

**CNEC**Consortium for  
Nonproliferation  
Enabling Capabilities**KICP****NNSA**  
National Nuclear Security AdministrationU.S. DEPARTMENT OF  
**ENERGY**

COHERENT  
SNS  
 $\nu$  Scattering

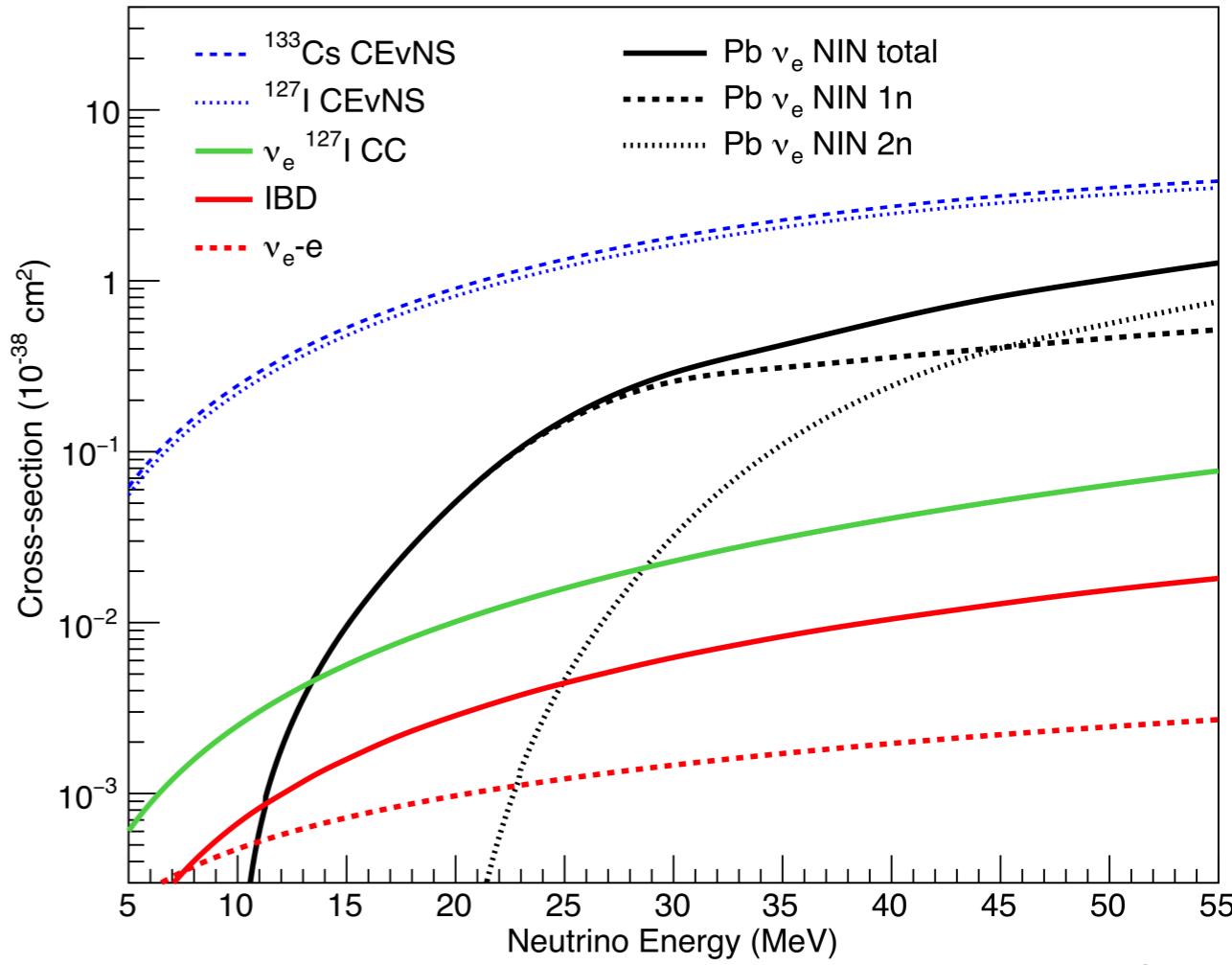
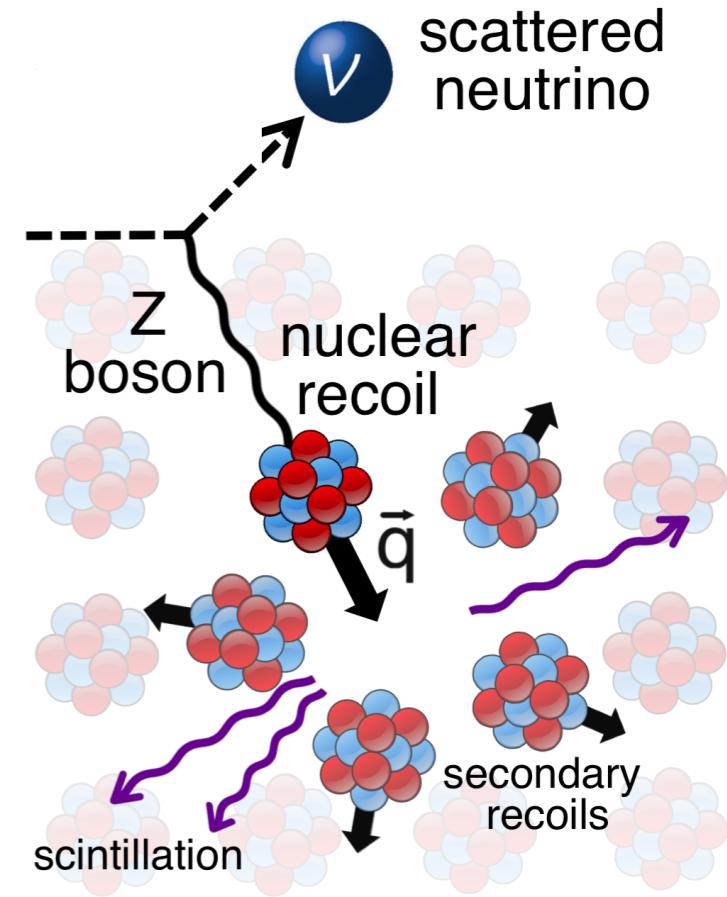
Phil Barbeau, Duke University

# Coherent $\nu$ -Nucleus Scattering

- 43 years ago, Coherent Elastic Neutrino-Nucleus Scattering (CEvNS) was predicted with the realization of the neutral weak current.

