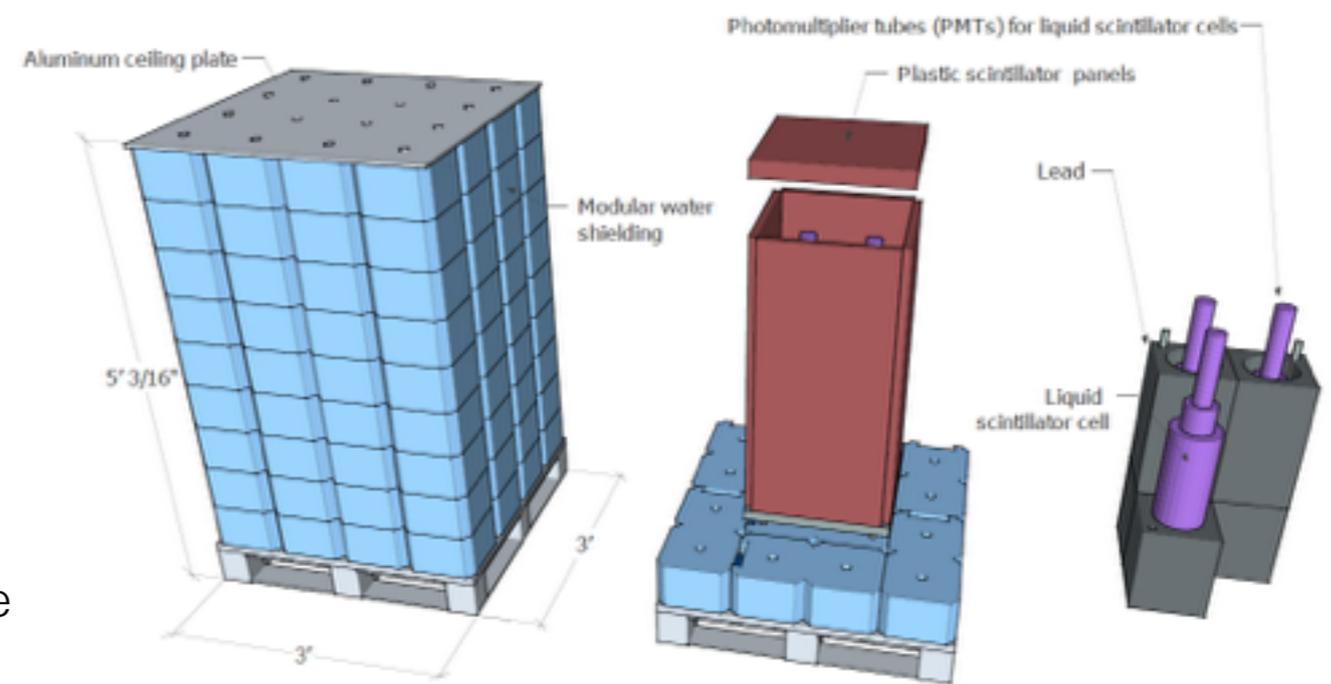
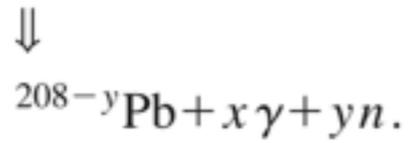
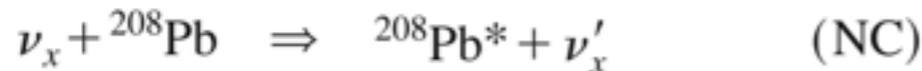
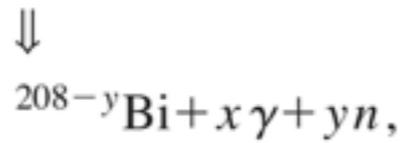
- Prior to CE ν NS search, neutron detection system installed at location of CsI[Na] detector
- Data informed model of prompt SNS neutron energy distribution
- Established understanding of beam timing w.r.t. SNS timing signal



In situ measurement of neutron backgrounds

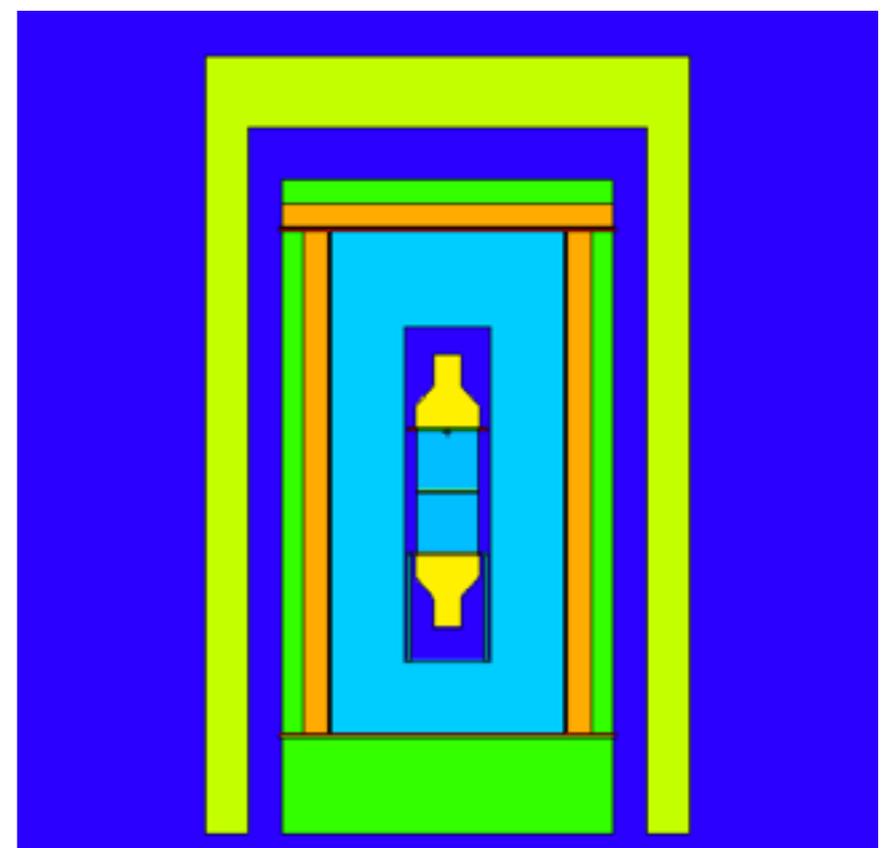


Neutrino-induced neutrons (NINs)



Poster from C. Awe on neutrino cubes at the SNS

- Dominant background for CEνNS measurement with naïve shielding configuration, but interesting physics of its own
 - Possible role in nucleosynthesis in certain astrophysical environments [1]
 - NIN production on Pb is the fundamental mechanism by which HALO intendeds to detect supernova neutrinos [2]
 - Process has never before been measured, considerable variation in theoretical predictions (~3x) [3]
- *In situ* measurements give rate limit, plus ongoing measurement of process with “neutrino cubes”



[1] Y-Z. Qian *et al.*, Phys. Rev. C 55 (1997)

[2] C.A. Duba *et al.* J. Phys. Conf. Series 136 (2008)

[3] C. Volpe, N. Auerbach, G. Colò, and N. Van. Giai, Phys. Rev. C 65 (2002)

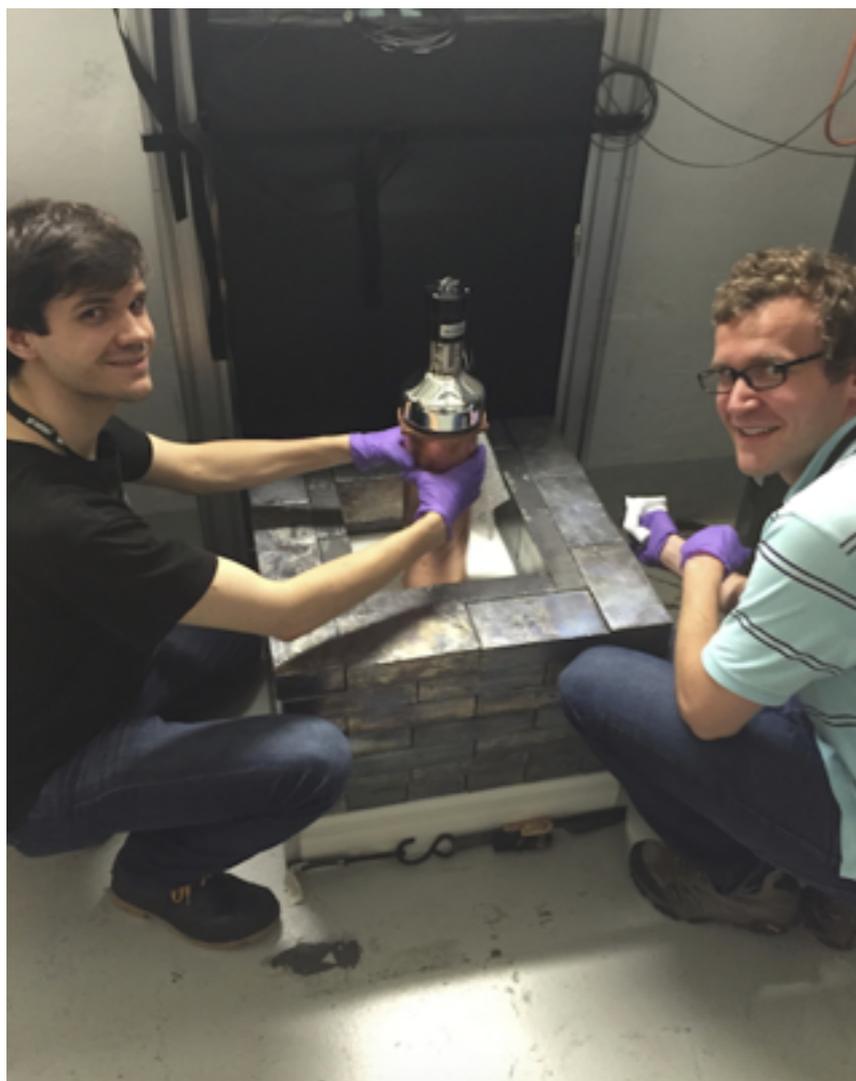
NIN pathways from S.R. Elliott, Phys. Rev. C (2000)



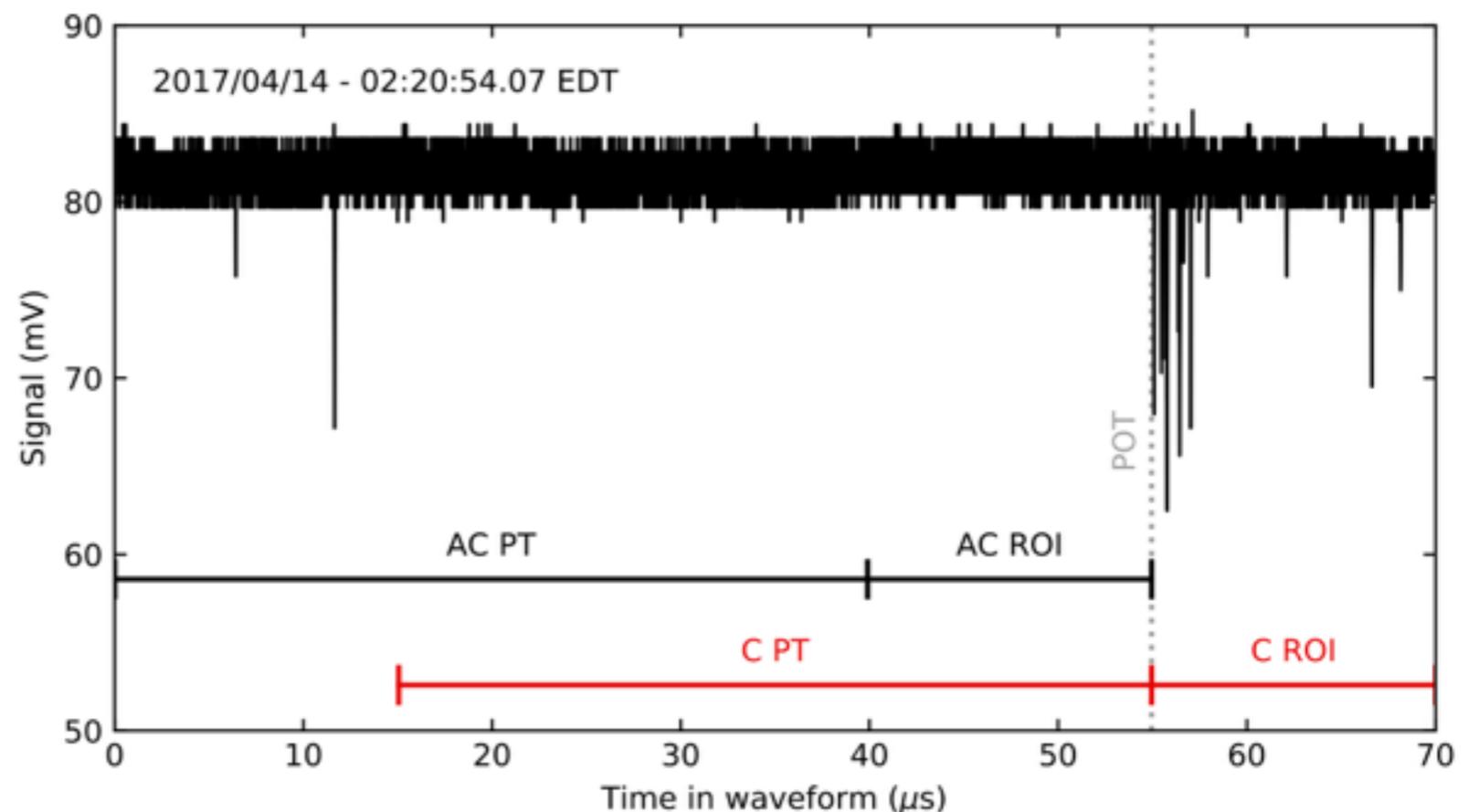
CE ν NS with CsI[Na]



