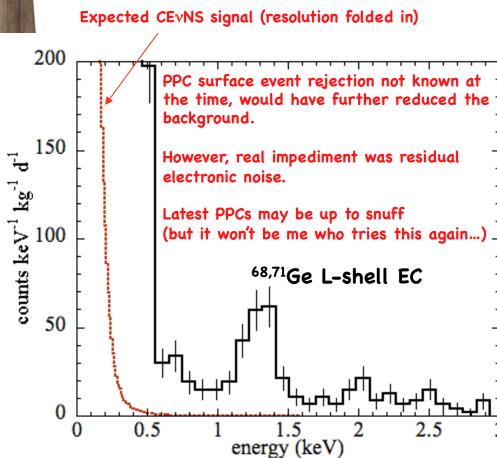


## So close, and yet so far



G. Gratta dixit: "first to put CEVNS signal and backgrounds on a linear-linear plot..."

#### Other reactor fans:



VOLUME 55, NUMBER 1

PHYSICAL REVIEW LETTERS

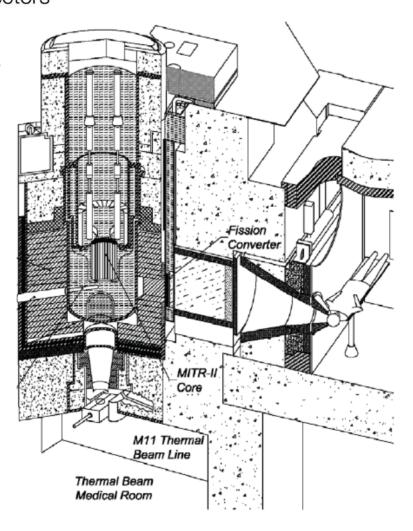
1 JULY 1985

#### **Bolometric Detection of Neutrinos**

Blas Cabrera, Lawrence M. Krauss, and Frank Wilczek
Department of Physics, Stanford University, Stanford, California 94305
Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 01238
Institute for Theoretical Physics, University of California, Santa Barbara, California 93106
(Received 14 December 1984)

Elastic neutrino scattering off electrons in crystalline silicon at 1–10 mK results in measurable temperature changes in macroscopic amounts of material, even for low-energy (<0.41MeV) pp  $\nu$ 's from the sun. We propose new detectors for bolometric measurement of low-energy  $\nu$  interactions, including coherent nuclear elastic scattering. A new and more sensitive search for oscillations of reactor antineutrinos is practical ( $\sim$ 100 kg of Si), and would lay the groundwork for a more ambitious measurement of the spectrum of pp,  $^7\text{Be}$ , and  $^8\text{B}$  solar  $\nu$ 's, and supernovae anywhere in our galaxy ( $\sim$ 10 tons of Si).

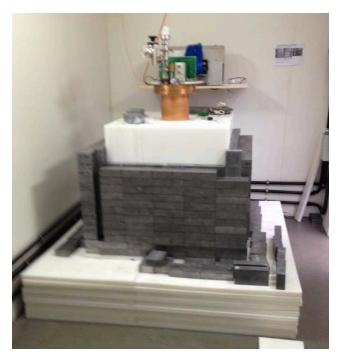


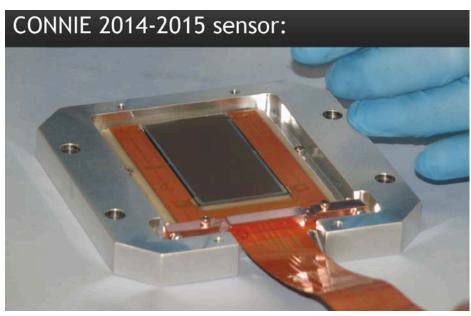


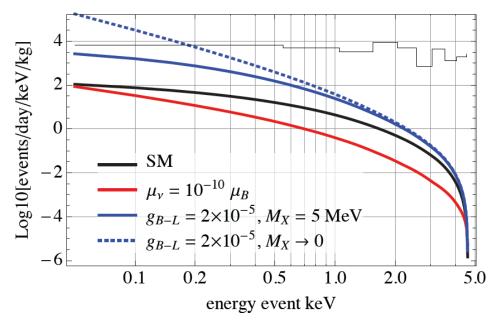
#### Other reactor fans:



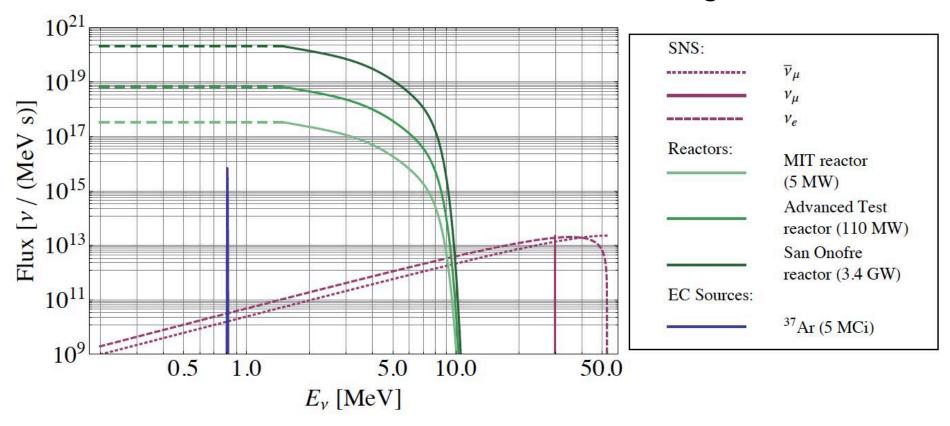








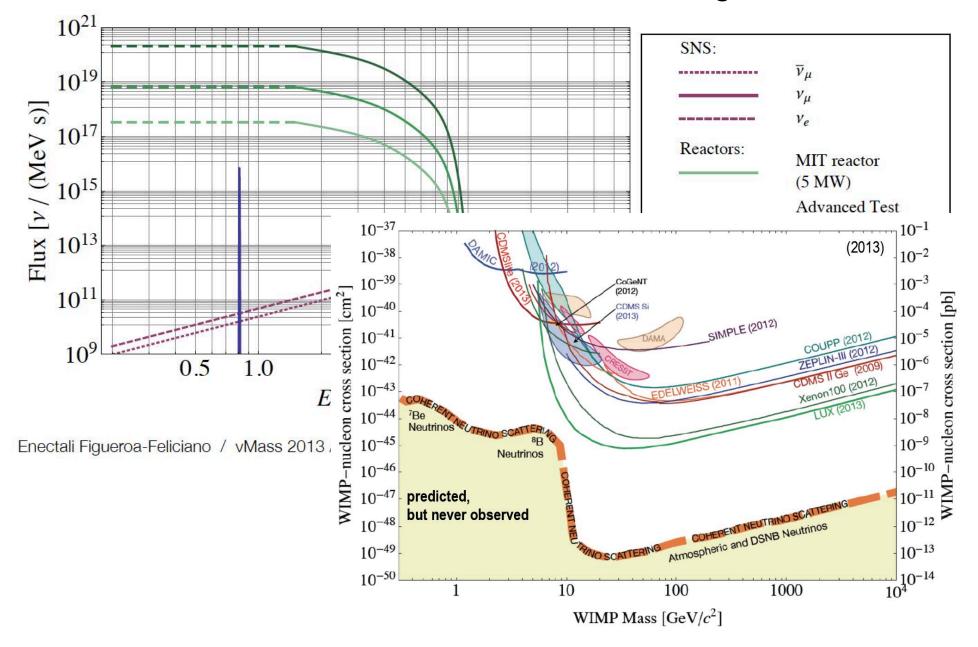
## Fortunately, reactors are not the only game in town:



Enectali Figueroa-Feliciano / vMass 2013 / Milano



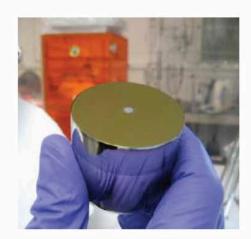
## Fortunately, reactors are not the only game in town:



### Enter COHERENT @ SNS:

## How to Make an Unambiguous Measurement

- Observe the pulsed v time-structure
- Observe the 2.2 µs characteristic decay of muon decay v's
- Observe the N<sup>2</sup> cross section behavior between targets



P-Type Point Contact HPGe



Low-Background Csl[Na]



Nal[TI]