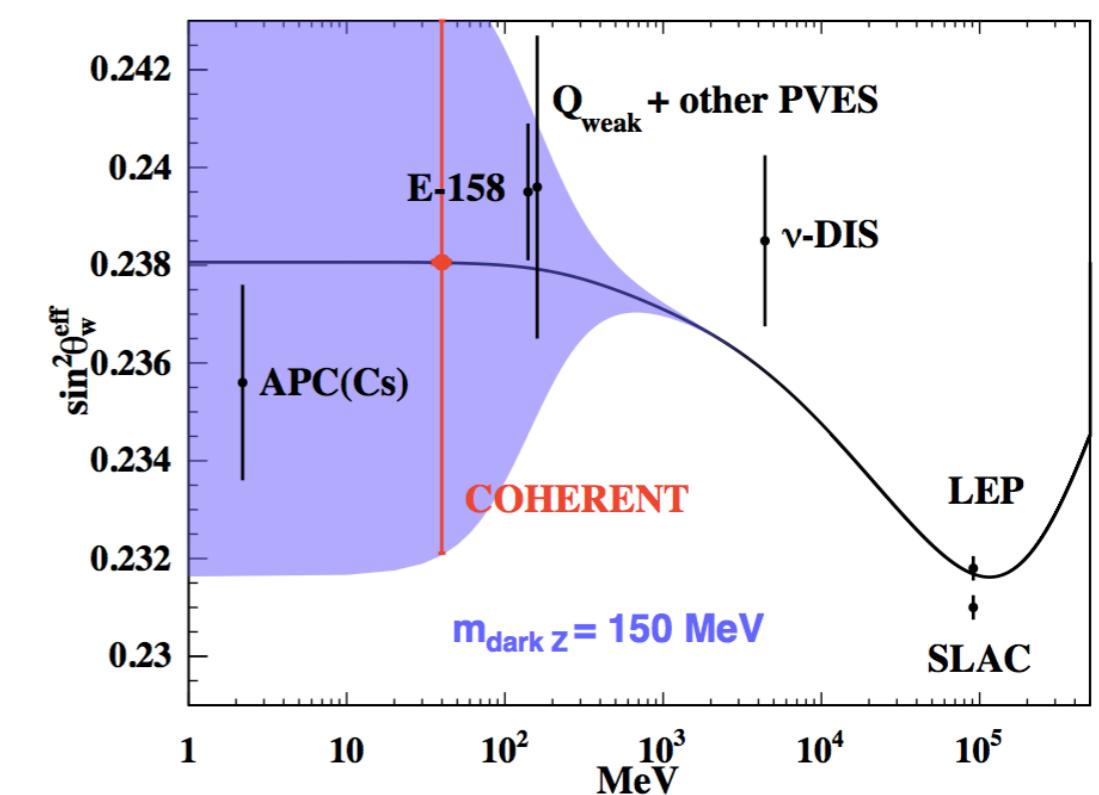
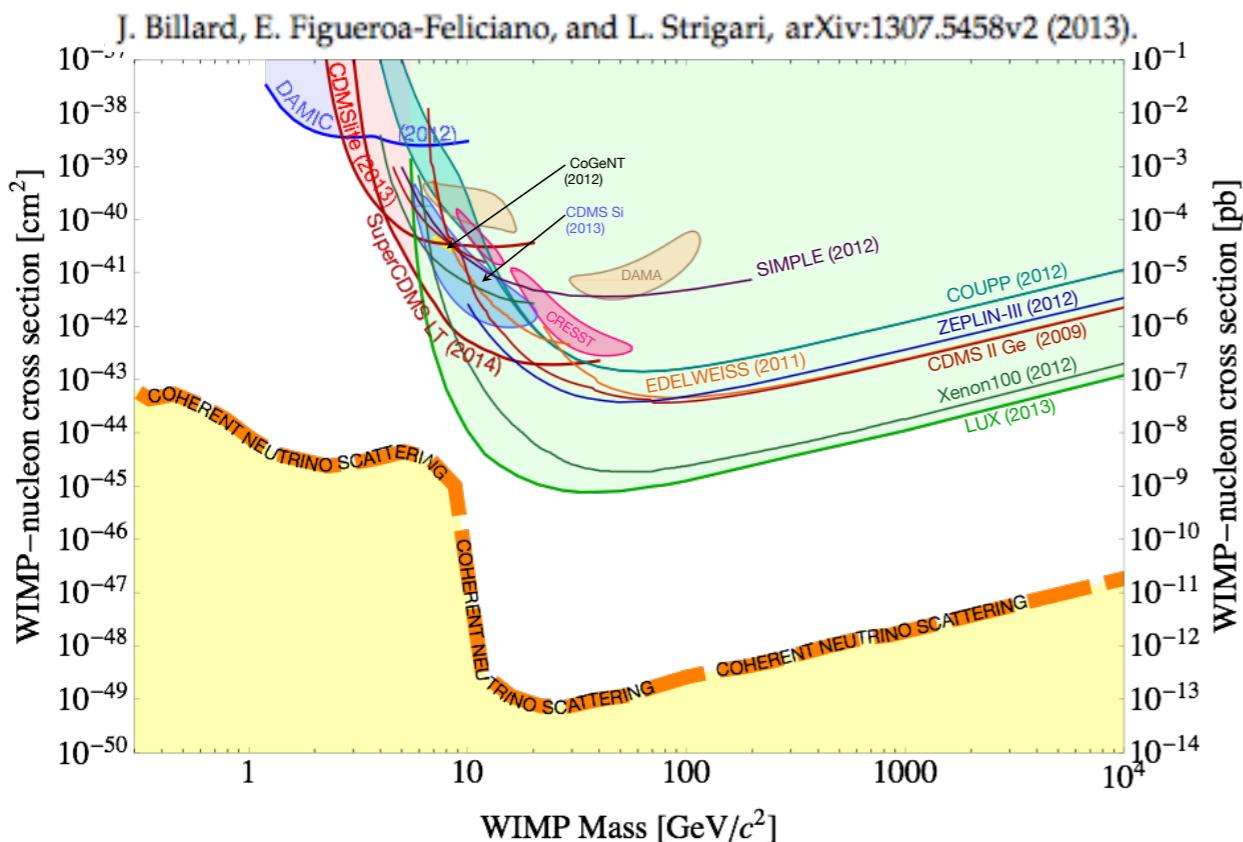
D. Z. Freedman, PRD 9 (5) 1974

- Neutrino scatters coherently off all Nucleons  $\rightarrow$  cross section enhancement:  $\sigma \propto N^2$
- Initial and final states must be identical: Neutral Current elastic scattering
- Nucleons must recoil in phase  $\rightarrow$  low momentum transfer  $qR < 1 \rightarrow$  very low energy nuclear recoil



# Coherent $\nu$ -Nucleus Scattering: a new tool

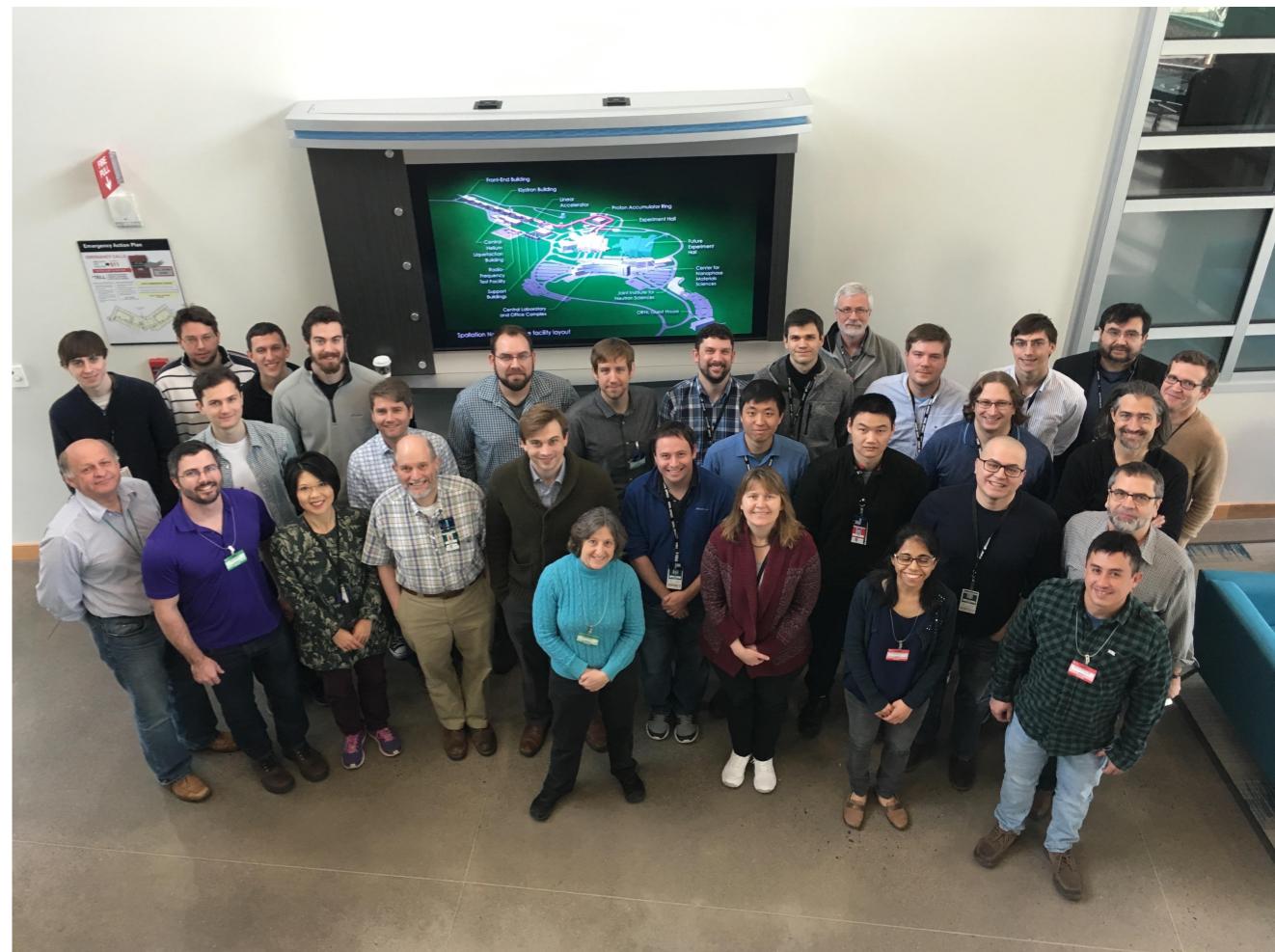
- Largest  $\sigma$  in Supernovae dynamics.  
**J.R. Wilson, PRL 32 (74) 849**
- Non-standard Neutrino Interactions relevant for DUNE & LBL CP violation.  
**K. Scholberg, Phys.Rev.D73:033005,2006**  
**J. Barranco et al., JHEP0512:021,2005**  
**Mehedi Masud, Poonam, Mehta, arXiv: 1603.01380**
- CEvNS is an irreducible background from WIMP searches.  
**J. Billard, E. Figueroa-Feliciano, and L. Strigari, arXiv:1307.5458v2 (2013).**
- Sensitive tool for Sterile neutrino searches  
**A. J. Anderson et al., PRD 86 013004 (2012)**  
**A. Drukier & L. Stodolsky, PRD 30 (84) 2295**
- A precision test of  $\sigma$  is a sensitive test of new physics above the weak scale.  
**L. M. Krauss, PLB 269, 407**
- Neutrino Magnetic Moments  
**A. C. Dodd, et al., PLB 266 (91), 434**
- Neutron distribution functions  
**K. Patton, et al., PRC 86, 024216**