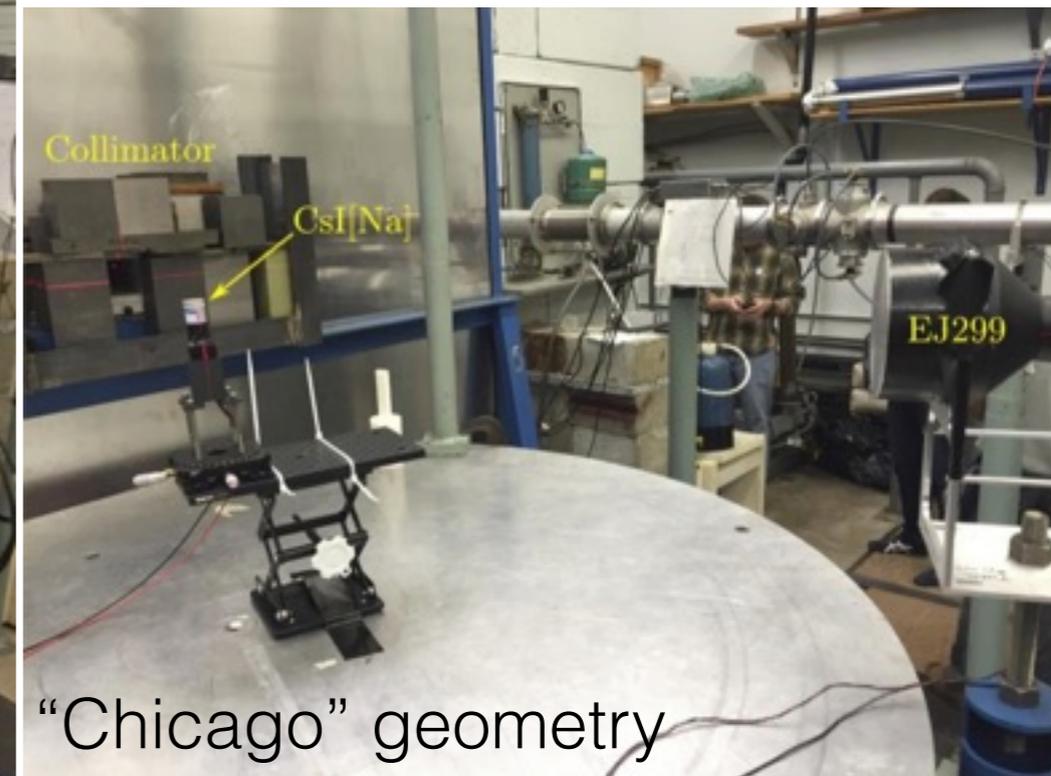
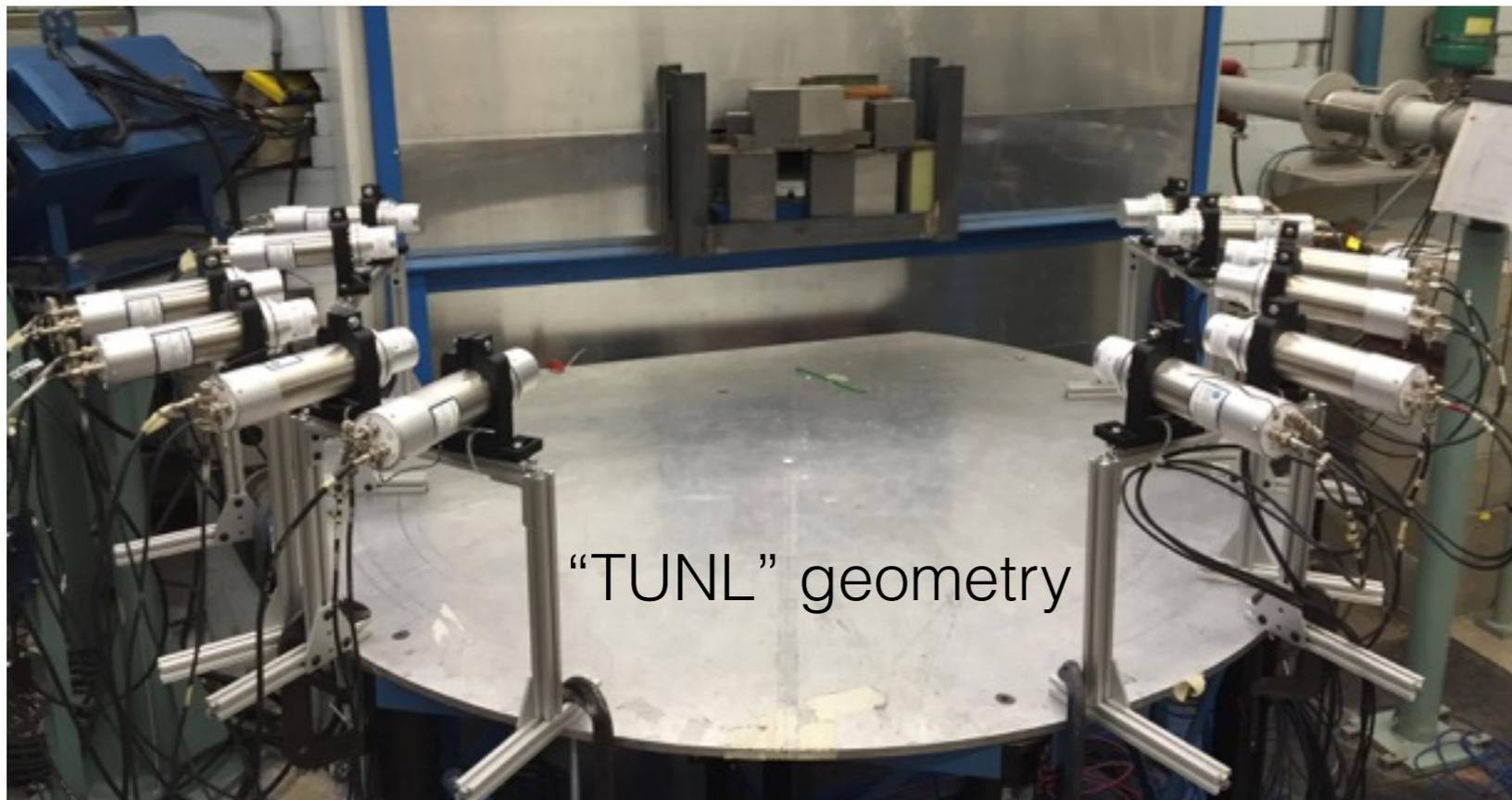
Deployed to SNS in June 2015



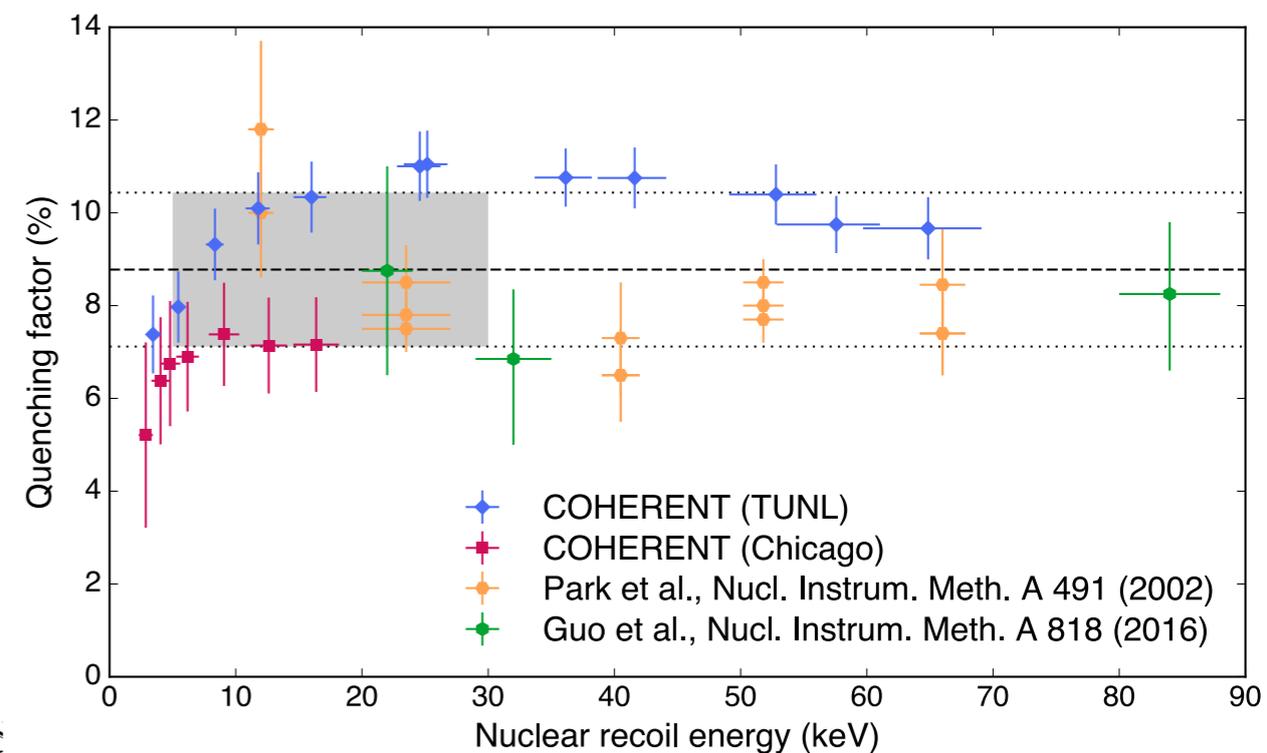
- 14.6-kg crystal made from low-background salts, encased in electroformed-copper can with PTFE reflector and synthetic silica window, surrounded by neutron and gamma shielding, including low-activity lead
- Development led by University of Chicago [1]
- Output of super-bialkali PMT with $\sim 30\%$ QE digitized for $70 \mu\text{s}$, triggered by SNS timing signal