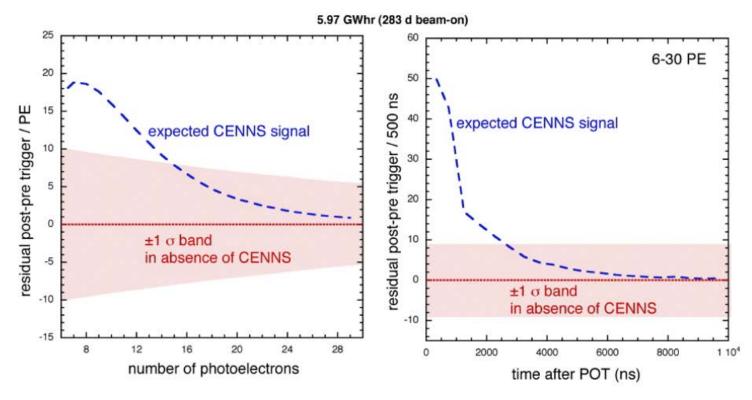


Single Phase LAr

#### Enter COHERENT @ SNS:

## How to Make an Unambiguous Measurement

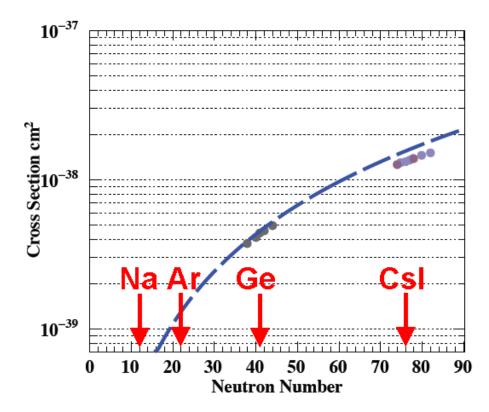
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### Enter COHERENT @ SNS:

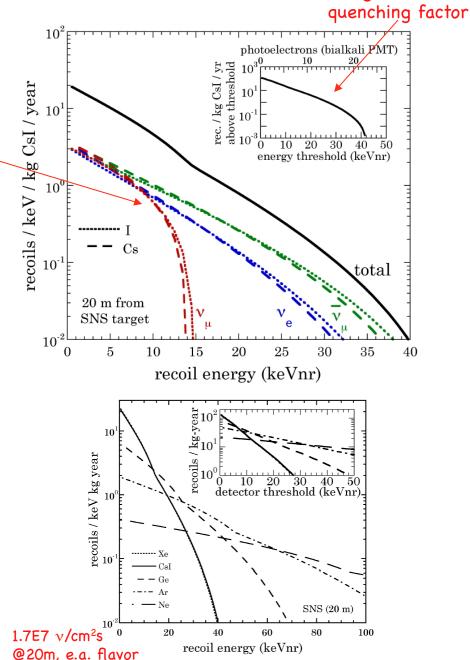
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## Why CsI[Na]? (NIM A773 (2014) 56)

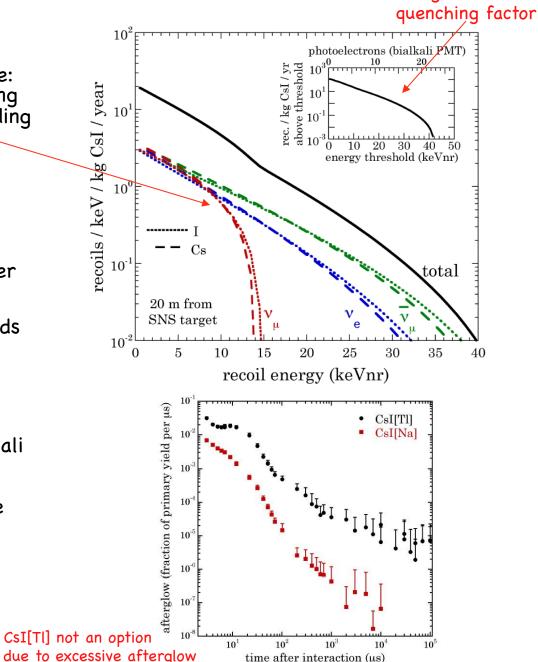
- Large  $N^2$  => large x-section.
- Cs and I surround Xe in Periodic Table: they behave much like a single recoiling species, greatly simplifying understanding the NR response.
- Quenching factor in energy ROI sufficient for ~5 keVnr threshold (we have measured this).
- Statistical NR/ER discrimination is possible at low-E (but will need further improved signal-to-background).
- Sufficiently low in intrinsic backgrounds (U, Th ,K-40, Rb-87, Cs-134,137)
   Measurements in complete SNS shield and 6 m.w.e. indicate we are ready)
- Practical advantages: High light yield (64 ph/keVee), optimal match to bialkali PMTs, rugged, room temperature, inexpensive (\$1/g), modest afterglow (CsI[Tl] not a viable option for surface experiment).
- Expect ~550 v recoils/year in 14 kg detector.