# The ((C)OHERENT Collaboration



Laurentian University  
Université Laurentienne



Consortium for  
Nonproliferation  
Enabling Capabilities

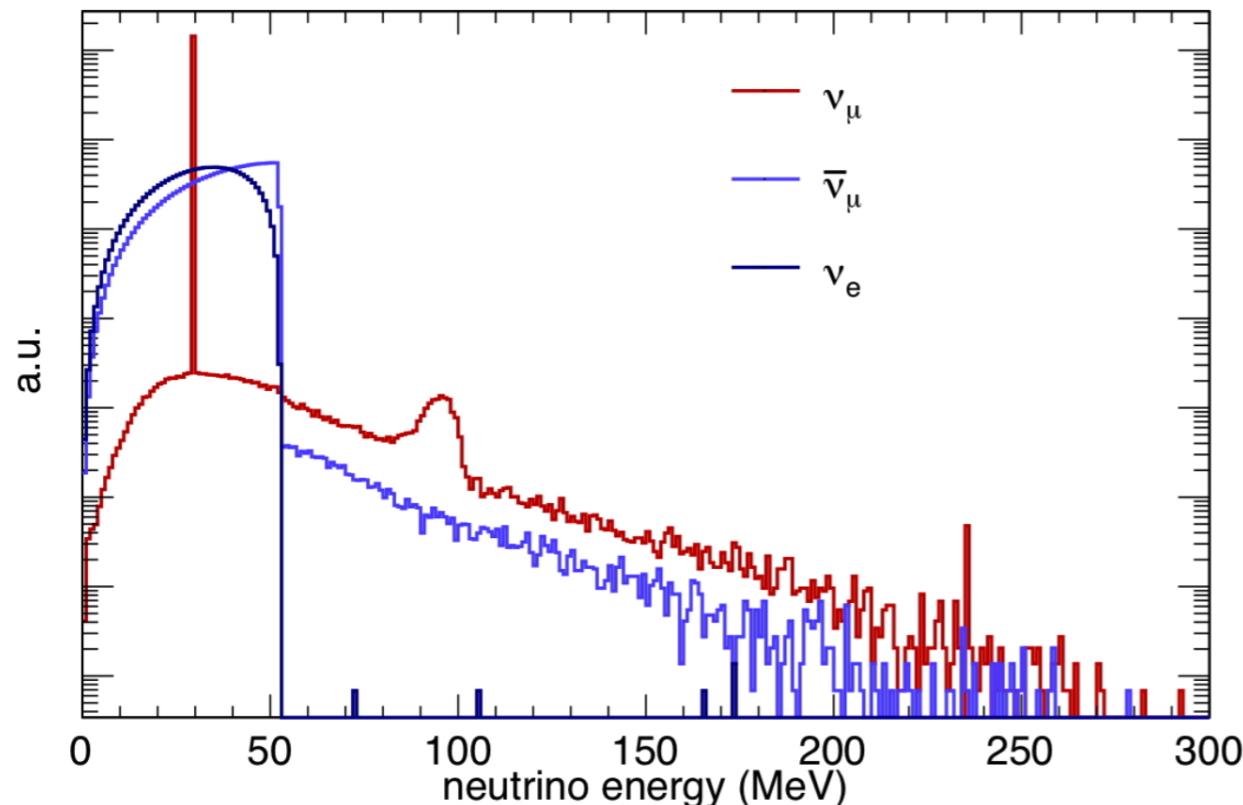


# The Spallation Neutron Source

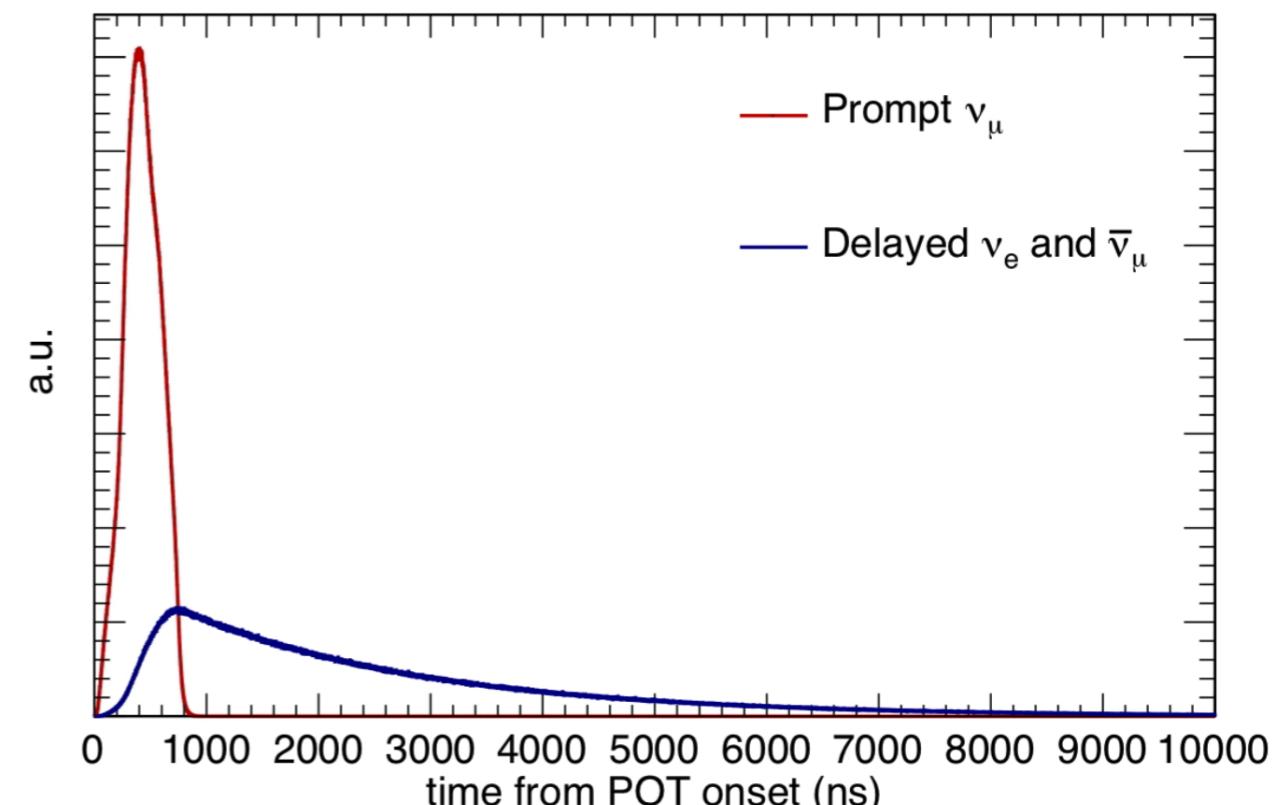
- Pion Decay-at-Rest Neutrino Source
- $\nu$  flux  $4.3 \times 10^7 \nu \text{ cm}^{-2} \text{ s}^{-1}$  at 20 m
- Pulsed: 800 ns full-width at 60 Hz