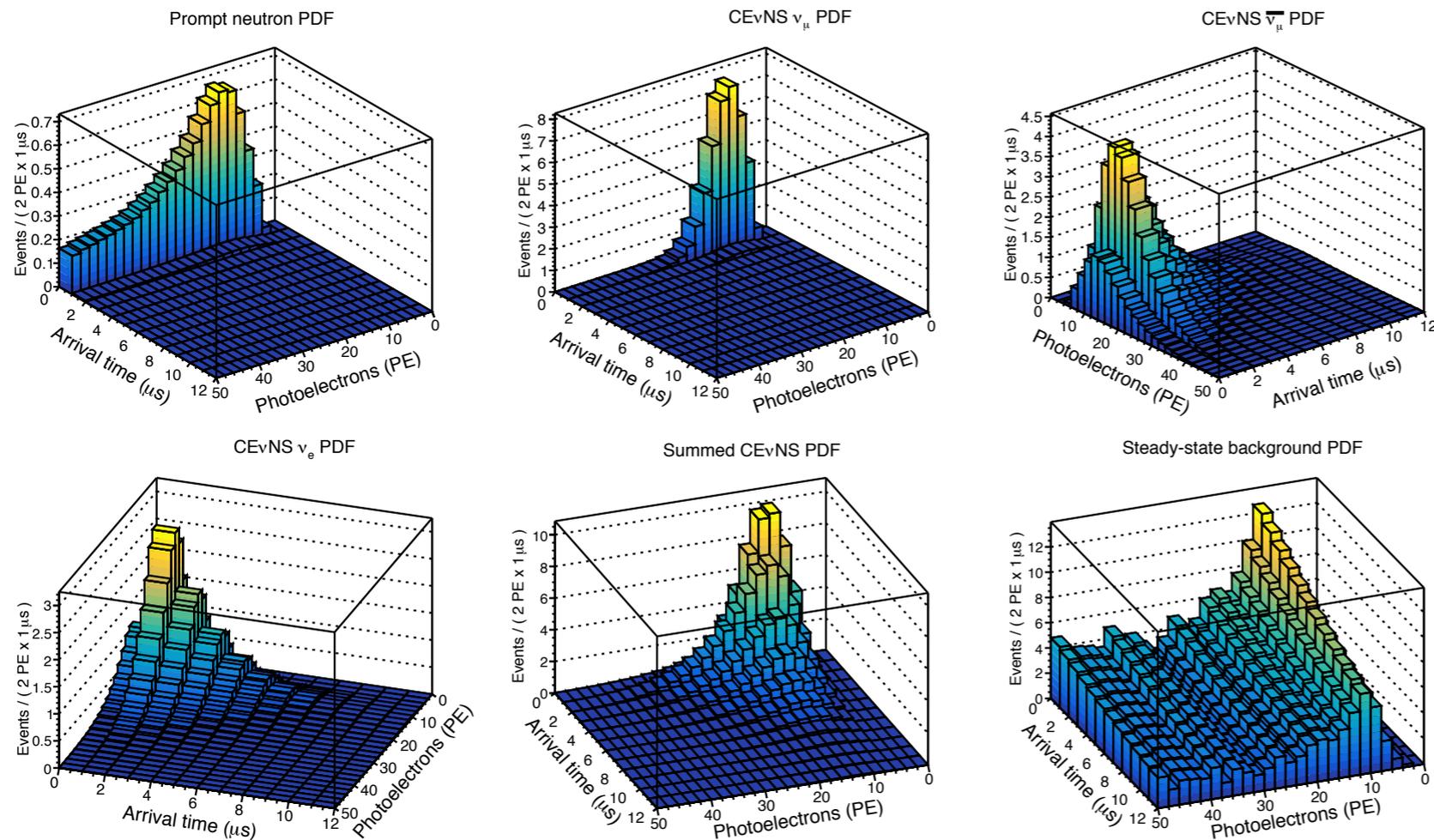
Quenching factor measurements at TUNL



- Elastically scatter quasi-monoenergetic neutrons into “backing detectors” at known angles
 - Each backing detector associated with events having well-defined nuclear recoil energies
- Determine QF from global values in range from 5 to 30 keVnr: $8.78 \pm 1.66\%$

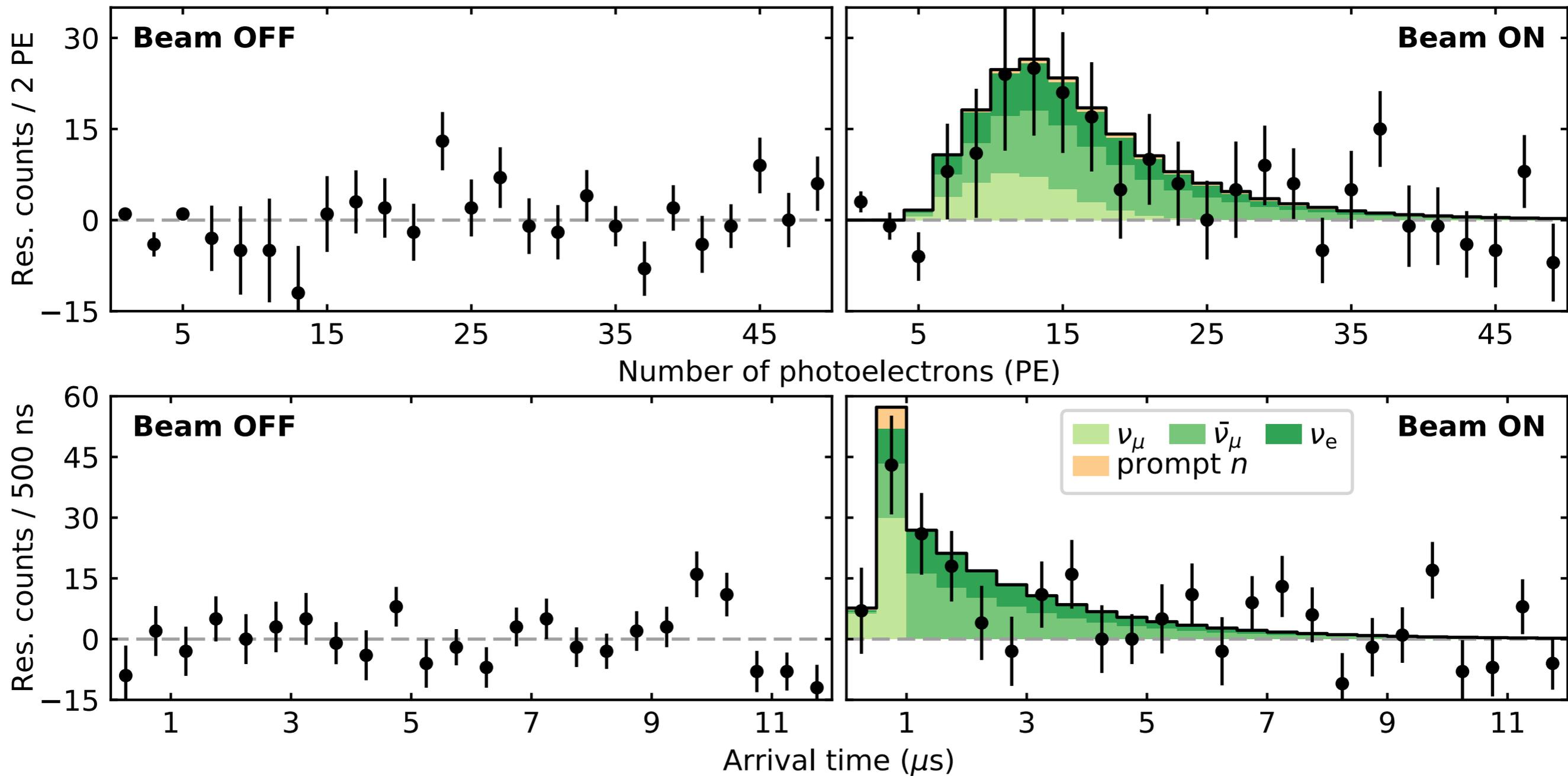


Rate and shape estimates



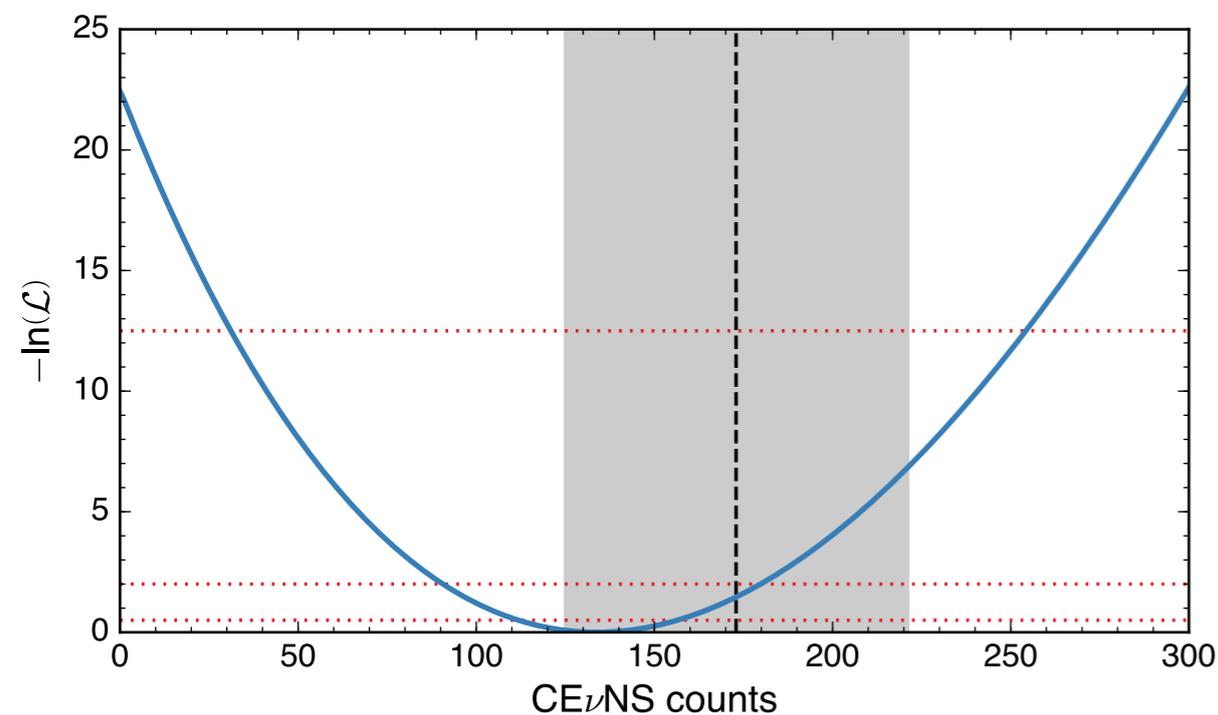
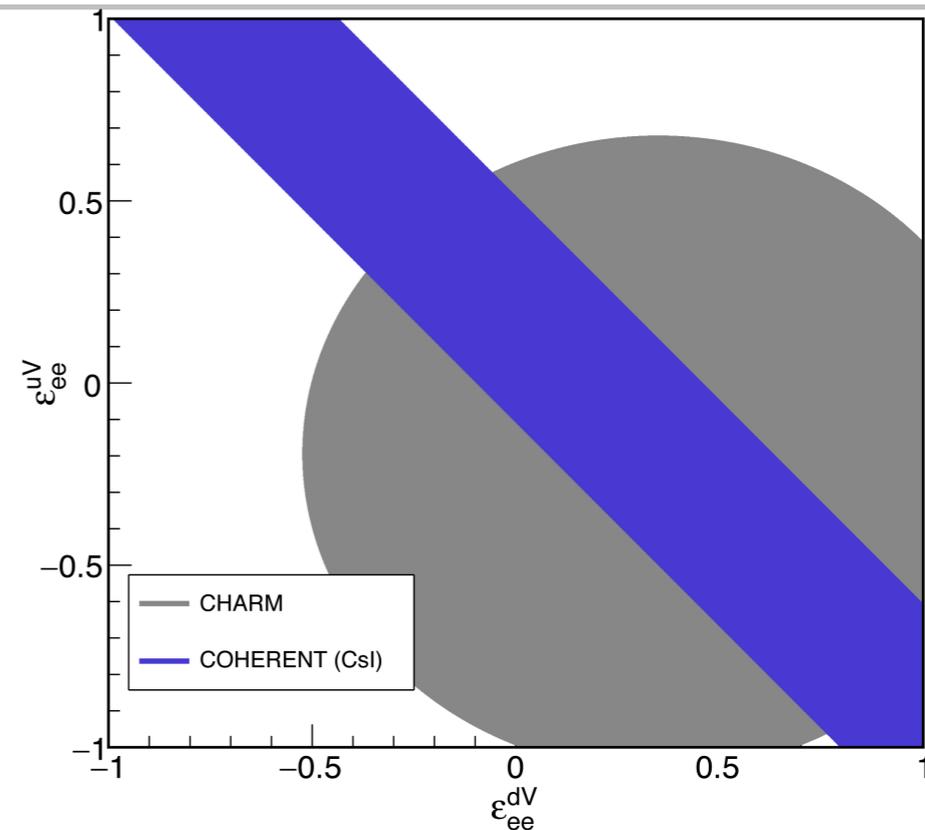
- Pulsed nature of beam facilitates analysis in time domain
- 2-D analysis (energy, time) makes use of all available information
- Ultimately performed binned 2-D profile likelihood analysis using PDFs shown here
 - Assumes Standard Model
 - Incorporates knowledge of detector response, analysis acceptance, etc

SM prediction and data



Results

- Beam exposure: ~ 6 GWhr, or $\sim 1.4 \times 10^{23}$ protons on target (0.22 grams of protons)
- Analyzed as a simple counting experiment
 - 136 ± 31 counts
- 2-D profile likelihood analysis
 - 134 ± 22 counts, within $1\text{-}\sigma$ of SM prediction of 173 ± 48
 - Null hypothesis disfavored at $6.7\text{-}\sigma$ level relative to best-fit number of counts
- Able to further constrain some NSI parameters



Dominant systematic uncertainties on predicted rates

Quenching factor	25%
ν flux	10%
Nuc. form factor	5%
Analysis acceptance	5%

CE ν NS observation data release

- Data that constituted CE ν NS observation has been packaged and is publicly available
 - <http://dx.doi.org/10.5281/zenodo.1228631>
 - <https://coherent.ornl.gov>
- Should include all information necessary to perform further analyses on CsI[Na] data
 - Binned data for coincidence and anticoincidence regions for both SNS on and off; prompt-neutron model
 - Descriptions and values for relevant systematics
- Collaboration intends to continue practice of data releases

The Zenodo logo consists of the word "zenodo" in a white, lowercase, sans-serif font, centered on a solid blue rectangular background.