Using **measured** 

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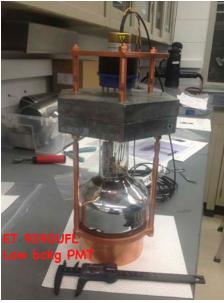


time after interaction (µs)

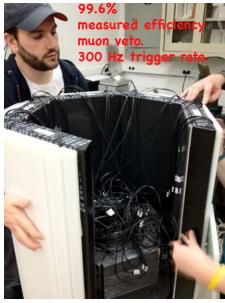
Using **measured** 

## Preliminaries: background studies w/ 2 kg prototype



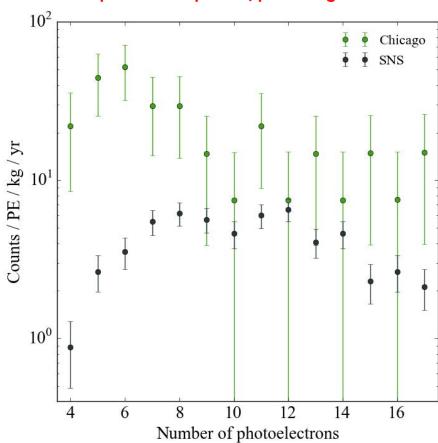




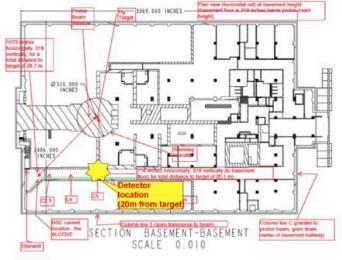


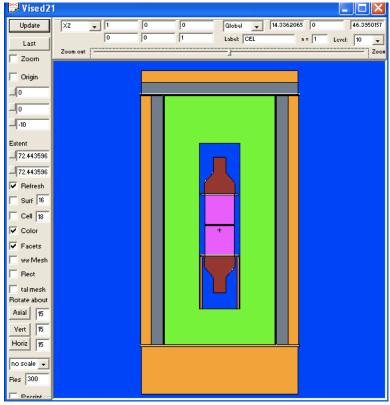
Pulsed SNS signal leads to very low bckg.

We improved on prototype background level!