**<1% contamination from non-CEvNS scatters**

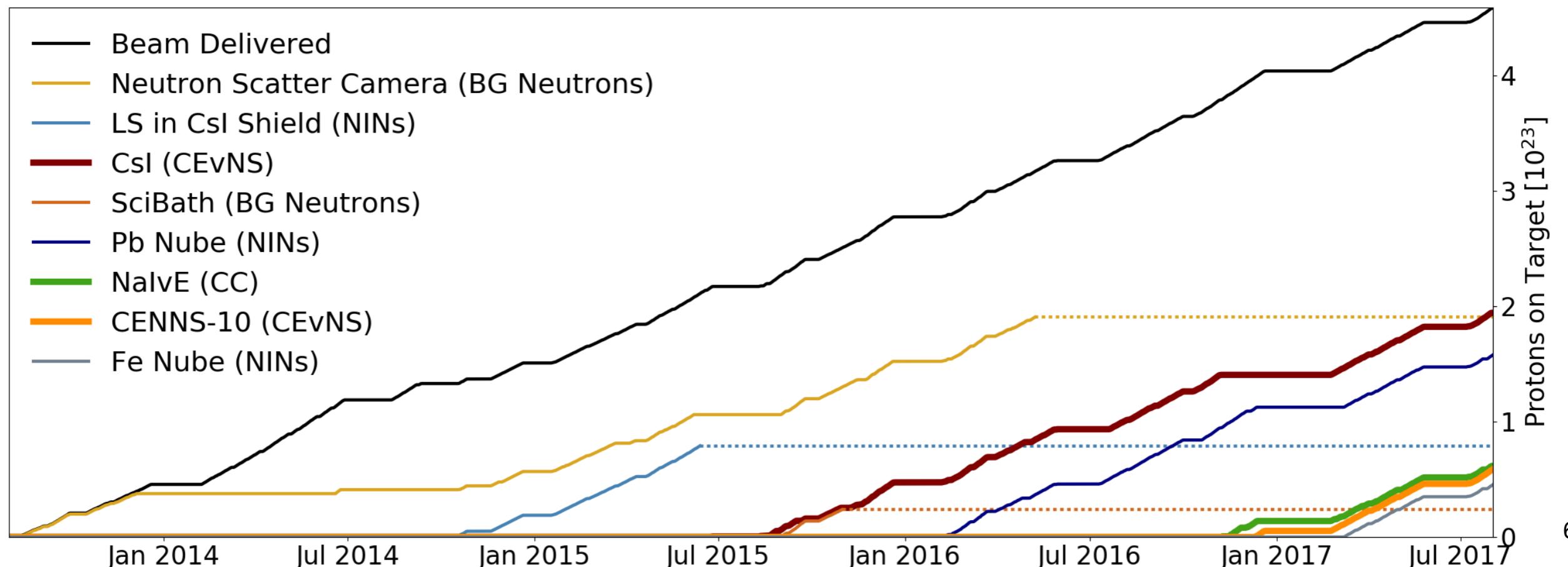
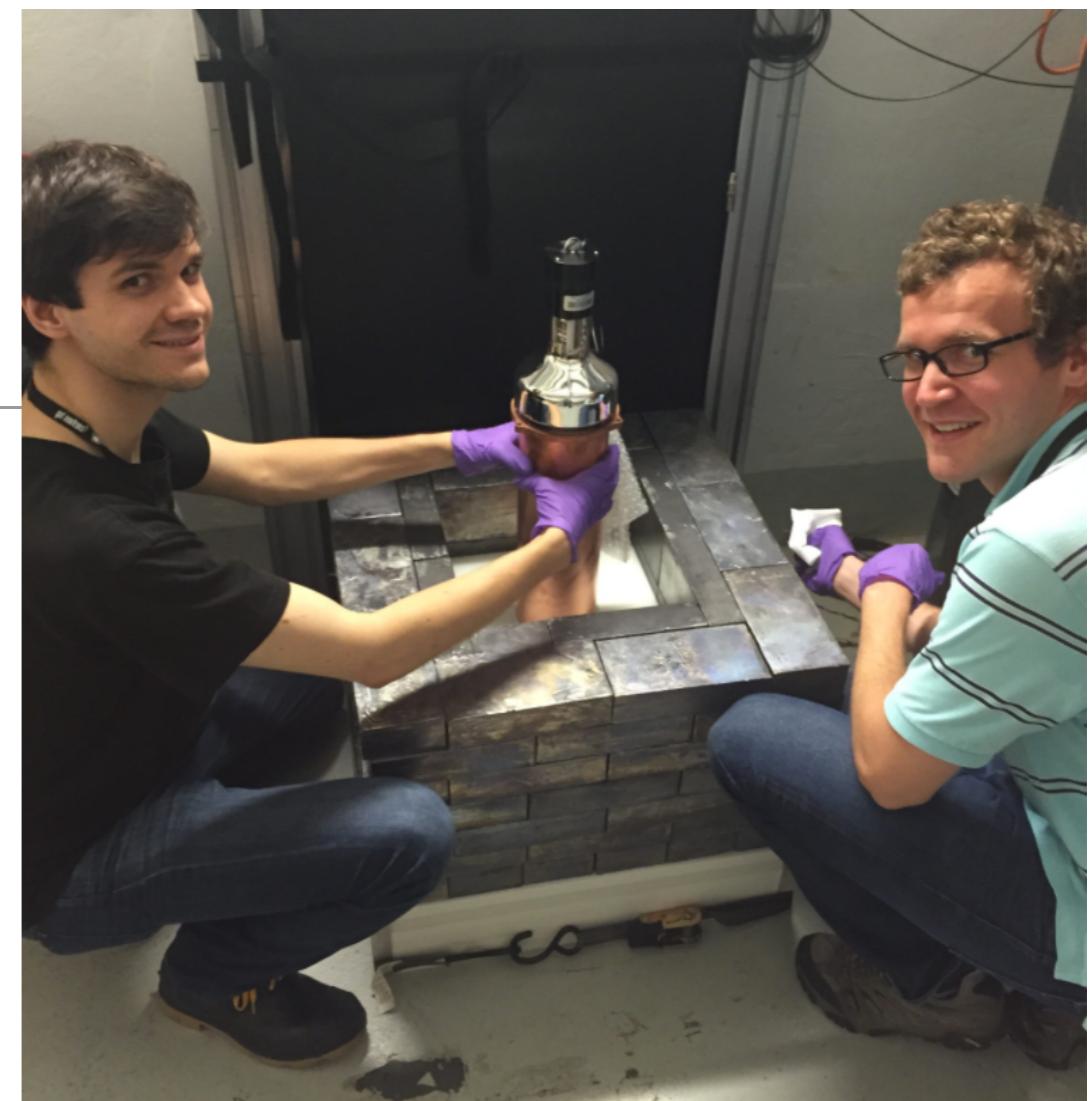