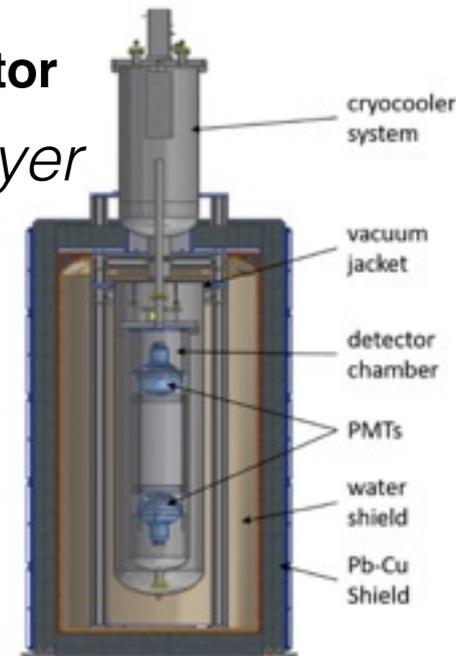
COHERENT physics moving forward

- Measure NINs cross section in ^{208}Pb , ^{56}Fe
 - Upgrades to detection system planned in cooperation with PROSPECT
- Measure ^{127}I CC cross section
 - 185-kg NaI ν E collecting low-gain CC data now; continue in 2-T phase in parallel with high-gain mode
 - Sensitivity to g_A quenching with $Q \sim \mathcal{O}(10 \text{ MeV})$
- N^2 dependence of $\text{CE}\nu\text{NS}$ cross section
 - Several distinct N values represented in COHERENT suite of experiments
 - 22-kg LAr detector already collecting $\text{CE}\nu\text{NS}$ data, plans for 10 kg of Ge PPCs and 2-T NaI[Tl]
- Begin to perform precision $\text{CE}\nu\text{NS}$ measurements
 - High-resolution, low-threshold detectors, such as Ge PPCs, enable access to exciting physics, e.g. electromagnetic properties of neutrinos

CENNS-10 LAr detector

Poster from J. Zetlemoyer

Also introduces future plan for 1-T LAr detector



NaI ν E: NaI[Tl] neutrino experiment

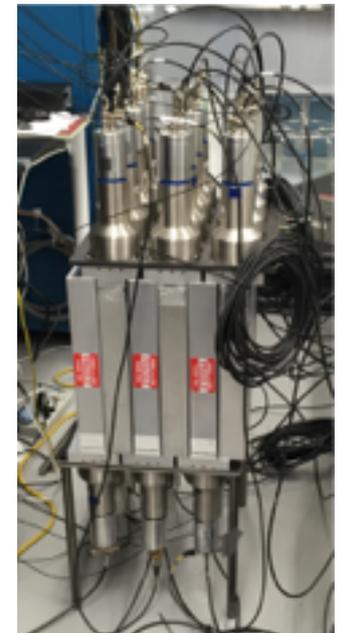
Posters

S. Hedges

Measurement of ^{127}I CC with 185-kg array

D. Markoff

Multi-ton NaI[Tl] array for $\text{CE}\nu\text{NS}$ measurement

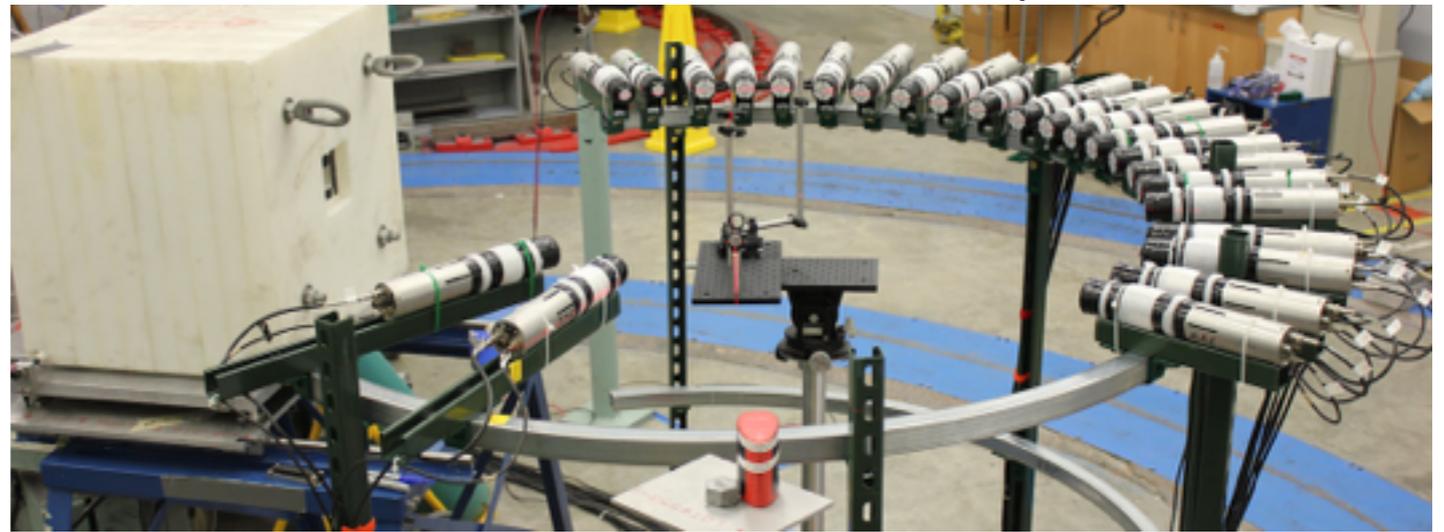


Reducing dominant systematic uncertainties

Quenching factors

- Understanding of QF is crucial for *all* CE ν NS measurements
 - Reanalyzing original data and collecting new data to resolve discrepancy in COHERENT QF measurements for CsI[Na]
 - Some data already collected and future measurements planned for Ge and NaI[Tl]

Poster from Long Li on QF measurement in Ge



- Indirect approaches to flux determination possible (e.g., improved input for models or direct measurement of pion production at SNS)
- Conceptual design stages of a D₂O detector for neutrino alley relying on CC interaction on D
 - D cross section is relatively well understood theoretically [1] and previous measurements agree with predictions [2]

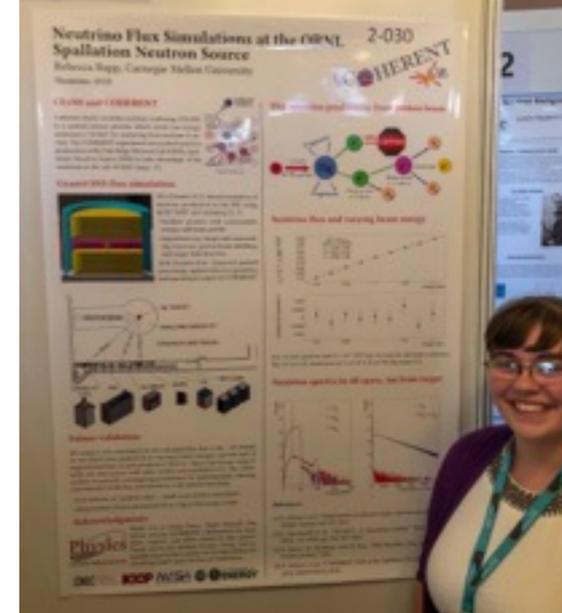
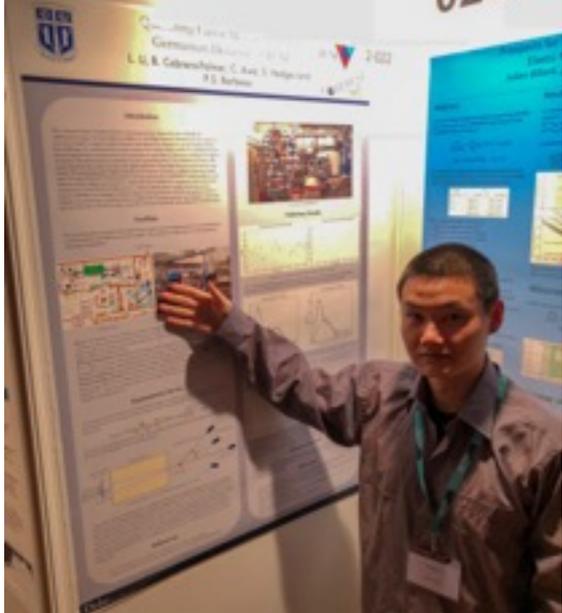
Poster on source simulation and flux measurement possibilities from R. Rapp

ν flux
normalization

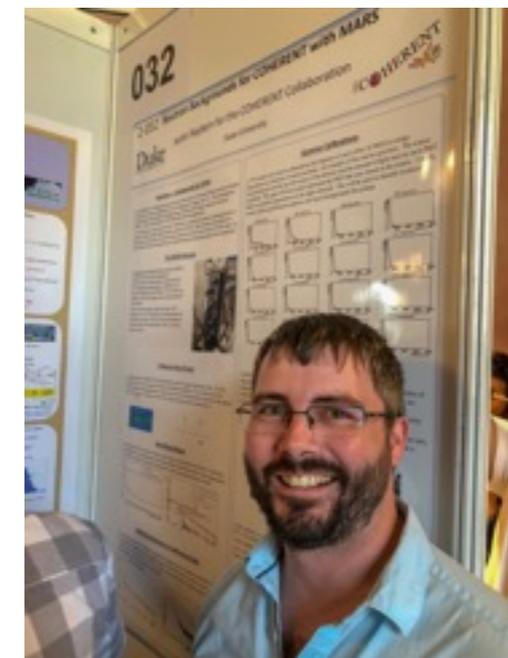
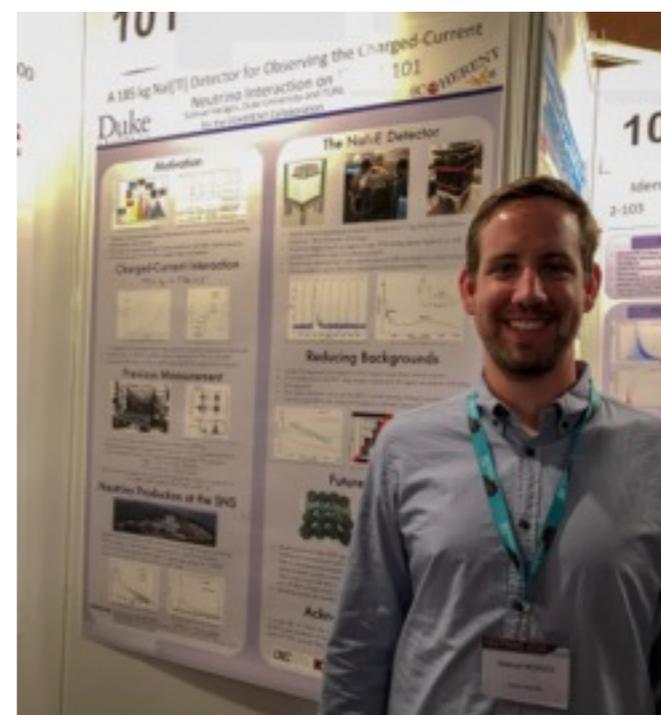
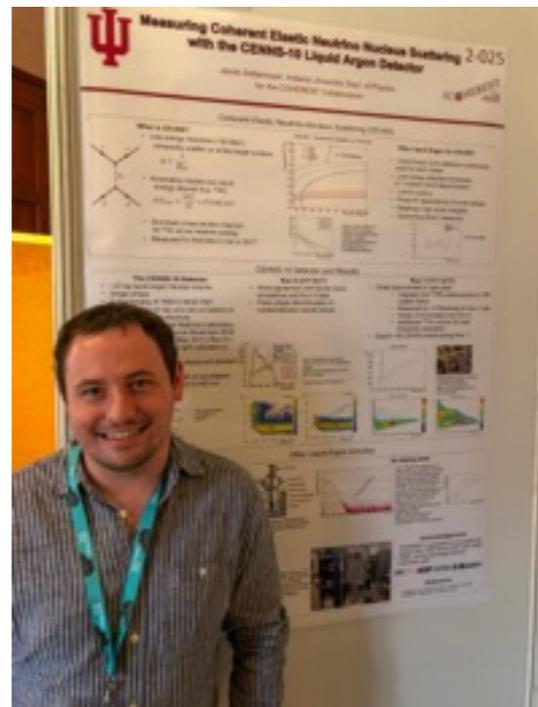
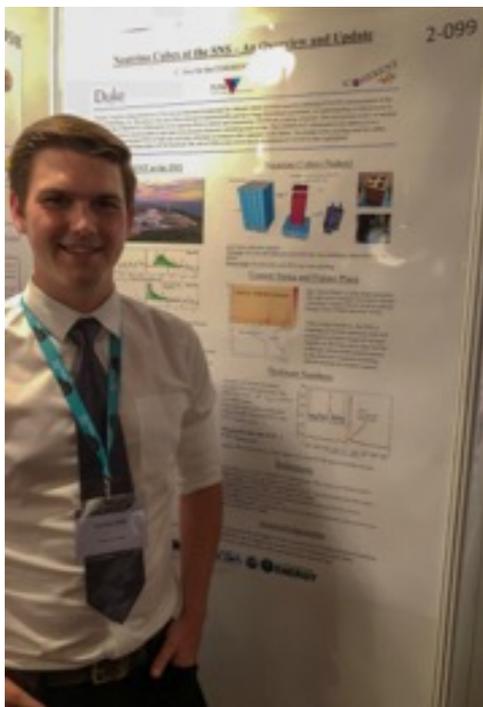
Only the beginning...