## Preliminaries: in situ neutron bckg measurements





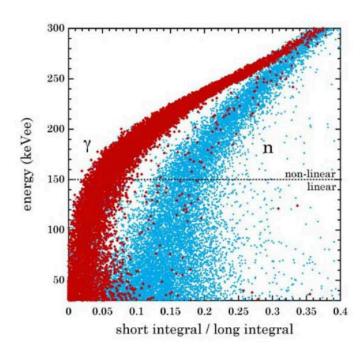


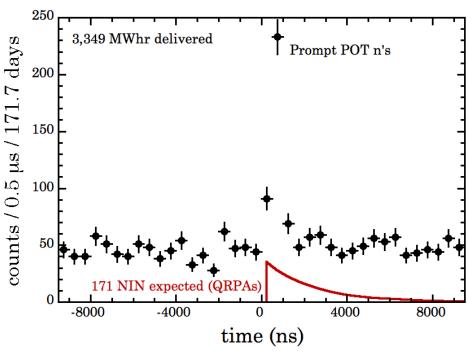




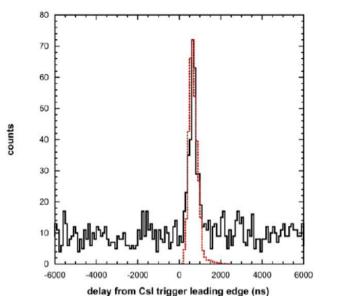


## Preliminaries: in situ neutron bckg measurements

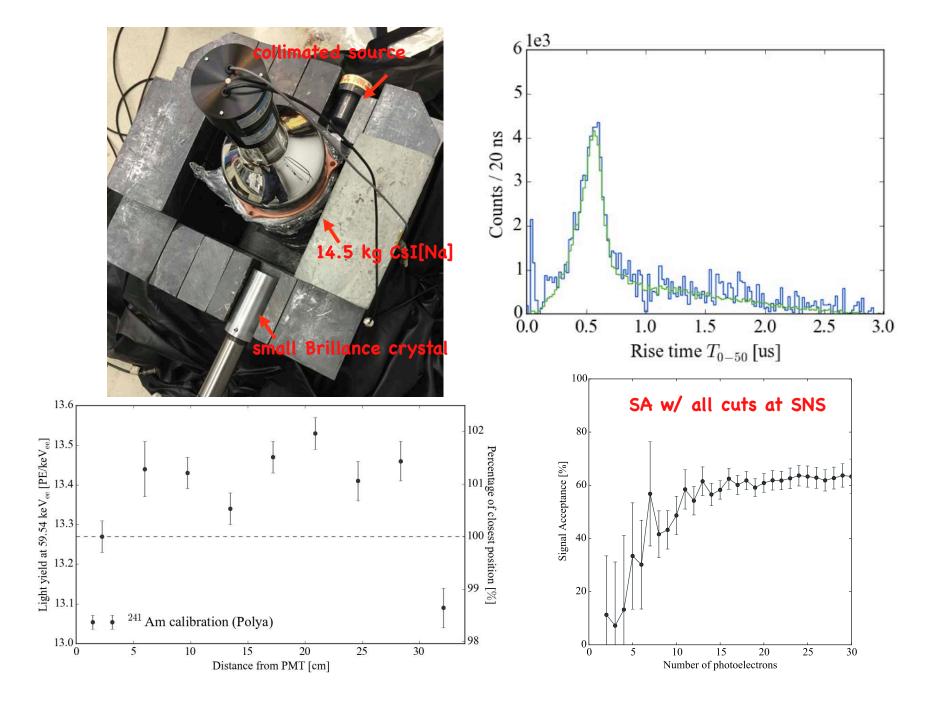




- Measurements informed final shielding design (3" internal HDPE to reduce NIN signal to 1/25 of CENNS).
- Prompt n's at this location also not an issue (contained in time, very low flux at ~2 x 10<sup>-8</sup> n/cm<sup>2</sup>/MeV/s)

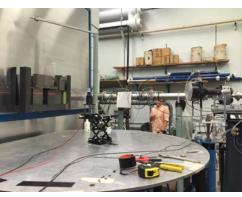


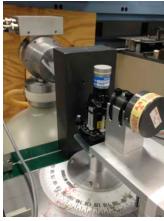
## Preliminaries: 14.5 kg detector characterization



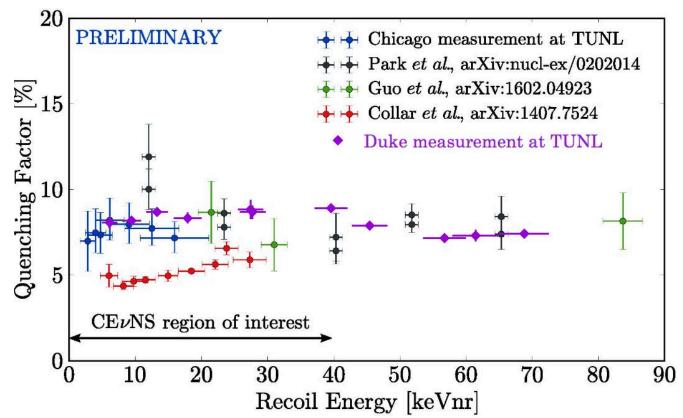
## Preliminaries: Quenching factor measurements





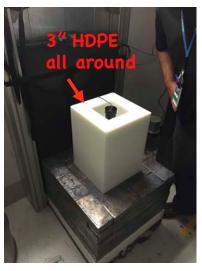


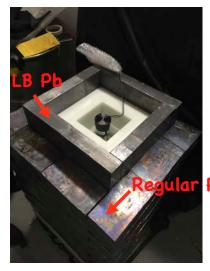




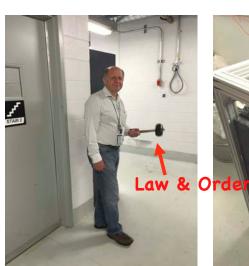
## Installation of 14.5 kg CsI[Na] June 2015