**$\sim 4 \times 10^{-5}$  background reduction**



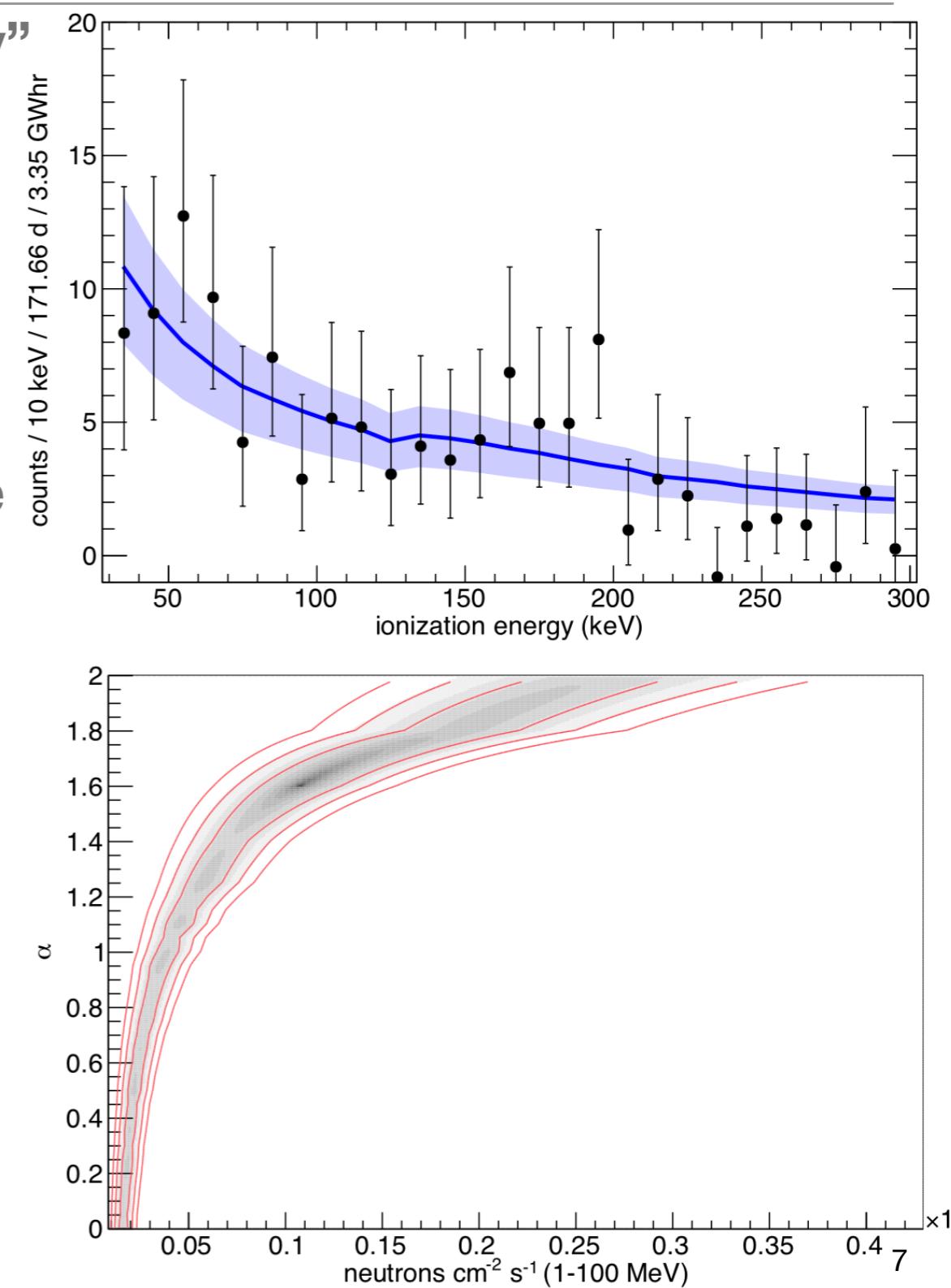
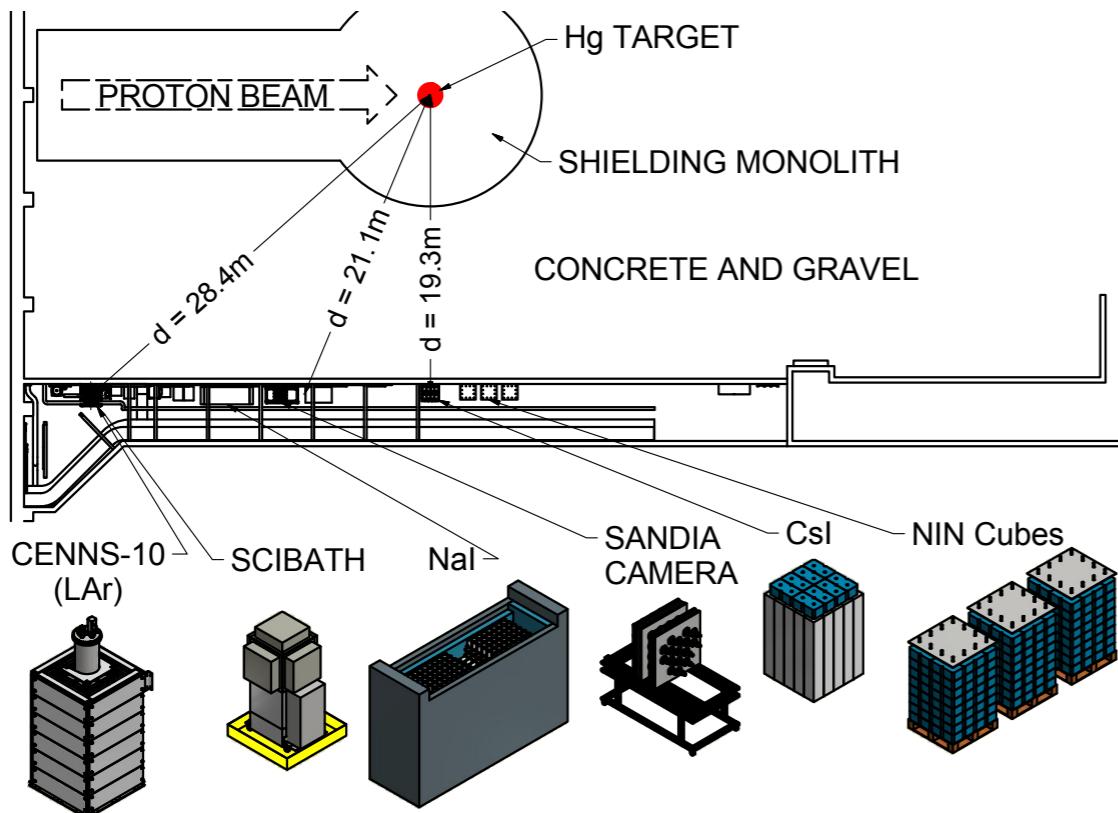
# A hand-held neutrino detector

- 14.6 kg low-background CsI[Na] detector deployed to a basement location of the SNS in the summer of 2015
- $\sim 2 \times 10^{23}$  POT delivered and recorded since CsI began taking data