- CE ν NS predicted in 1974 but unobserved until 2017
 - Observed at 6.7- σ level using 14.6-kg CsI[Na] scintillator deployed at pulsed, stopped-pion ν source (SNS)
- COHERENT continues to search for CE ν NS with numerous detectors (LAr, NaI[Tl], Ge PPCs) in addition to several other efforts
 - Working towards performing *precision* CE ν NS measurements
- Many other groups seeking observation with many different kinds of detectors, different neutrino sources
 - Examples: CONNIE, CONUS, MINER, Nu-CLEUS, nuGEN, RICOCHET, RED-100
 - These efforts are *complementary!* Joint analyses using different detectors and/or sources are very powerful [1]
- Tremendous amount of physics left to be done with CE ν NS
 - Important complement to oscillation measurement program through lifting of LMA-D ambiguity





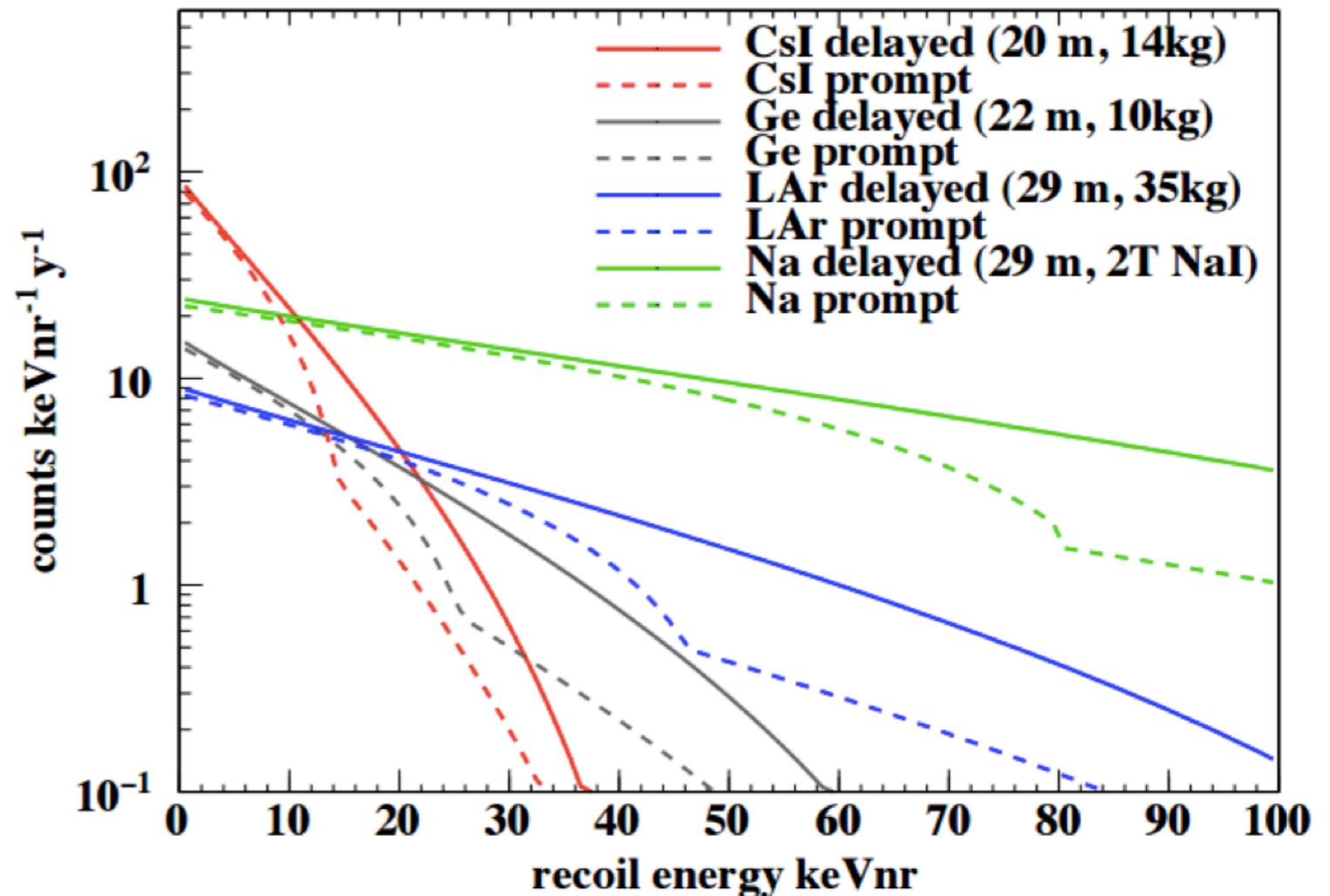
Long Li (Gleb Sinev represented by Rebecca Rapp
 Connor Awe Kate Scholberg) Justin Raybern
 Jacob Zettlemoyer Sam Hedges



Backup

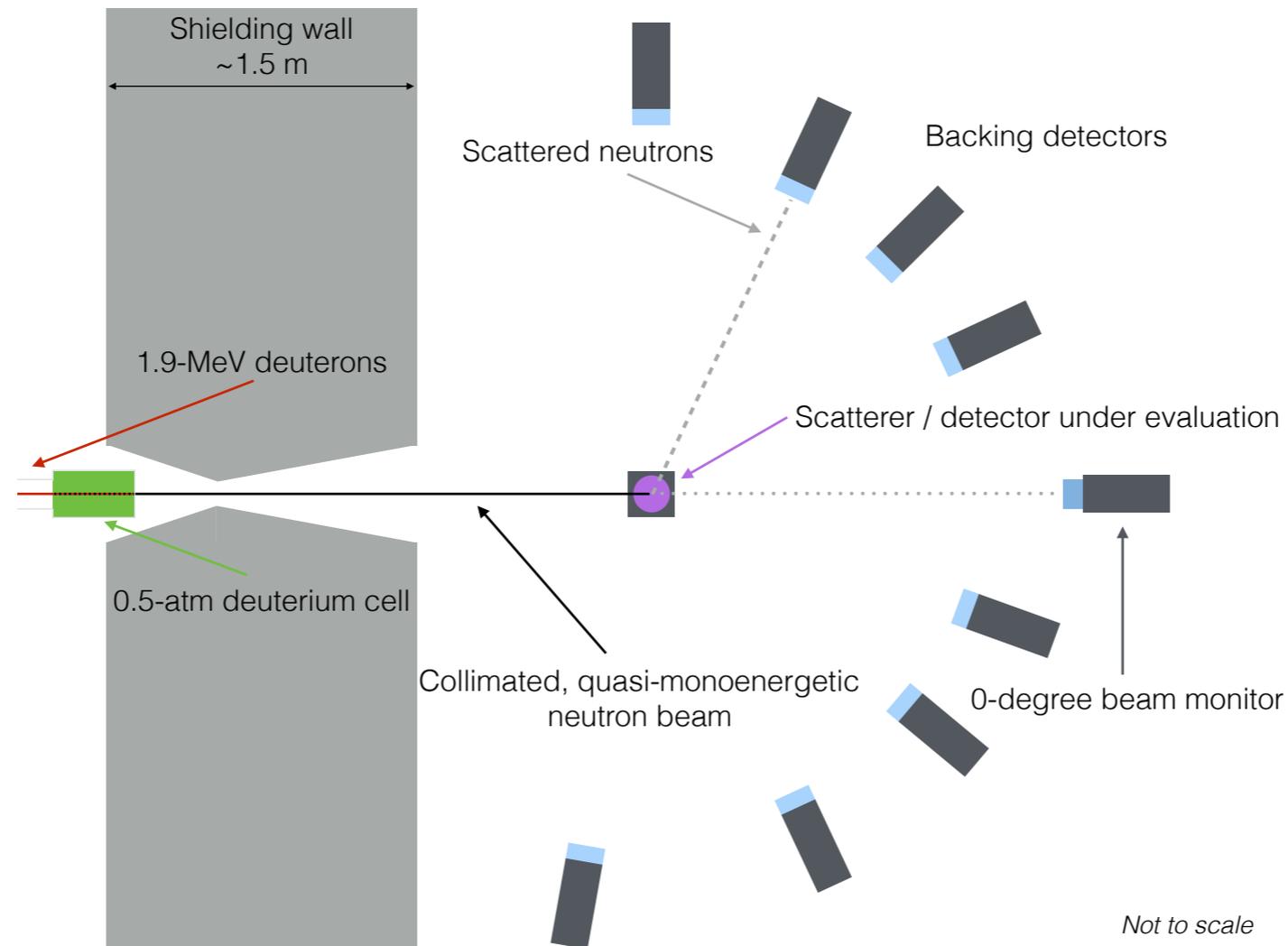
Low-energy nuclear recoils from CE ν NS

- Signature of CE ν NS in a detector is a low-energy nuclear recoil
- To properly interpret collected data, it is of paramount importance that detector response at these *nuclear recoil* energies be well understood
- Uncertainty in detector threshold translates into uncertainty in measured cross section
 - Situation worse for heavier targets



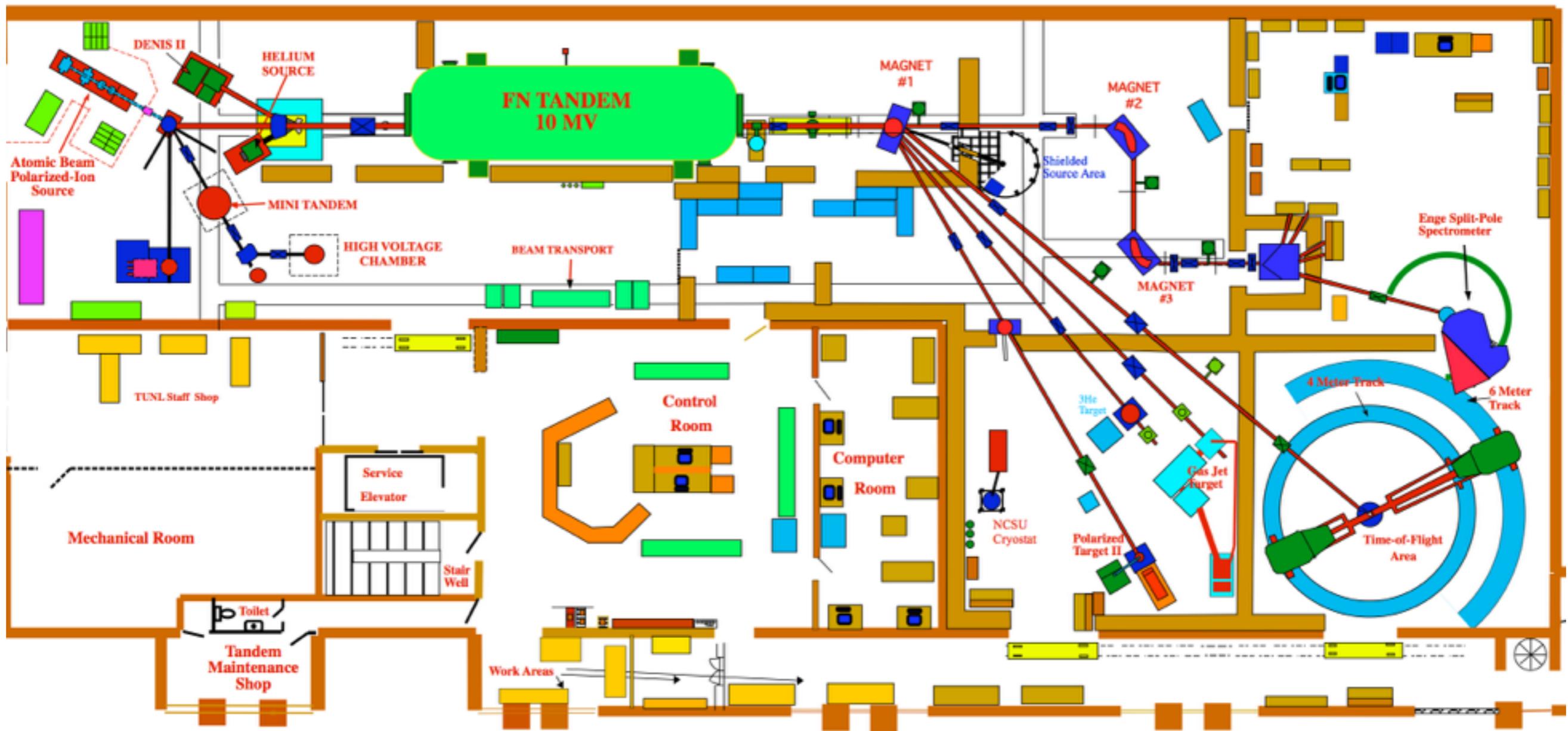
Low-energy nuclear recoils from neutron scattering