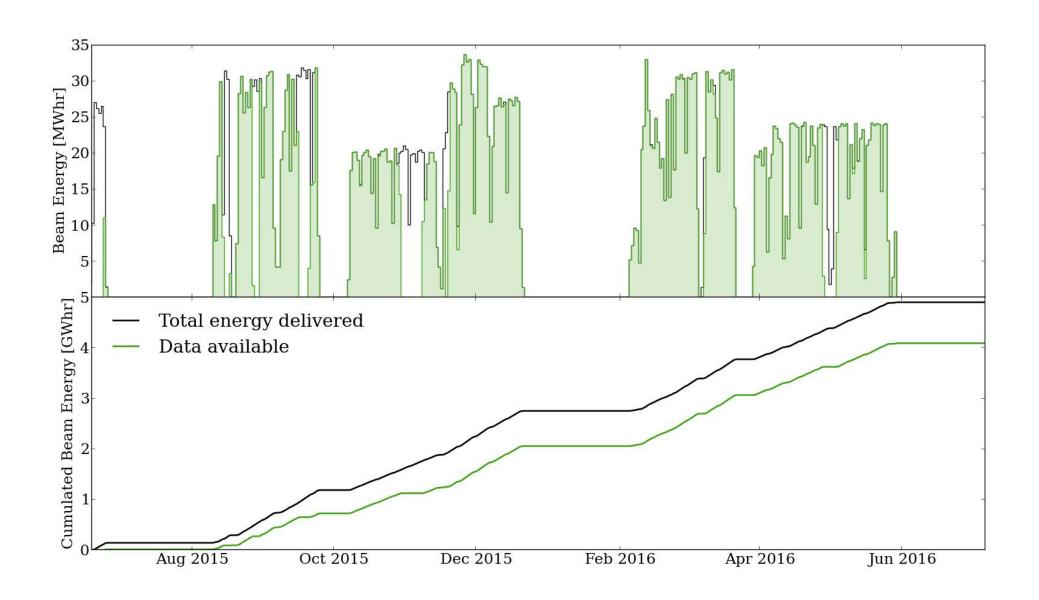






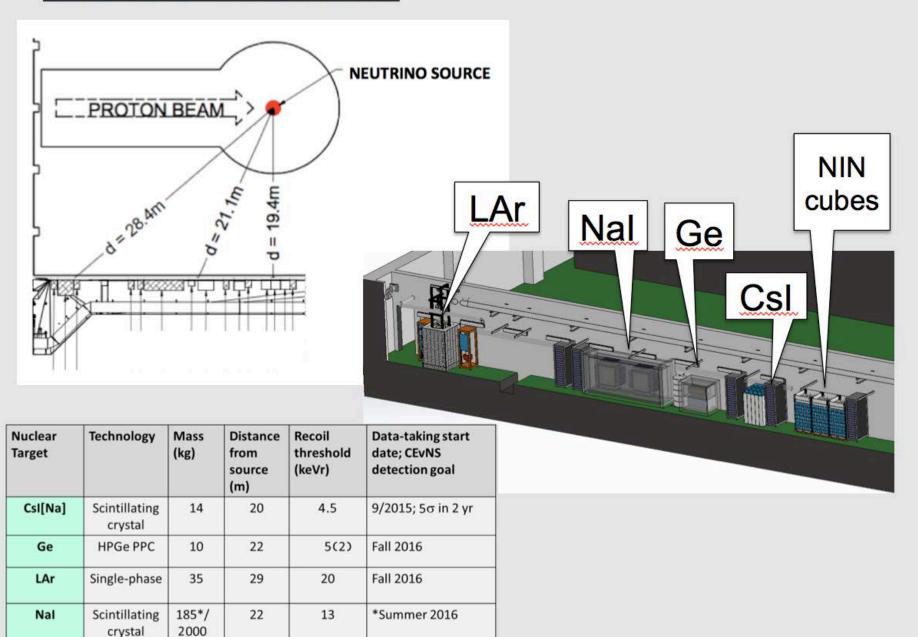


## Installation of 14.5 kg CsI[Na] June 2015





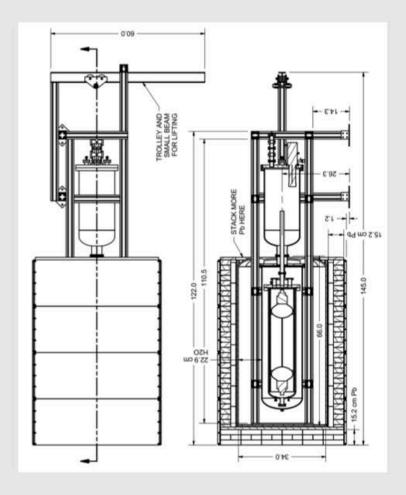
#### **COHERENT DETECTORS AND STATUS**



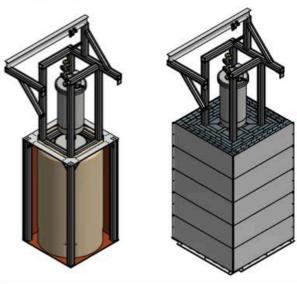


#### CENNS-10 LAr detector for COHERENT

- · CENNS-10 detector built at Fermilab, modified by IU-group
- 35kg LAr fiducial mass. NR/ER discrimination.
- . Pb, Cu, H2O shielding structure built for SNS neutrino corridor
- ~300/700 (prompt/delayed) CEvNS events/vr on 100/900 est. background evts
- · QF measured









#### LAr QF measurements

Measurement of Scintillation and Ionization Yield and Scintillation Pulse Shape from Nuclear Recoils in Liquid