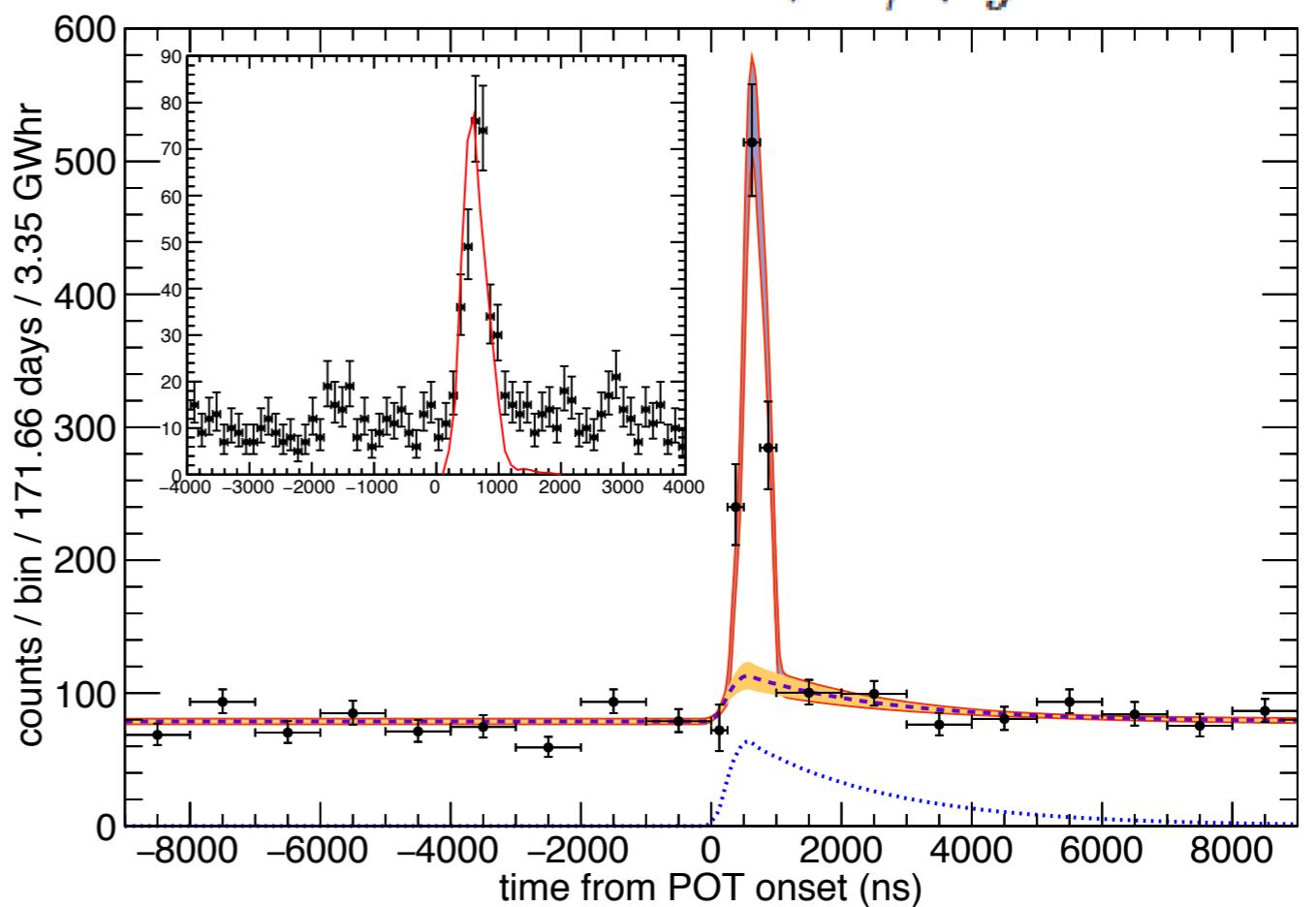
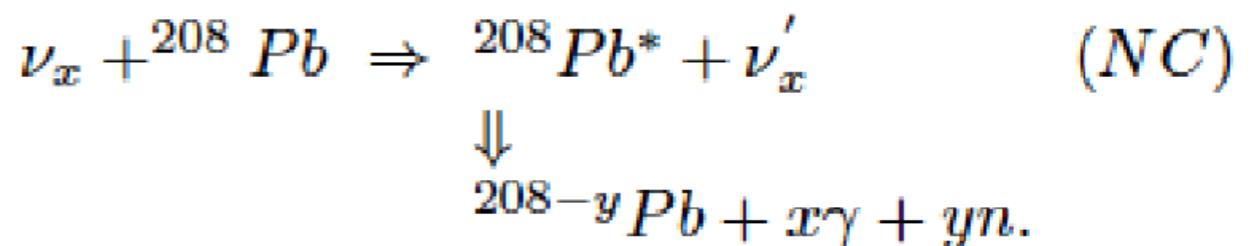
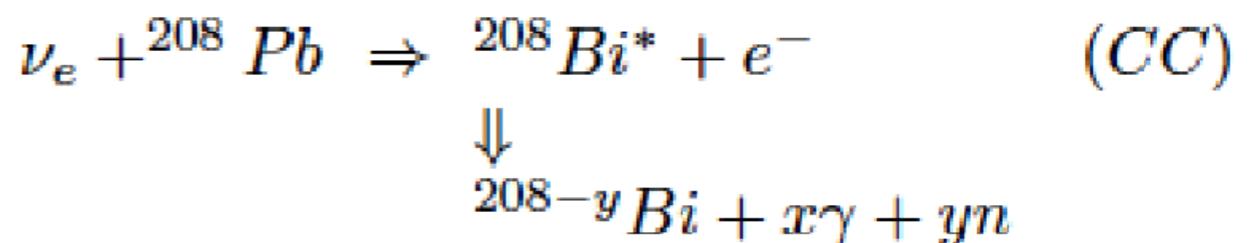
# What about neutron backgrounds?

- A basement location dubbed “**Neutrino Alley**” shields us from the copious flux of neutrons produced at the SNS
- Location identified after a detailed hunt with neutron detectors (Scatter-Camera and SciBath)
- In situ measurement of the neutrons within the CsI shield prior to installation of CsI[Na] detector