- Quasi-monoenergetic neutron beam scattered by central detector into fixed angles covered by “backing” detectors; nuclear recoil energy kinematically well defined



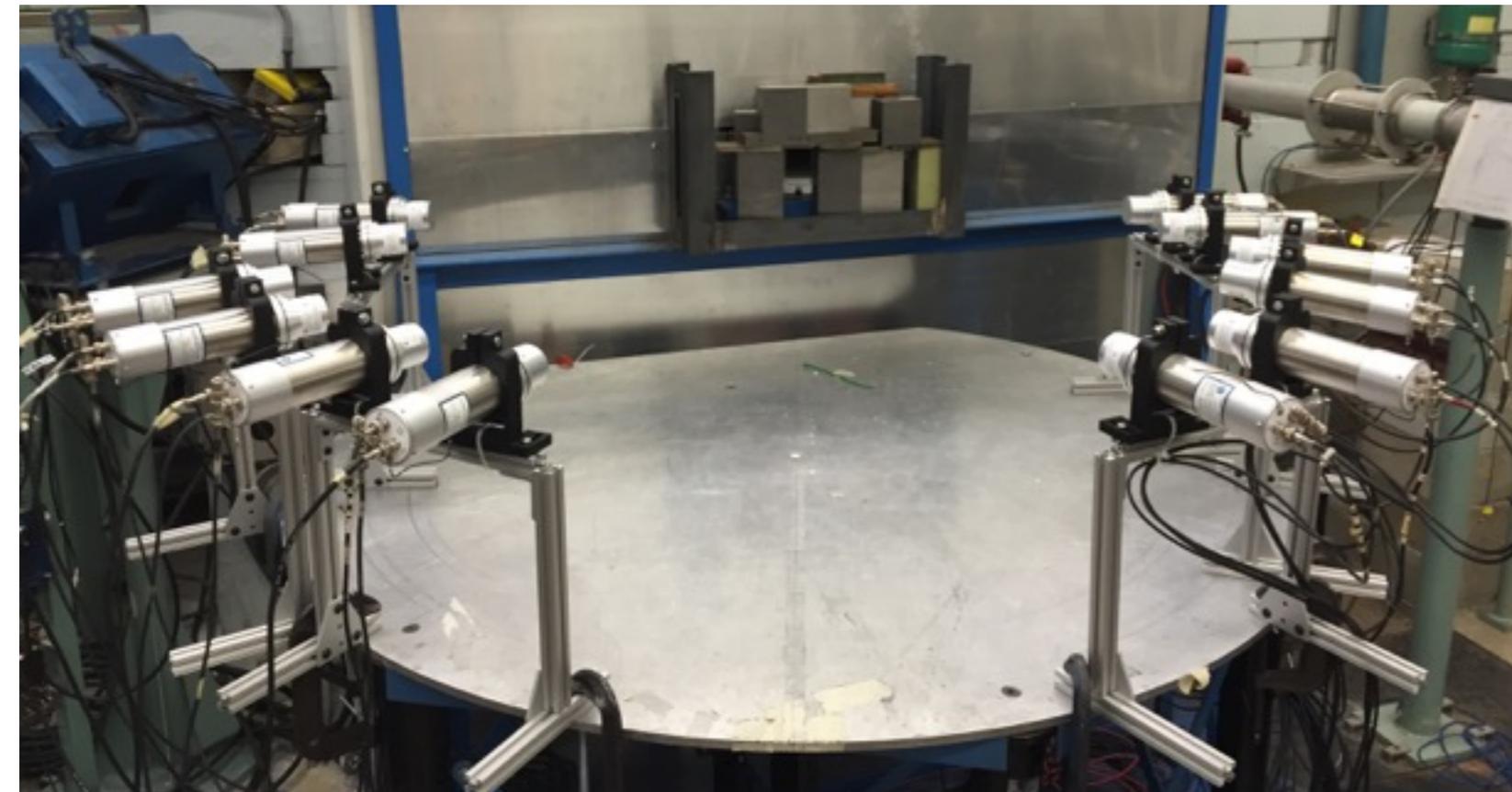
$$\Delta E = 2E_n \frac{M_n^2}{(M_n + M_T)^2} \left(\frac{M_T}{M_n} + \sin^2 \theta - (\cos \theta) \sqrt{\left(\frac{M_T}{M_n} \right)^2 - \sin^2 \theta} \right)$$

Tandem accelerator lab at TUNL



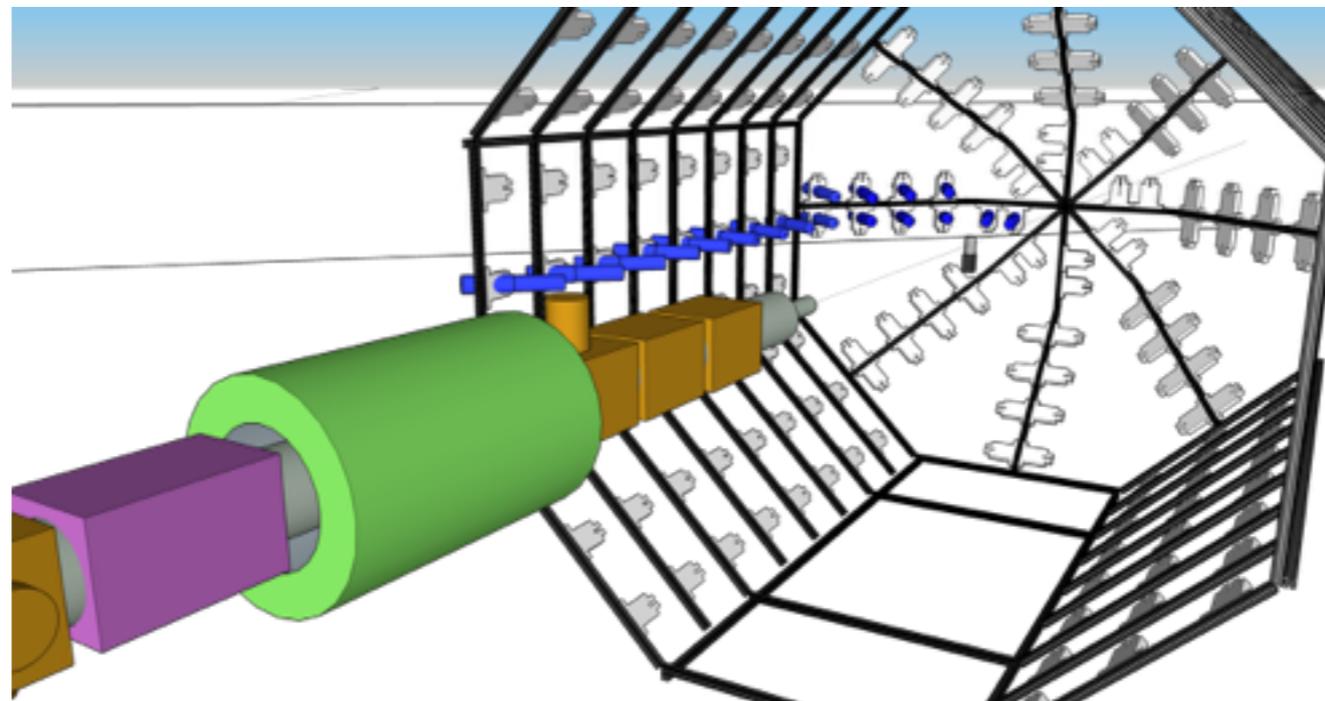
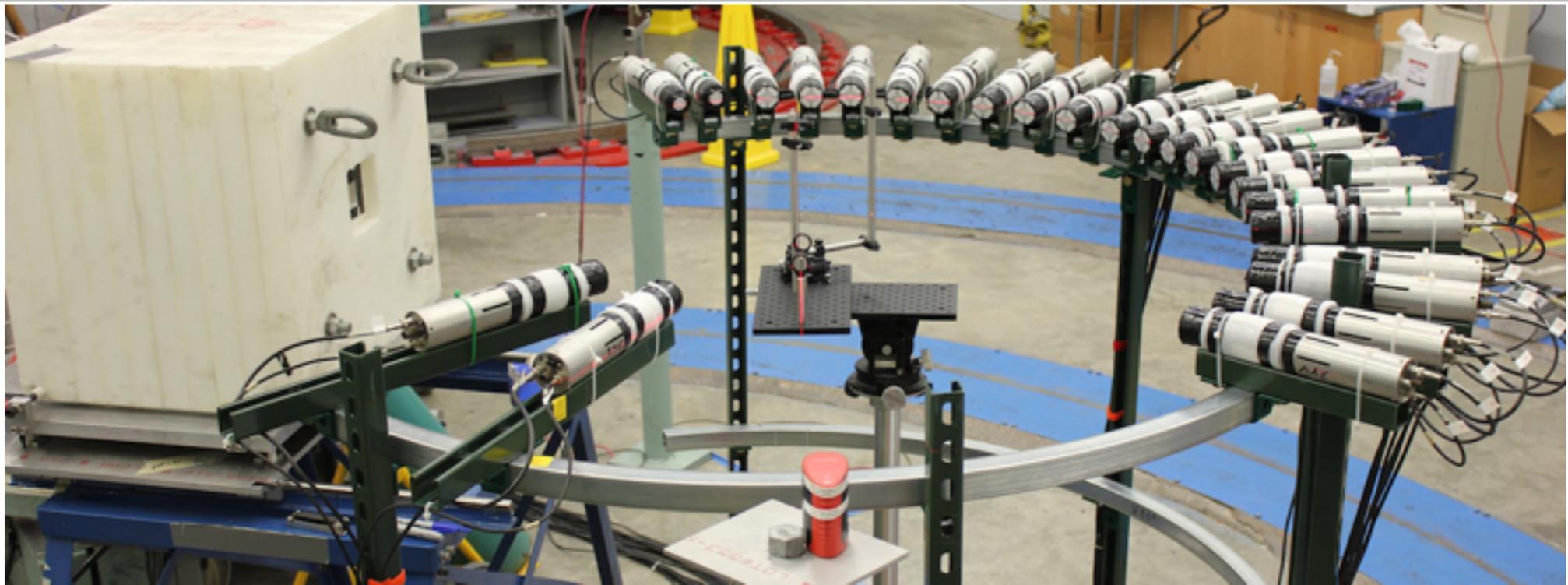
- 3 ion sources
- Beam can be bunched and chopped
- 10-MV maximum terminal voltage
- Numerous beam lines and experimental areas

Quenching factor measurements at TUNL



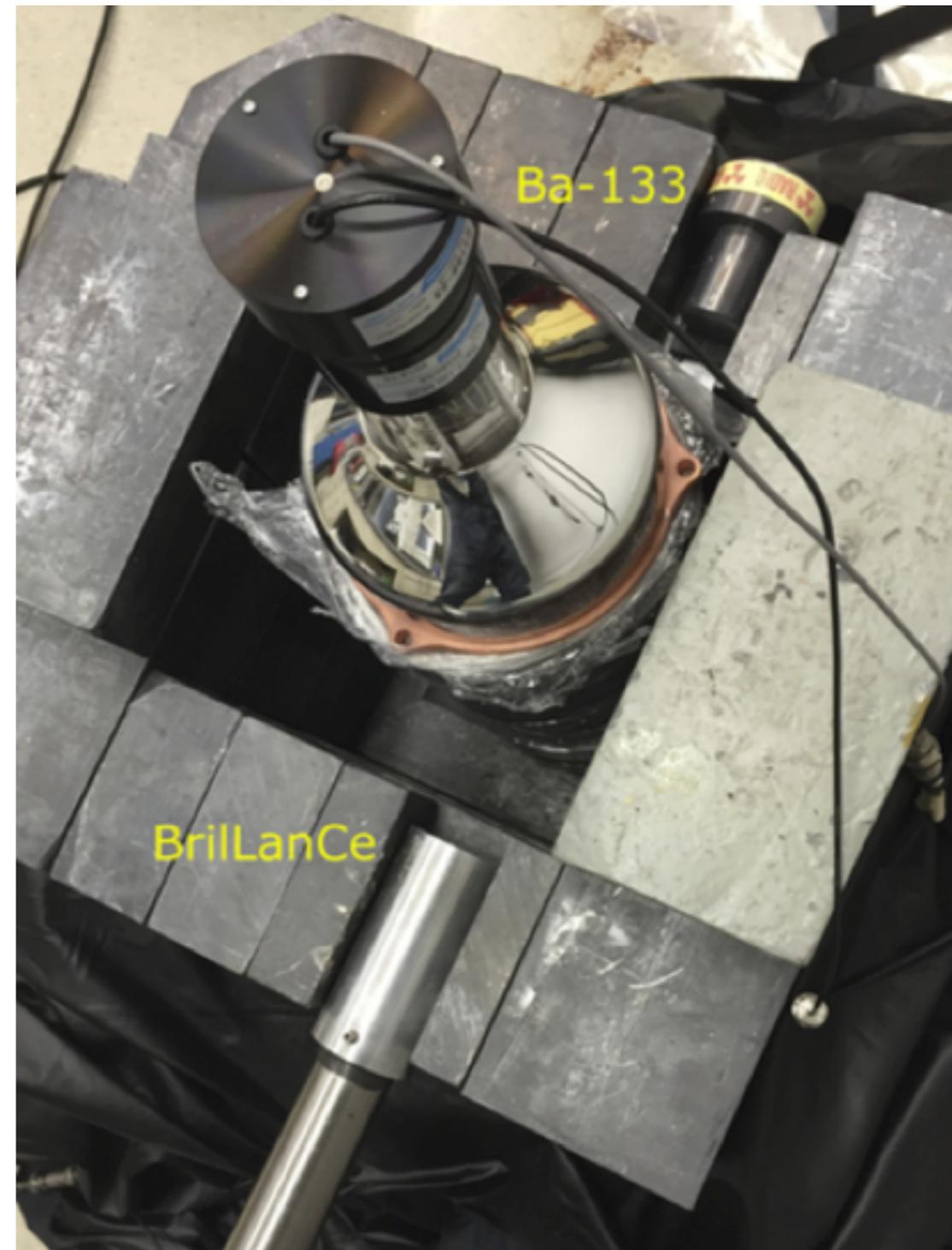
- Neutron beam produced by pulsed deuteron beam incident on deuterium gas cell
- Scattered neutrons detected by “backing detectors”
- Angle of backing detector selects well-defined nuclear recoil energy

