## CEvNS: the shape of things to come



J.I. Collar

**APS Pittsburgh 2017** 

# A 10c introduction to coherent v-N scattering (CEvNS)

- <u>Uncontroversial</u> Standard Model process
- Large enhancement to cross-section  $P_{\text{Phys. Rev. D 9 (1974) 1389}}^{\text{D.Z. Freedman}}$ for E<sub>V</sub> < few tens of MeV ( $\sigma \propto N^2$ , possible only for neutral current)
- 43 yrs until successful detection... combination of source & detector technology was missing:

Detector mass must be at least ~1 kg (reactor experiment) + <u>recoil</u> energy threshold << 1keV

(recoils lose just 10-20% of E to ionization or scintillation)

Cryogenic bolometers and many other methods proposed over the last four decades.



#### Fundamental physics:

- Largest σ<sub>V</sub> in SN dynamics: should be measured to validate models (J.R. Wilson, PRL 32 (74) 849)
- A large detector can measure total E and T of SN  $v_{\mu}, v_{\tau} \Rightarrow$  determination of v oscillation pattern and mass of v star (J.F.Beacom, W.M.Far & P.Vogel, PRD 66(02)033011)
- Coherent ♂ same for all known v...
  oscillations observed in a coherent detector
  ⇒ evidence for v<sub>sterile</sub> (A.Drukier & L.Stodolsky, PRD 30 (84) 2295)
- Sensitive probe of weak nuclear charge
  ⇒ test of radiative corrections due to new physics above weak scale (L.M.Krauss, PLB 269, 407)
- More sensitive to NSI and new neutral bosons
  than v factories. Also effective v charge ratio
  (J. Barranco et al., hep-ph/0508299, hep-ph-0512029)
- $\sigma$  critically depends on  $\mu_V$ : observation of SM prediction would increase sensitivity to  $\mu_V$
- by > an order of magnitude (A.C.Dodd et al, PLB 266 (91) 434)
- Sensitive probe of neutron dens. distribution

Smallish detectors... <u>"v technology"?</u>

- Monitoring of nuclear reactors against illicit operation
- or fuel diversion: present proposals using conventional
- 1-ton detectors reach only > ~3 GWt reactor power
- Geological prospection, planetary tomography...
  the list gets much wilder.

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**ONE IN EVERY HOME** 

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#### Everyone needs a hobby



"Tendons"







#### Other reactor enthusiasts:



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PHYSICAL REVIEW LETTERS

1 JULY 1985

#### **Bolometric Detection of Neutrinos**

Blas Cabrera, Lawrence M. Krauss, and Frank Wilczek Department of Physics, Stanford University, Stanford, California 94305 Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 01238 Institute for Theoretical Physics, University of California, Santa Barbara, California 93106 (Received 14 December 1984)

Elastic neutrino scattering off electrons in crystalline silicon at 1–10 mK results in measurable temperature changes in macroscopic amounts of material, even for low-energy (<0.41MeV)  $pp \nu$ 's from the sun. We propose new detectors for bolometric measurement of low-energy  $\nu$  interactions, including coherent nuclear elastic scattering. A new and more sensitive search for oscillations of reactor antineutrinos is practical (~100 kg of Si), and would lay the groundwork for a more ambitious measurement of the spectrum of pp, <sup>7</sup>Be, and <sup>8</sup>B solar  $\nu$ 's, and supernovae anywhere in our galaxy (~10 tons of Si).







#### CONNIE 2014-2015 sensor:





Technological applications: reactor monitoring is around the corner (compact microbolometer arrays and CCDs are almost there)

#### Fortunately, reactors are not the only game in town:



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## Enter COHERENT @ SNS:

### How to Make an Unambiguous Measurement

- Observe the pulsed v time-structure
- Observe the 2.2  $\mu s$  characteristic decay of muon decay v's
- Observe the N<sup>2</sup> cross section behavior between targets



P-Type Point Contact HPGe



Low-Background Csl[Na]



Nal[TI]



Single Phase LAr

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#### Why CsI[Na]? (NIM A773 (2014) 56) scattered neutrino Large N<sup>2</sup> => large x-section. Cs and I surround Xe in Periodic Table: Ζ nuclear they behave much like a single recoiling boson recoil species, greatly simplifying understanding Quenching factor in energy ROI sufficient for ~5 keVnr threshold (we have measured this). recoils scintillation $10^{\circ}$ photoelectrons (bialkali PMT) $10^{30}$ 10 20 $10^{10}$ above threshold / kg CsI / yı recoils / keV / kg CsI / year 10 1010 $10 \ 20 \ 30 \ 40$ 50energy threshold (keVnr) $10^{0}$ $10^{-10}$ total 20 m from SNS target $10^{-2}$

15

10

5

20

recoil energy (keVnr)

25

30

35

40

 Statistical NR/ER discrimination is possible at low-E with large statistics.

the NR response.

- Sufficiently low in intrinsic backgrounds (U, Th ,K-4Ó, Rb-87, Cs-134,137) Measurements in complete SNS shield and 6 m.w.e. indicated we were ready)
- Practical advantages: High light yield (64 ph/keVee), optimal match to bialkali PMTs, rugged, room temperature, inexpensive (\$1/g), modest afterglow (CsI[Tl] not a viable option for surface experiment).
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#### Preliminaries: background studies w/ 2 kg prototype



#### Preliminaries: in situ neutron bckg measurements



#### Preliminaries: 14.5 kg detector characterization





#### Preliminaries: Quenching factor measurements



### Installation of 14.5 kg CsI[Na] June 2015 (first "handheld" v detector)



#### What next: precision CEvNS studies with PPCs

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- Best-measured quenching factor (and getting better).
- 10 kg of Ge in compact 4-crystal array (possible through CoGeNT/C-4 R&D).







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CEvNS holds great promise to expand our knowledge of neutrino properties.

However, we must expand our knowledge of detector response to low-E NRs first!