# Another background: neutrino-induced neutrons

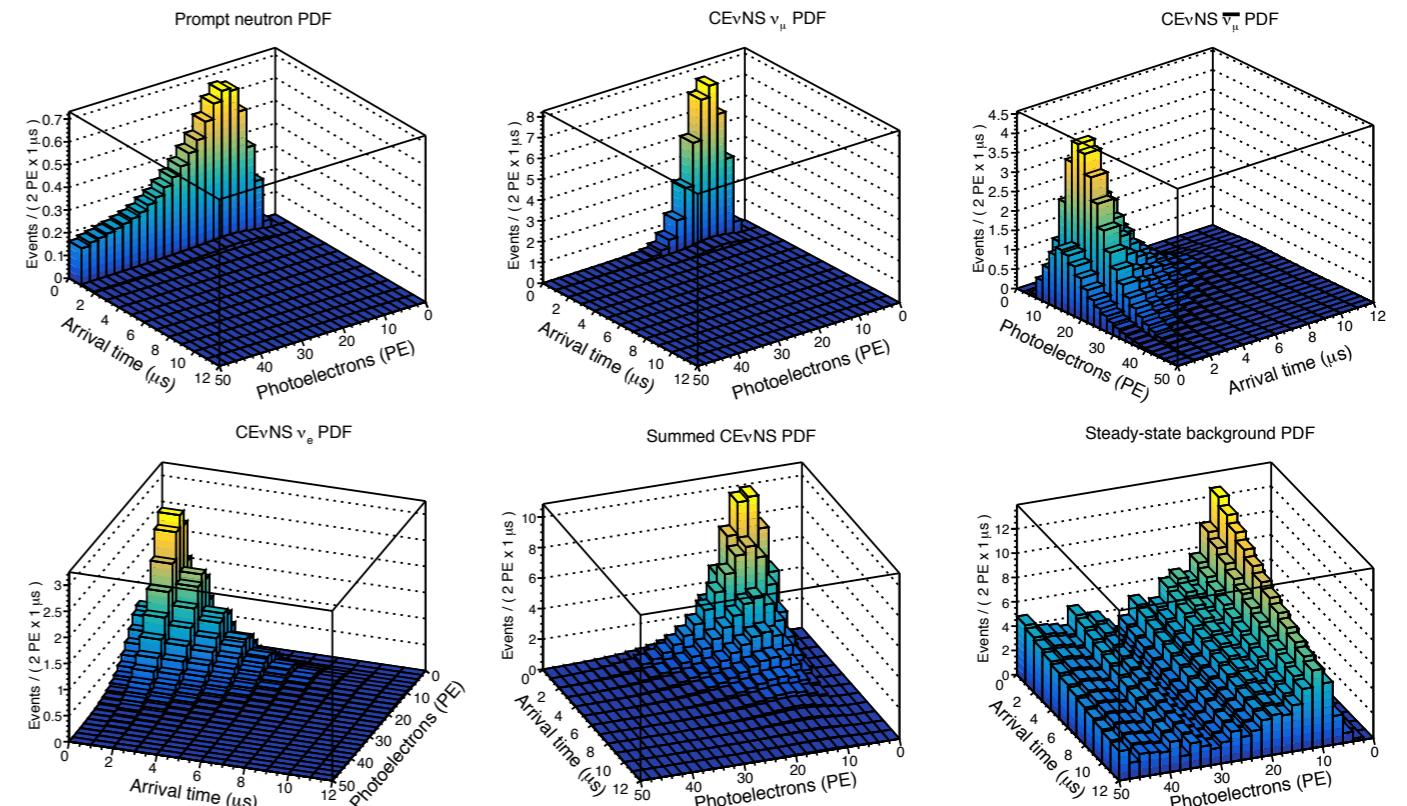
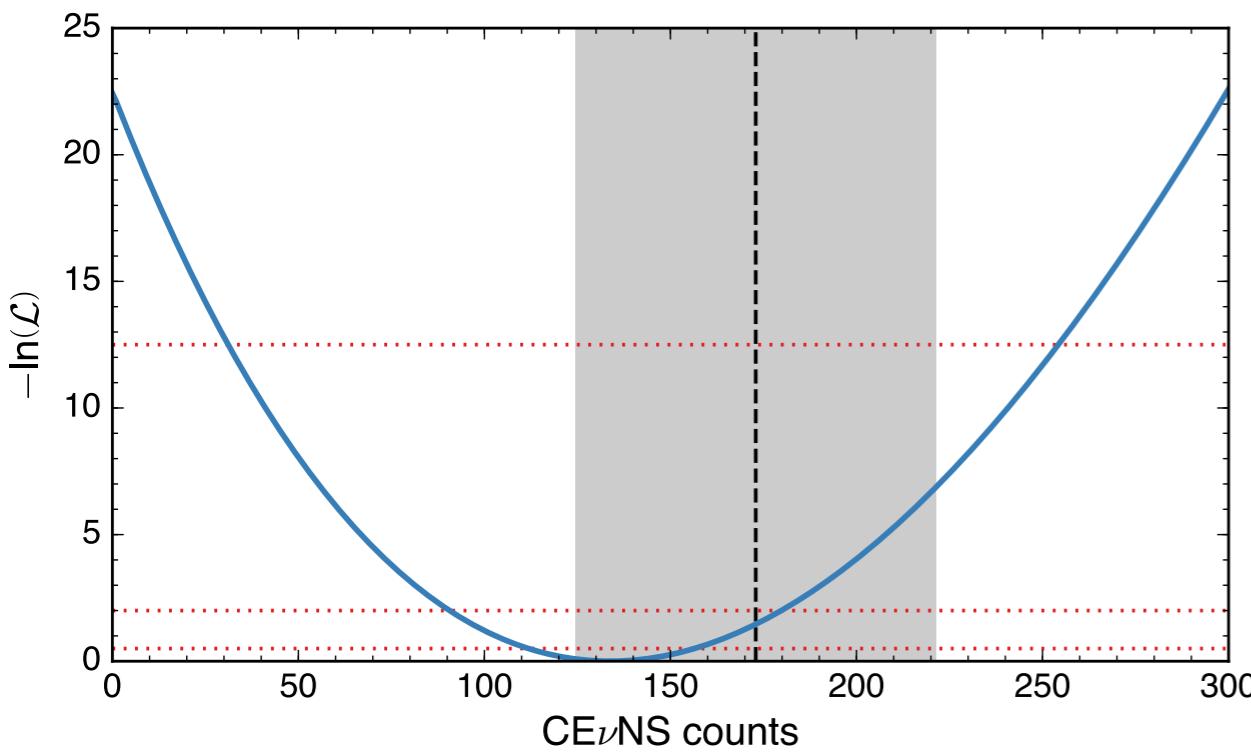
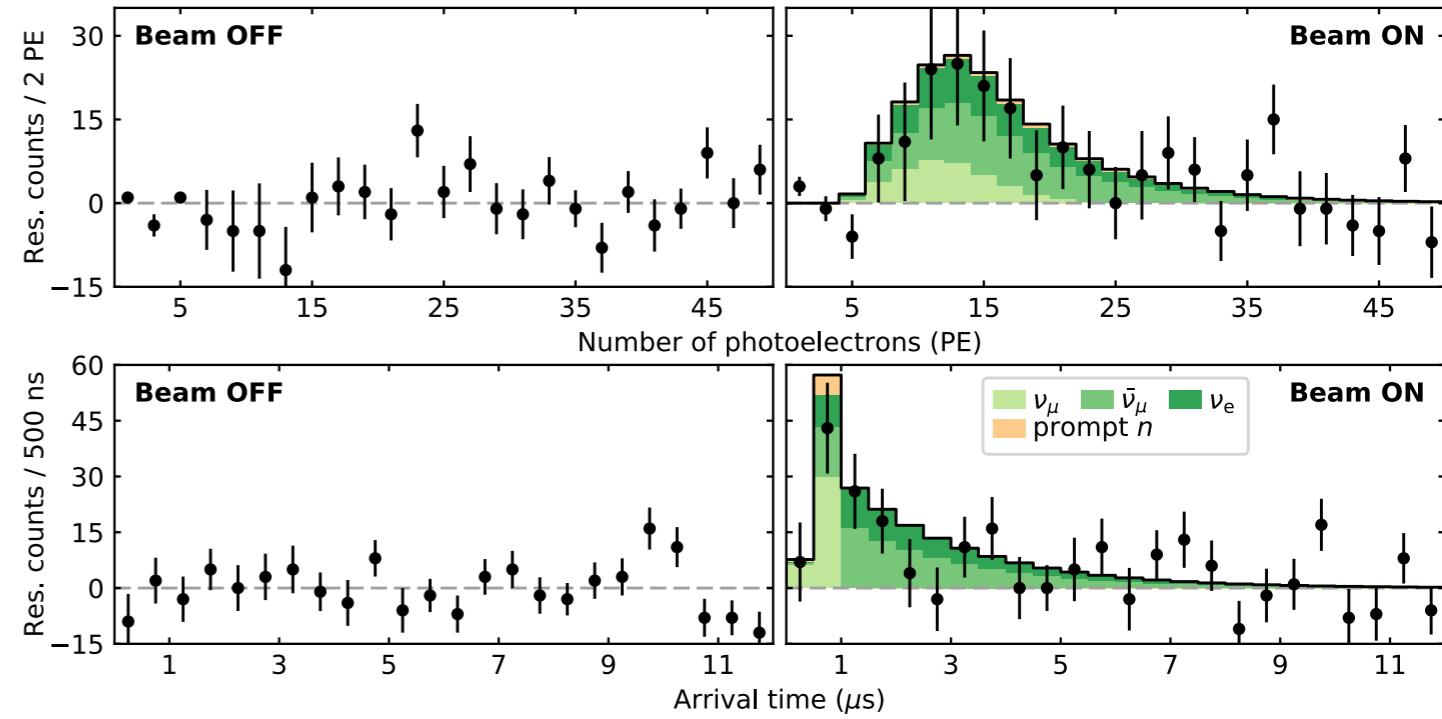
- In-situ measurement also provides a constraint on a neutron-producing background in the lead shield of the detector
- First indications of neutrino-induced neutrons in Pb (a factor of 1.7 below theory prediction)
- Can be important process in many stellar environments
  - C.A. Duba *et al.* J.Phys.Conf.Series 136 (2008)
  - Y-Z. Qian *et al.*, Phys. Rev. C 55 (1997)
  - M. Athar, S. Ahmad and S. K. Singh., Nucl. Phys. A 764 (2006) 551-568
- Both neutron measurements provide constraints that are used in the final analysis