Quenching factor measurements at TUNL

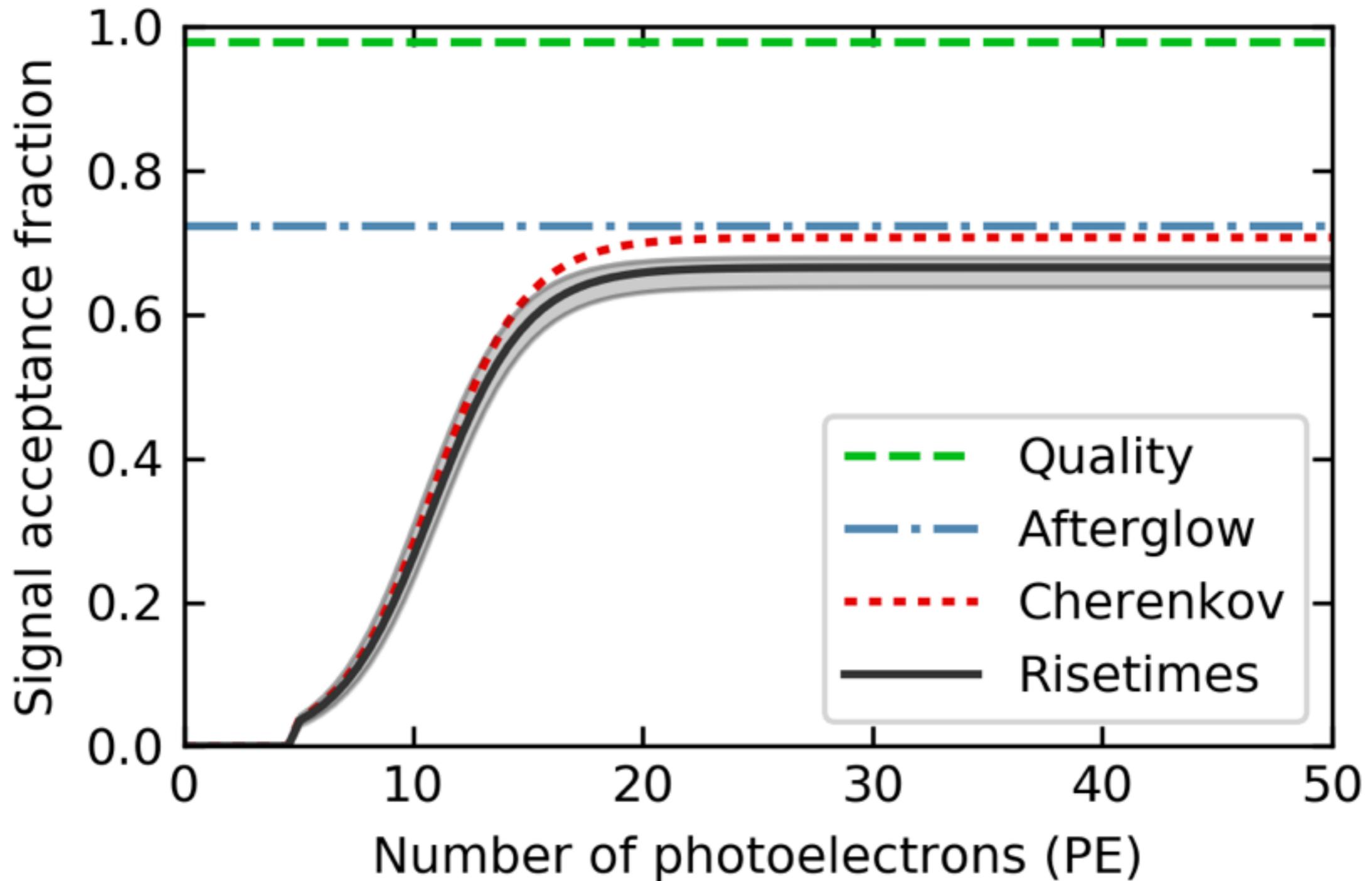


CE ν NS with CsI[Na]

- Prior to deployment, careful characterizations in Chicago
- Uniformity along length confirmed
- Response to low-energy gamma rays assessed via small-angle Compton scattering
- Allows tuning of cuts to reject spurious events but accept low-energy depositions in the CsI

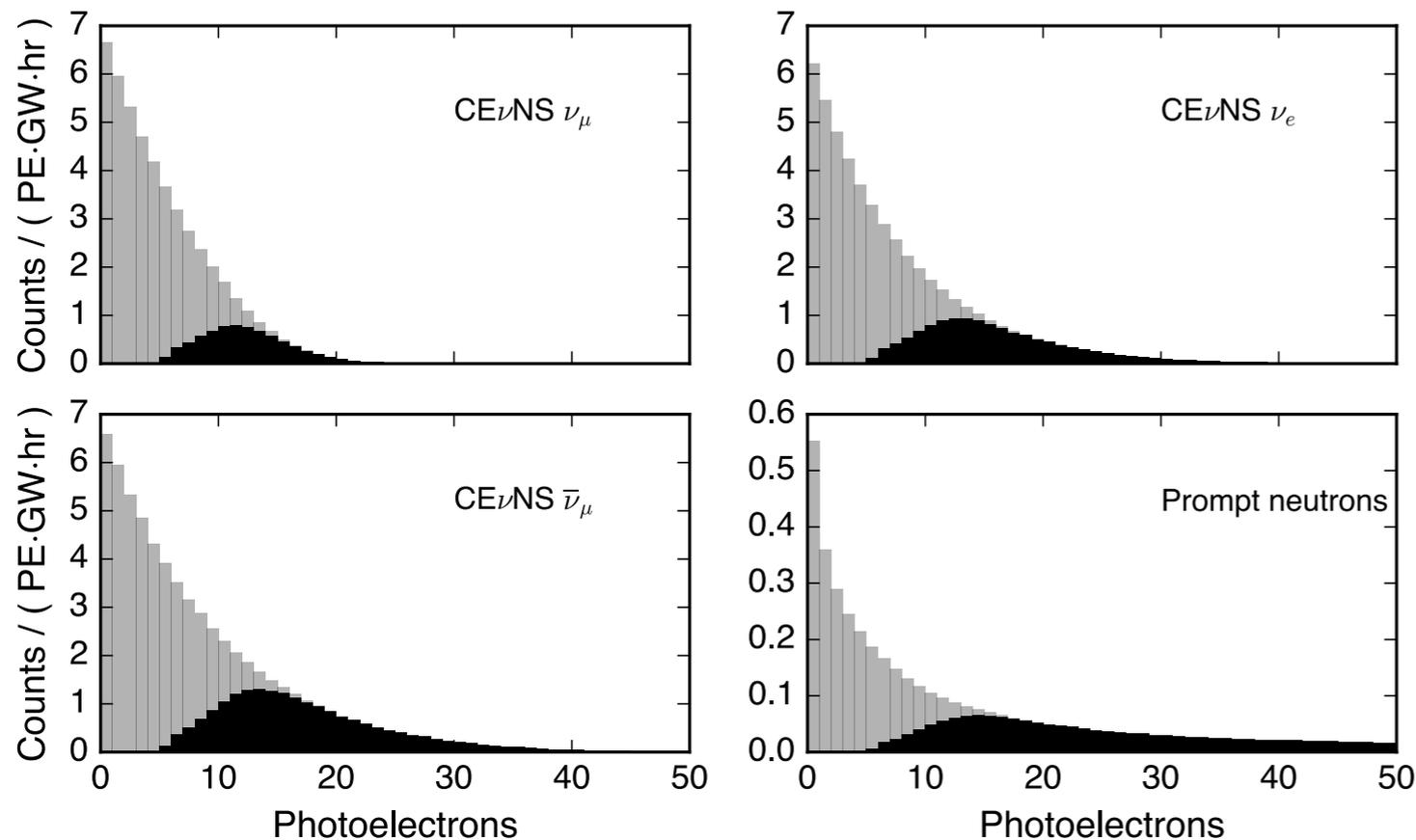


Analysis acceptance efficiency



Rate and shape estimates

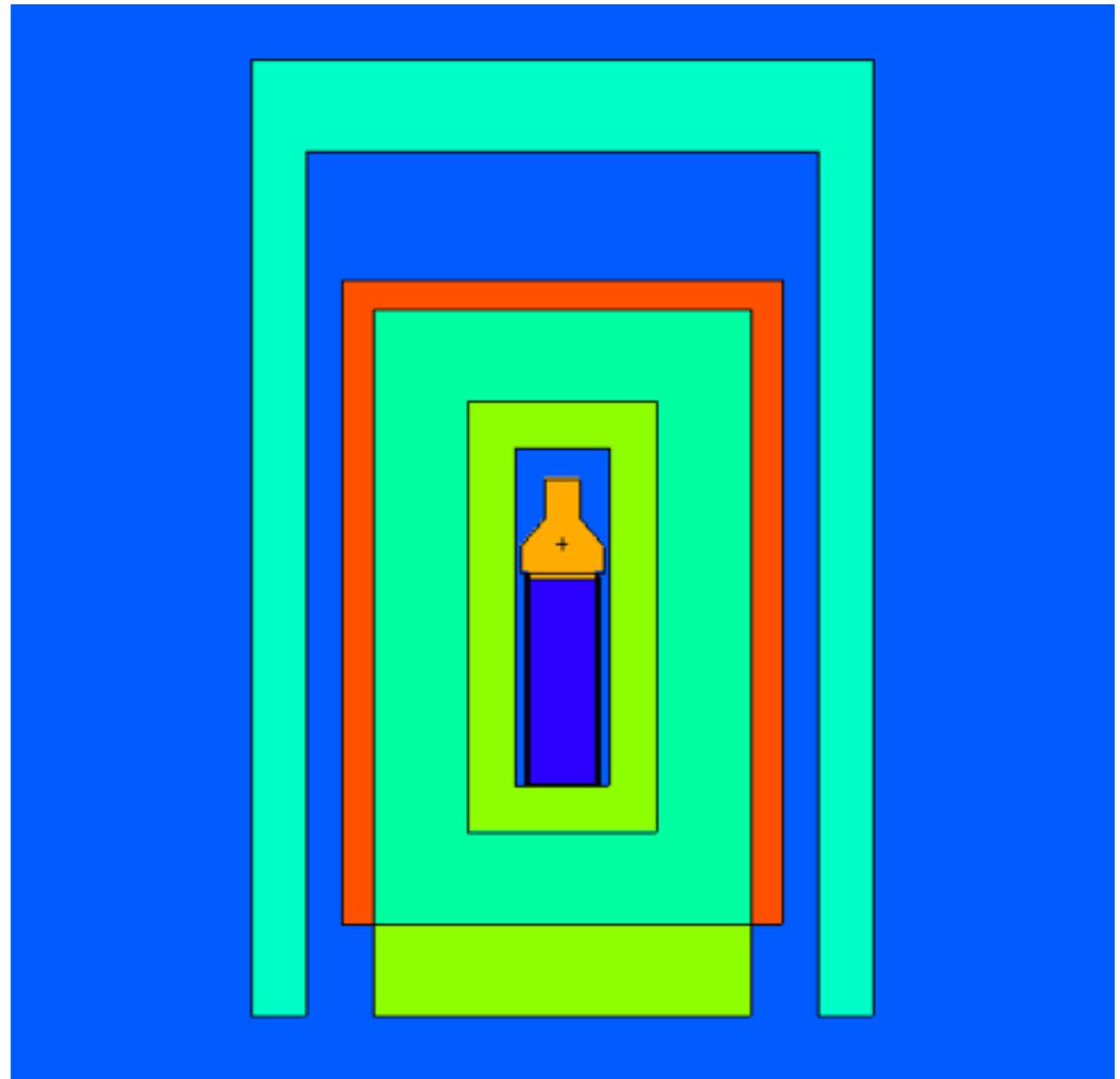
Raw CEvNS recoils Observed CEvNS recoils