Argon - SCENE Collaboration (Cao, H. et al.) Phys.Rev. D91 (2015) 092007 arXiv:1406.4825 [physics.ins-det] FERMILAB-PUB-14-204-AE-E

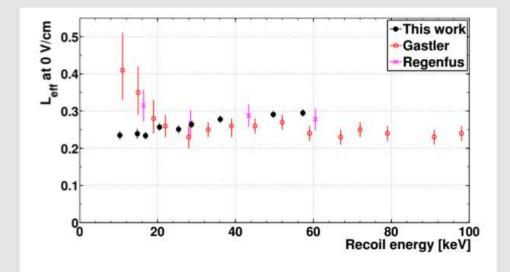
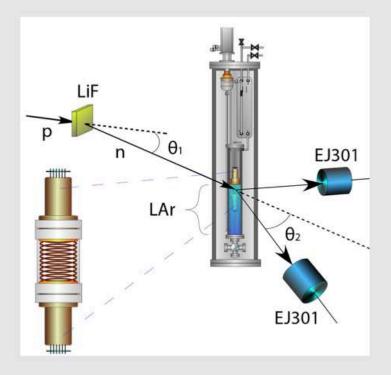


FIG. 10. S1 yield as a function of nuclear recoil energy measured at zero field relative to the light yield of <sup>83m</sup>Kr at zero field, compared to previous measurements[8, 9].

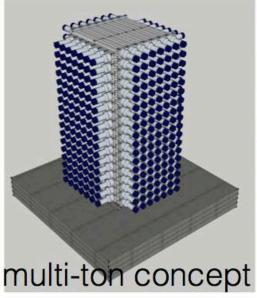


## Detector Subsystems: Nal[TI]

#### https://twitter.com/NalvE\_SNS

- Initial deployment 185 kgs
- Up to 9 T in hand
- N = 23 for Na
- Instrumentation tests underway at Duke and UW
- QF measured by collaboration