# The Result

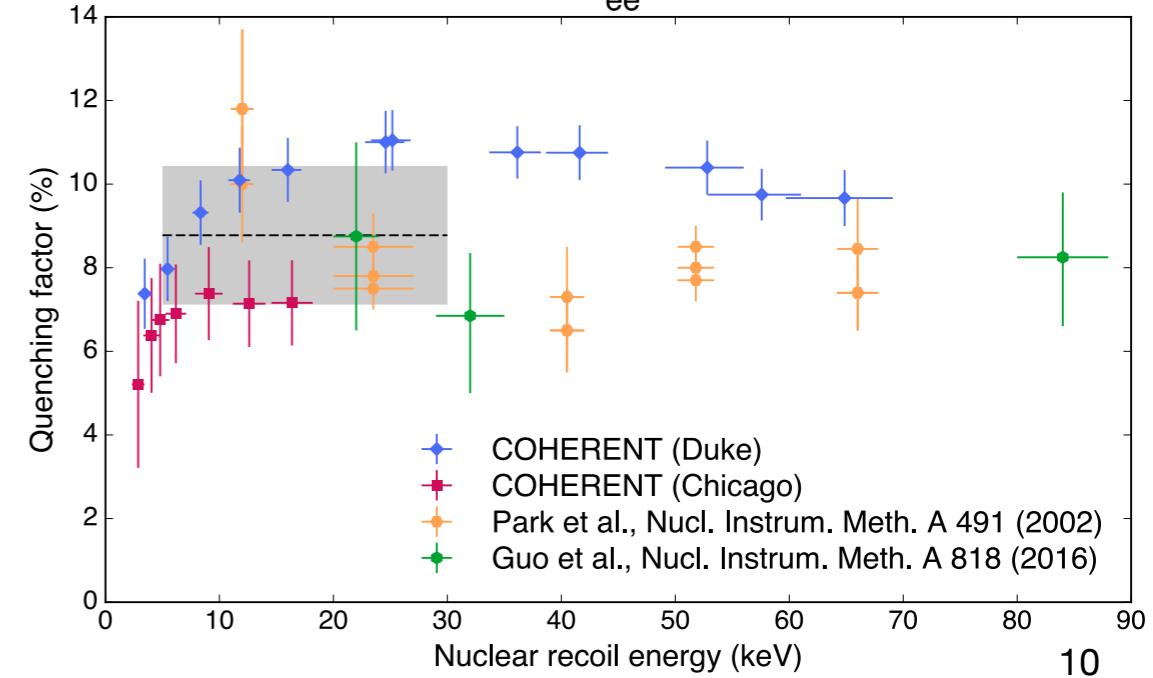
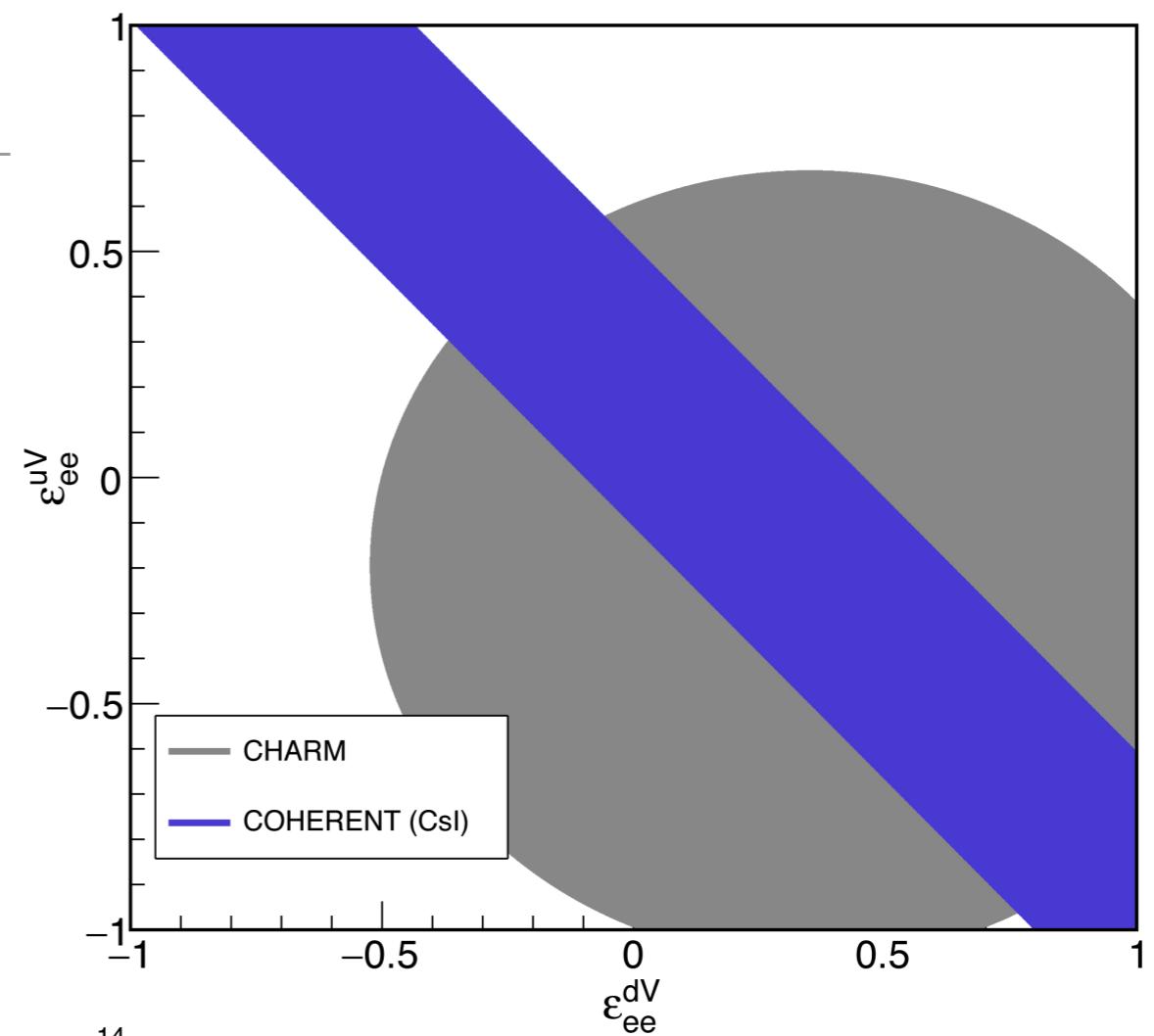
D. Akimov et al., Science 10.1126/science.aao0990 (2017).

- We perform a binned ML fit for the CEvNS signal, including the constraints on the neutron backgrounds, and taking steady-state backgrounds from an anti-coincident window
- We report a **6.7 sigma** significance for an excess of events, that agrees with the standard model prediction to within **1 sigma**



# Implications for Non-Standard Neutrino Interactions

- First result improves constraints on non-universal NSI
- Low hanging fruit. We can expect significant improvement with more data, and when more COHERENT detectors report their results
- Uncertainty currently dominated by our knowledge of the quenching factor in CsI[Na] ( $\sim 25\%$ )
- Can factor out the  $\sim 10\%$  neutrino flux uncertainty by measuring the ratio of interaction rates when the other COHERENT detectors report their results



# Where do we go from here

- More exposure for CsI[Na] detector & better understanding of QF
- A number of COHERENT detectors are now online (LAr, NaI[Tl]-185 kg, MARs neutron detector) with PPC HPGe coming soon and ton-scale NaI being designed!
- Rich neutrino physics program for non-CEvNS cross-sections including CC cross-sections on  $^{127}\text{I}$ ,  $^{56}\text{Fe}$  and  $^{208}\text{Pb}$
- We also look forward to a new era of miniaturized neutrino detector technology with several other collaborations coming on line soon.