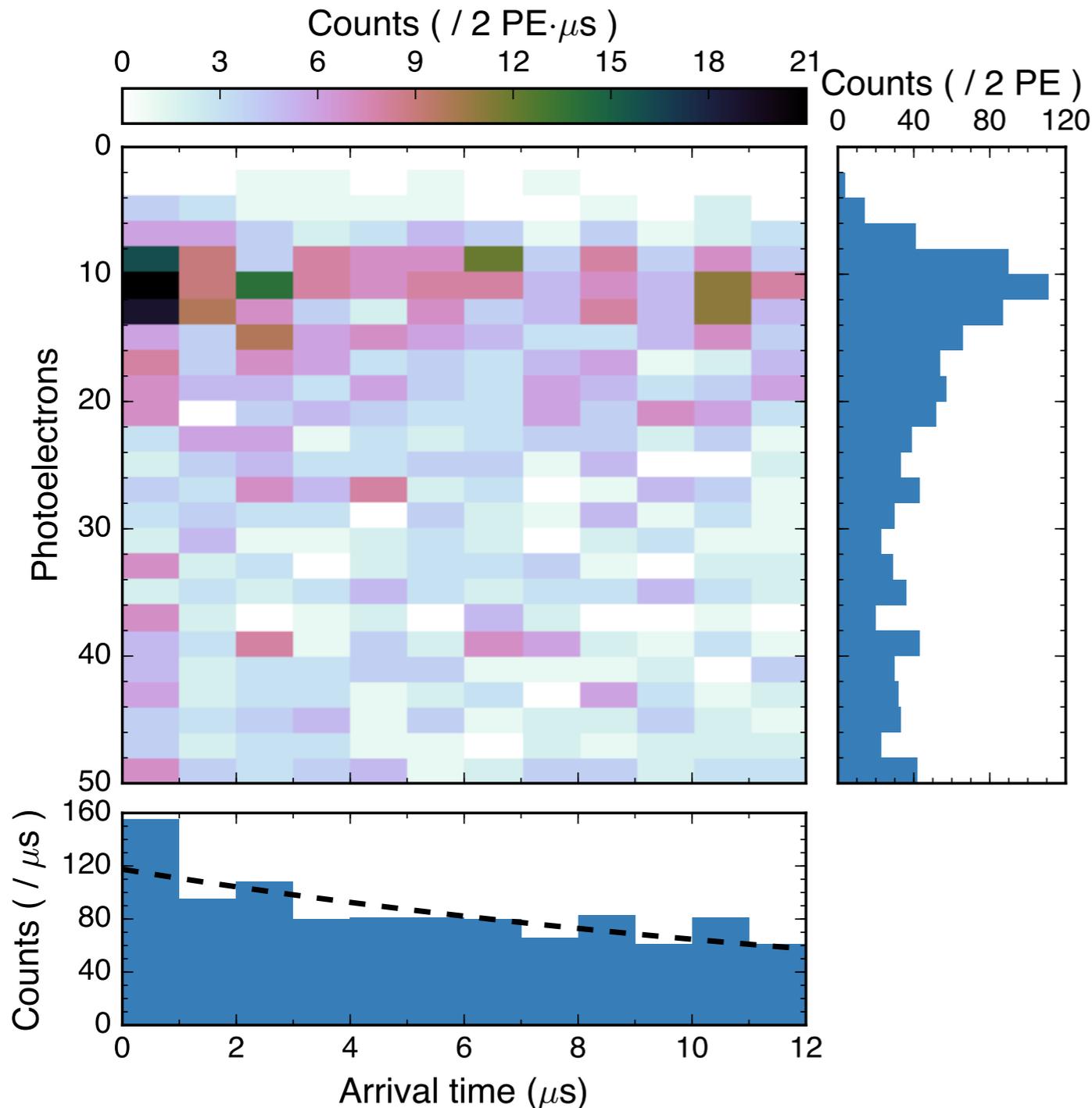
- Predict recoil distributions assuming SM - convert to photoelectrons using carefully determined calibrations
- *In situ* neutron measurements inform spectral model of prompt SNS neutrons
- Acceptance efficiency applied to models to produce beam-power-normalized PDFs in energy space

CE ν NS with CsI[Na]

- Several layers of shielding
 - 7.5-cm-thick inner HDPE layer (addressing NINs)
 - 5-cm low-activity lead
 - 10-cm contemporary lead
 - 5-cm plastic-scintillator muon veto
 - 9+ cm water shielding on sides and top

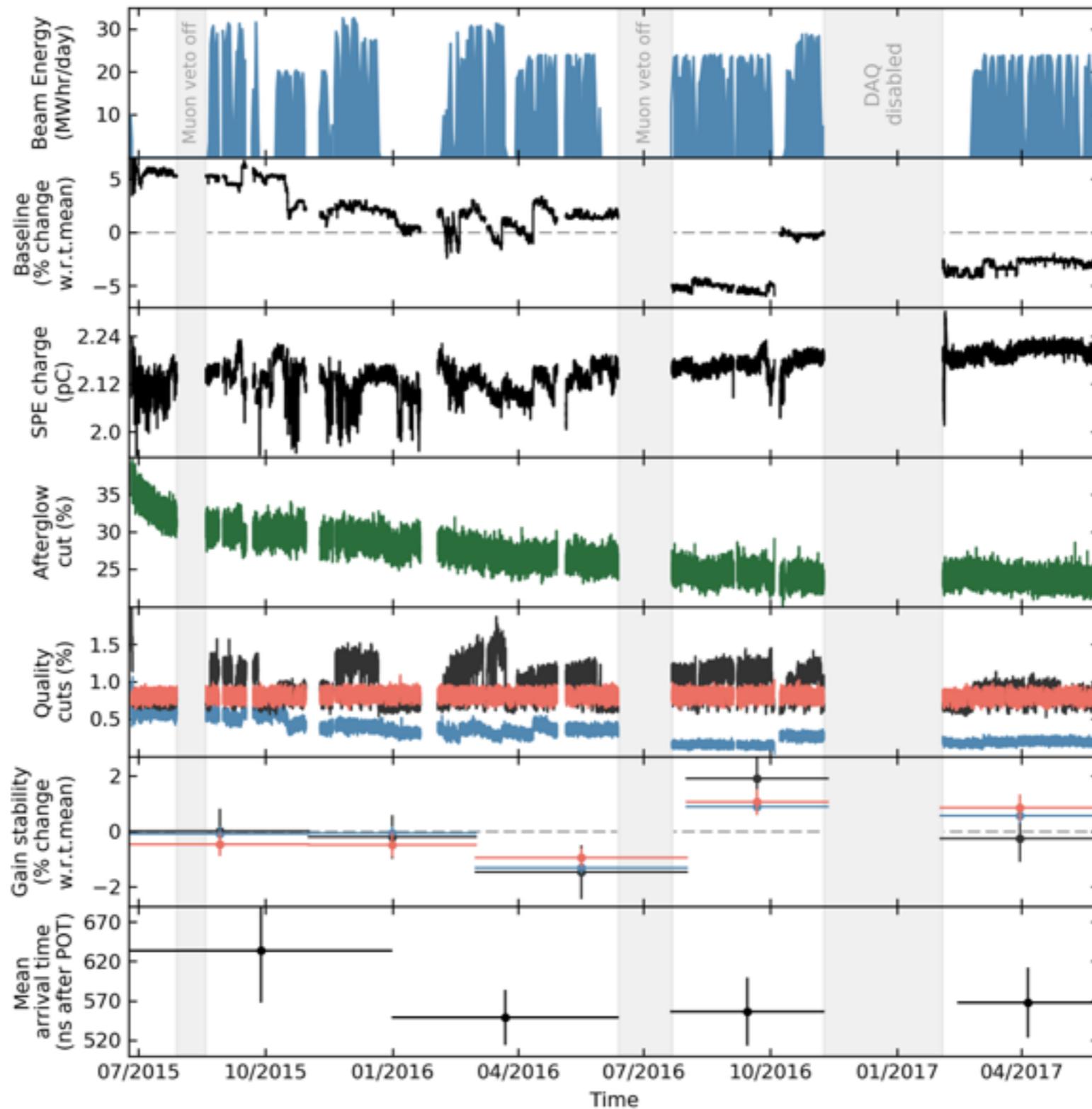


Background model for 2-D

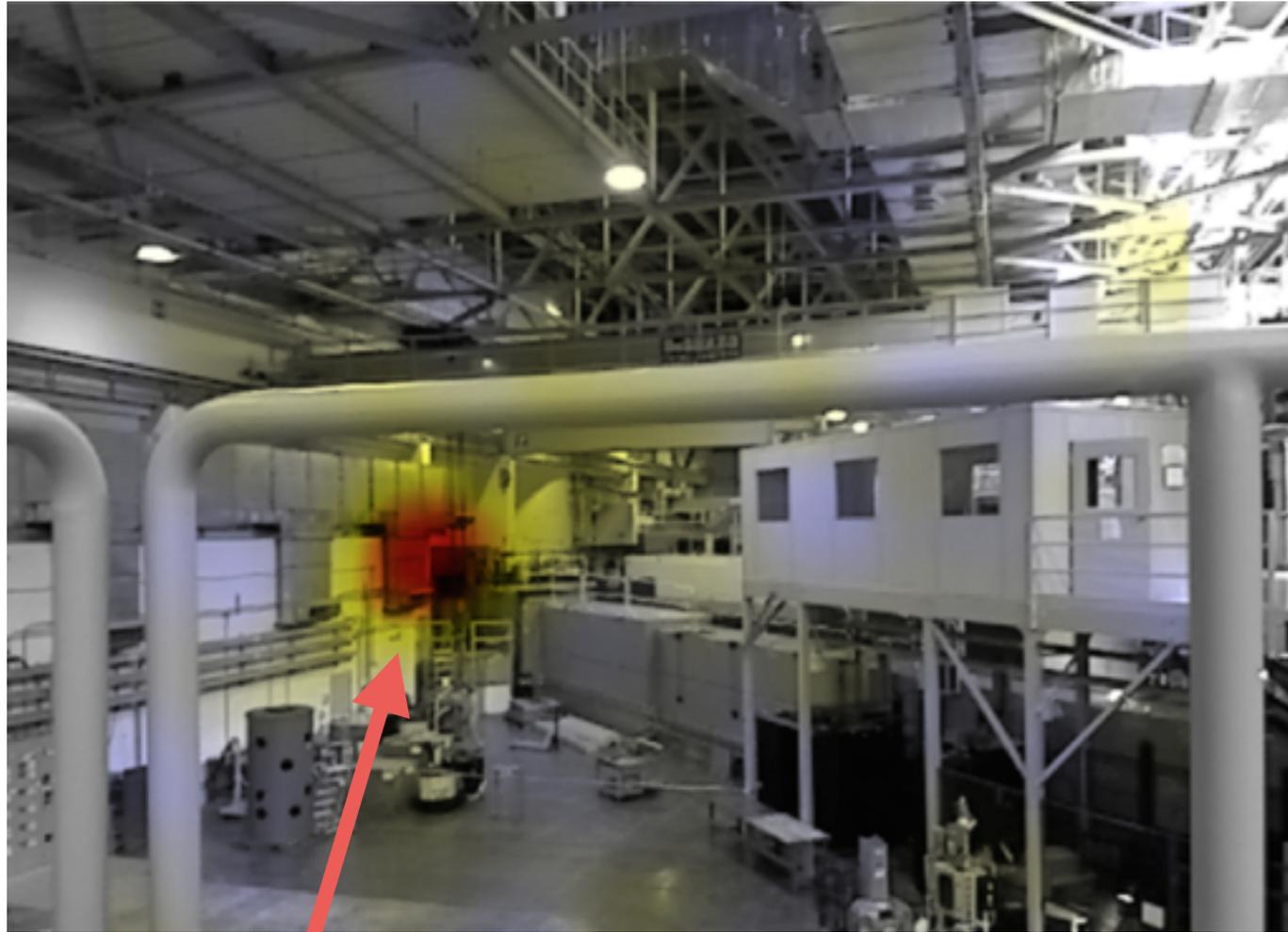


- Background model informed by anti-coincidence dataset
- Use “factorized” approach taking advantage of uncorrelated energy/time features
- Exponential fit to time projection, then used with energy projection to define model

Stability and general health checks



Neutrons at the SNS



Coded-aperture neutron imager

- Built by ORNL collaborators
- Intended for nuclear security applications
- Takes a picture of target area “in neutrons”

In case you forgot: SNS is a billion-plus dollar facility dedicated to neutrons

Target is “visible” through monolith shielding on the instrument floor