## Nal[TI]: Two primary measurement goals

· CEvNS on Na

The electron neutrino Charged & Neutral-Current interaction on <sup>127</sup>I

Isotope	Reaction Channel	Source	Experiment	Measurement (10 <sup>-42</sup> cm <sup>2</sup> )	Theory (10 <sup>-42</sup> cm <sup>2</sup> )
<sup>2</sup> H	$^2{ m H}( u_e,e^-){ m pp}$	Stopped $\pi/\mu$	LAMPF	$52 \pm 18 (tot)$	54 (IA) (Tatara et al., 1990)
<sup>12</sup> C	$^{12}{ m C}( u_e, e^-)^{12}{ m N}_{ m g.s.}$	Stopped $\pi/\mu$ Stopped $\pi/\mu$ Stopped $\pi/\mu$	KARMEN E225 LSND	$\begin{array}{c} 9.1 \pm 0.5 (\mathrm{stat}) \pm 0.8 (\mathrm{sys}) \\ 10.5 \pm 1.0 (\mathrm{stat}) \pm 1.0 (\mathrm{sys}) \\ 8.9 \pm 0.3 (\mathrm{stat}) \pm 0.9 (\mathrm{sys}) \end{array}$	9.4 [Multipole](Donnelly and Peccei, 1979) 9.2 [EPT] (Fukugita <i>et al.</i> , 1988). 8.9 [CRPA] (Kolbe <i>et al.</i> , 1999b)
	$^{12}{ m C}( u_e,e^-)^{12}{ m N}^*$	Stopped $\pi/\mu$ Stopped $\pi/\mu$ Stopped $\pi/\mu$	KARMEN E225 LSND	$5.1 \pm 0.6 (\mathrm{stat}) \pm 0.5 (\mathrm{sys})$ $3.6 \pm 2.0 (\mathrm{tot})$ $4.3 \pm 0.4 (\mathrm{stat}) \pm 0.6 (\mathrm{sys})$	5.4-5.6 [CRPA] (Kolbe <i>et al.</i> , 1999b) 4.1 [Shell] (Hayes and S, 2000)
	$^{^{12}\text{C}(\nu_{\mu},\nu_{\mu})^{12}\text{C}^*}_{^{12}\text{C}(\nu,\nu)^{12}\text{C}^*}$	Stopped $\pi/\mu$ Stopped $\pi/\mu$	KARMEN KARMEN	$3.2 \pm 0.5 (\mathrm{stat}) \pm 0.4 (\mathrm{sys})$ $10.5 \pm 1.0 (\mathrm{stat}) \pm 0.9 (\mathrm{sys})$	2.8 [CRPA] (Kolbe <i>et al.</i> , 1999b) 10.5 [CRPA] (Kolbe <i>et al.</i> , 1999b)
	$^{12}\mathrm{C}( u_{\mu},\mu^{-})\mathrm{X}$	Decay in Flight	LSND	$1060 \pm 30 ({\rm stat}) \pm 180 ({\rm sys})$	1750-1780 [CRPA] (Kolbe <i>et al.</i> , 1999b) 1380 [Shell] (Hayes and S, 2000) 1115 [Green's Function] (Meucci <i>et al.</i> , 2004)
	$^{12}{ m C}( u_{\mu},\mu^{-})^{12}{ m N}_{ m g.s.}$	Decay in Flight	LSND	$56 \pm 8(\mathrm{stat}) \pm 10(\mathrm{sys})$	68-73 [CRPA] (Kolbe <i>et al.</i> , 1999b) 56 [Shell] (Hayes and S, 2000)
<sup>56</sup> Fe	$^{56}{\rm Fe}(\nu_e,e^-)^{56}{\rm Co}$	Stopped $\pi/\mu$	KARMEN	$256 \pm 108 ({ m stat}) \pm 43 ({ m sys})$	264 [Shell] (Kolbe et al., 1999a)
<sup>71</sup> Ga	$^{71}\mathrm{Ga}( u_e,e^-)^{71}\mathrm{Ge}$	<sup>51</sup> Cr source <sup>51</sup> Cr <sup>37</sup> Ar source	GALLEX, ave. SAGE SAGE	$0.0054 \pm 0.0009 (tot)$ $0.0055 \pm 0.0007 (tot)$ $0.0055 \pm 0.0006 (tot)$	0.0058 [Shell] (Haxton, 1998) 0.0070 [Shell] (Bahcall, 1997)
$^{127}I$	$^{127}{ m I}( u_e,e^-)^{127}{ m Xe}$	Stopped $\pi/\mu$	LSND	$284 \pm 91 \mathrm{(stat)} \pm 25 \mathrm{(sys)}$	210-310 [Quasi-particle] (Engel et al., 1994)

# Detector Subsystems: HPGe PPCs

- Smaller N: 38-44
- Excellent resolution at low energies
- Well-measured quenching factor
- Phase I: 5-10kg PPC Ge detector array:
  - Repurposing on-hand MAJORANA DEMONSTRATOR/LANL natGe detectors.
  - Copper/Lead/Poly shield with Plastic scintillator μ-veto.
  - Installation in Fall 2016
- Potential Phase II: Expansion of target with larger-mass (C4-style) point contact detectors.

